



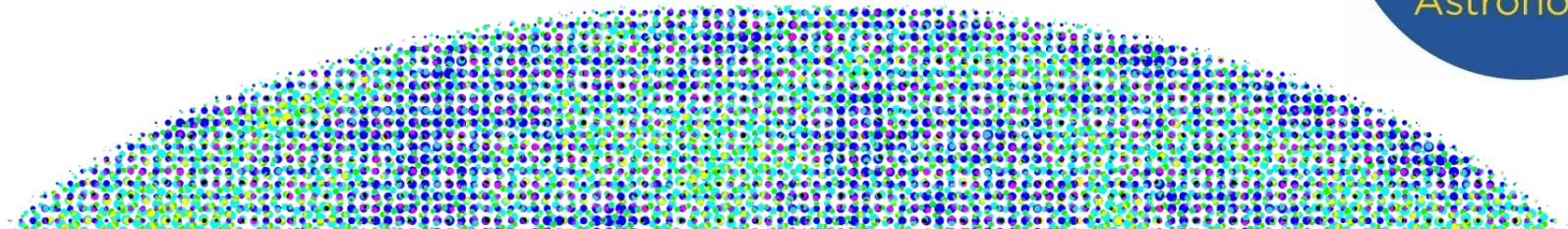
# 2024 Advances in Neutrino Physics Driven by Cosmological Data

CMB-S4 Summer Collaboration meeting

Helena Garcia Escudero (she/her)

[garciaeh@uci.edu](mailto:garciaeh@uci.edu)

08/01/2024



# Cosmological observations and Neutrino physics

- CMB-S4 impact on our understanding of the Universe:
    - Inflation.
    - Neutrino physics
- 1.- Sum of the neutrino masses
  - 2.- Effective number of light relativistic particles
  - 3.- Beyond Standard Model Scenarios



# What is the value of the neutrino mass?



IMAGE CREDIT: <https://phys.org/news/2018-06-katrin-neutrino-mass.html>

- KATRIN upper limit on effective electron anti-neutrino mass 0.8 eV  
(doi: 10.1038/s41567-021-01463-1)
- Cosmology can constrain neutrino mass > **20 times better** today already!\* [arxiv 2404.03002](https://arxiv.org/abs/2404.03002)

$$\Lambda\text{CDM} + \sum m_\nu$$

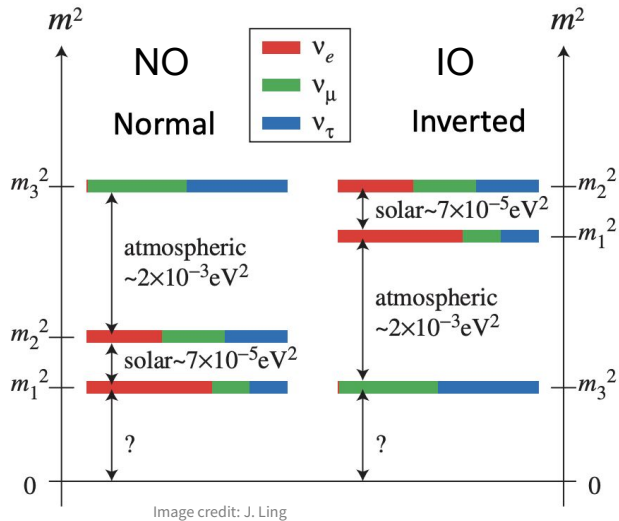
$$\text{DESI} + \text{CMB}$$

$$\sum m_\nu < 0.072 \text{ eV} \quad [95\% \text{ CL}]$$

\*model dependent

# Sum of the neutrino masses and Cosmology

- DESI+CMB-S4 sensitivity 2-3  $\sigma$  to the minimal neutrino mass (arxiv: 1610.02743)  
(58 meV from oscillation experiments, i.e.: Super-Kamiokande, SNO, KamLAND).



- We found employing most robust data sets, statistical validations, theory accuracy :

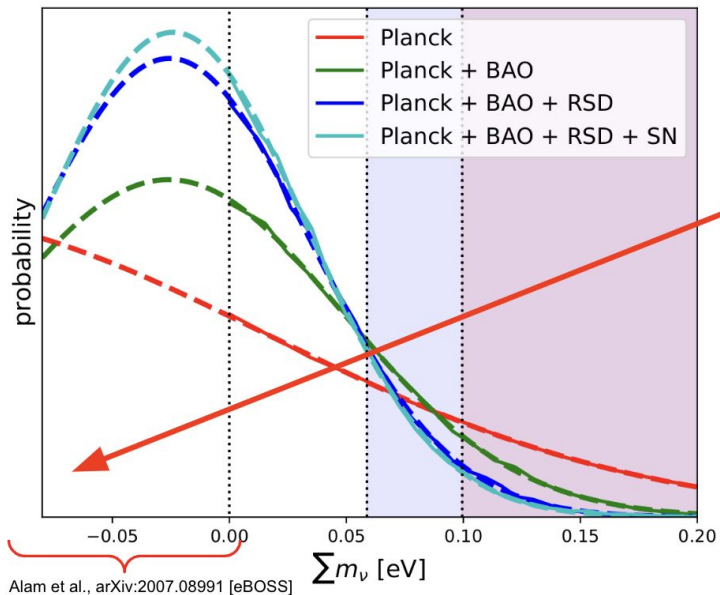
P18+DESI(BAO)	$\sigma$
Preference $\sum m_\nu = 0$ over NO	1.54
Preference NO over IO	1.86

H.Garcia Escudero and K. Abazajian 2024 (in prep)

- We are already approaching it today!

CMB-S4 expected to be sensitive to  $\sum m_\nu$  in the minimal mass scenario.

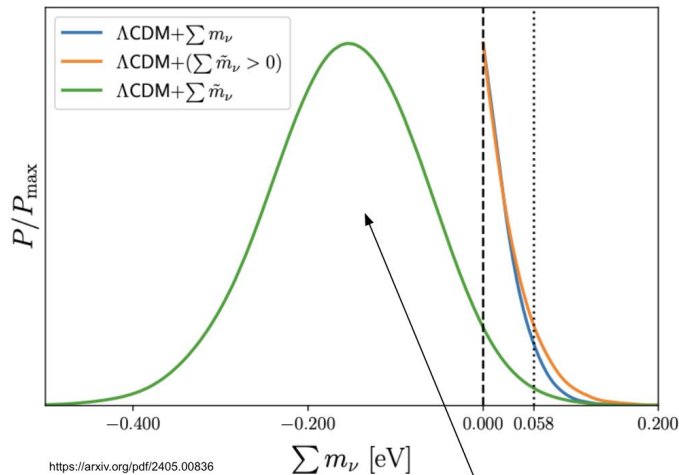
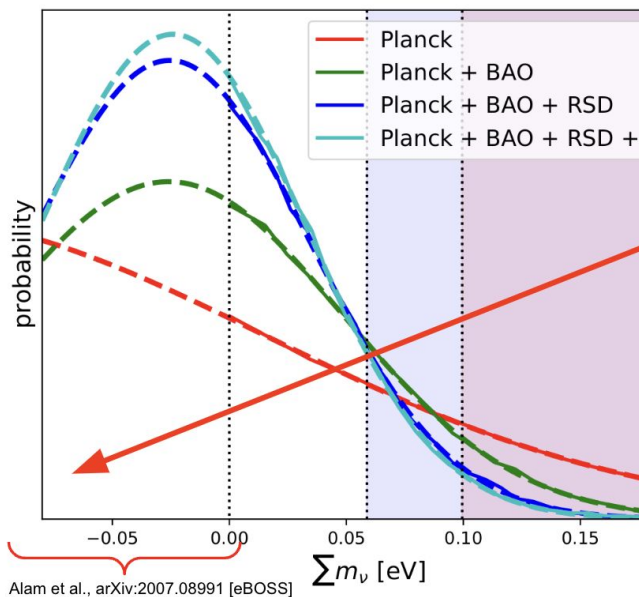
# Negative neutrino mass anomaly



Where is the stringency from Planck 2018 coming from?



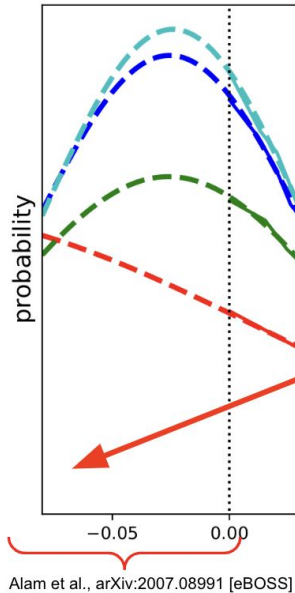
# Negative neutrino mass anomaly



Posterior of  $\sum m_\nu$  (eV) inferred from Planck + ACT Lensing + DESI

$$\sum m_\nu = -160 \pm 90 \text{ meV} (68\%)$$

# Negative neutrino mass anomaly



- Possible origin of “negative” neutrino mass bounds:
  - CMB lensing arXiv: 2407.07878
  - CMB polarization determination of  $\tau$
  - $z = 0.7$  BAO systematics. arXiv: 2404.03002
- Unphysical negative masses weak tension at  $< \sim 2\sigma$  but worth exploring.
- Future CMB-S4 data study this discrepancy.

or of  $\sum m_\nu$ ,  
ferred from  
+ ACT  
g + DESI  
68%)

# Effective Number of Neutrino Species, $N_{\text{eff}}$

- Total cosmological number density of neutrinos

$$n_\nu = N_{\text{eff}} \left(\frac{3}{4}\right) \left(\frac{4}{11}\right) n_\gamma,$$

- $N_{\text{eff}} = 3 \Rightarrow$  neutrinos instantaneously decoupled from the primordial plasma.

- Neutrino properties early universe  $\rightarrow N_{\text{eff}} = 3.044$

arxiv 2005.07047

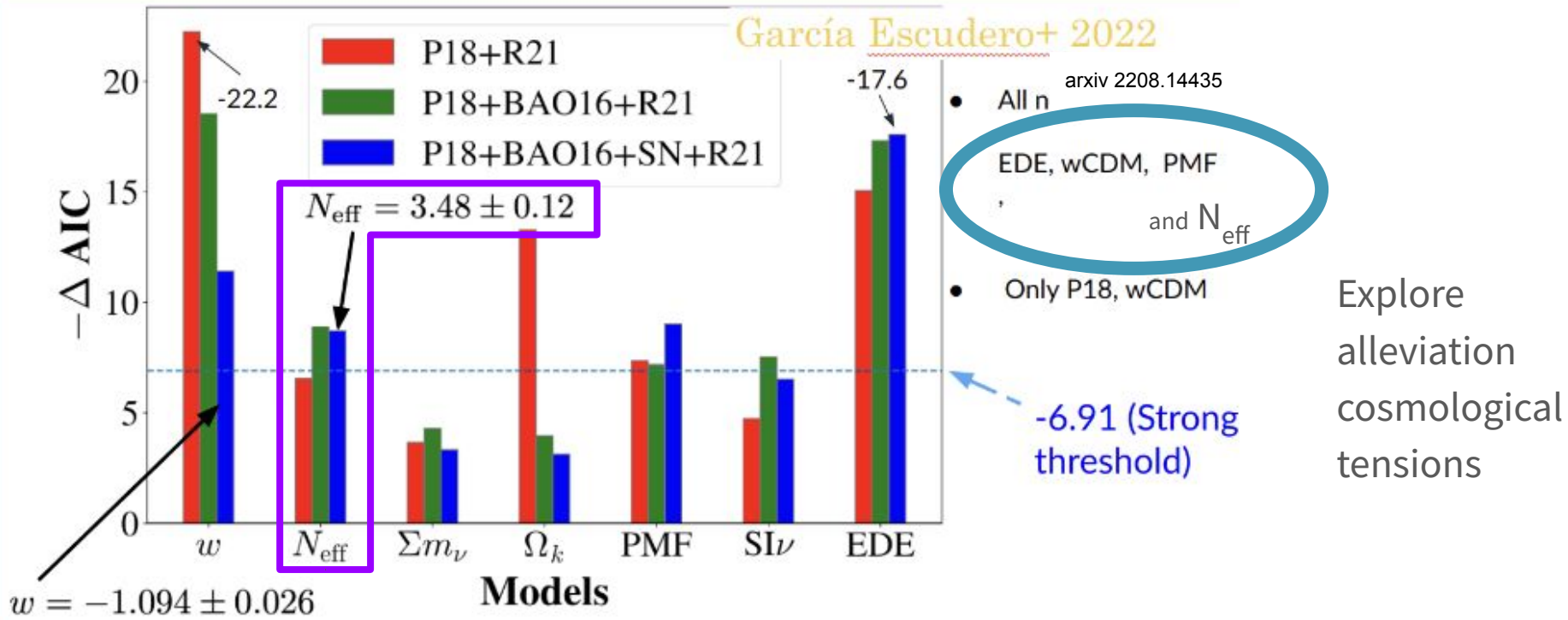
CMB-S4 observation with 2-3  $\sigma$  sensitivity on  $N_{\text{eff}}$  can find:

$N_{\text{eff}}$  consistent with 3.044  $\rightarrow$

- Confirmation of standard cosmology arxiv 1807.06209
- Precise understanding thermal conditions universe.



$N_{\text{eff}}$  significantly different from 3.044 = signature of new physics!



$N_{\text{eff}}$  is a preferred model comparable with other exiting candidates

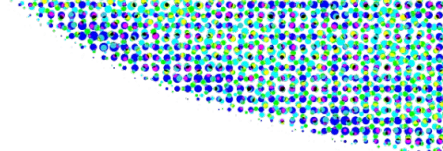
# Short baseline anomalies and Sterile neutrinos

- $N_{\text{eff}}$  higher than 3.044 = new physics =?? hidden neutrino sector!
- Short baseline neutrino oscillations results ( Super-Kamiokande, LSND, MiniBooNE) and Gallium anomaly (i.e.: SAGE, GALLEX, BEST), hint richer neutrino sector.
- More than 3 active neutrinos with one (or more ) sterile flavors.

## Preliminary results:

- We find  $>2\sigma$  preference eV sterile model compared to  $\Lambda$ CDM using Planck CMB +BAO (DESI).

H.Garcia Escudero and K. Abazajian 2024 (in prep)



# Conclusions and upcoming future

- Cosmological observations are a **very powerful tool** for high precision determination neutrino cosmological parameters.
- Combined cosmological probes (CMB-S4 + BAO) **revolutionize understanding of neutrino physics**:
  - 1.- Test standard cosmological model predictions.
  - 2.- Shred light to existing cosmological anomalies.
  - 3.- Test new physics.
- Upcoming observational results next decade critical understanding Universe!



**Thank you for attention!  
Questions?**

**Helena Garcia Escudero**

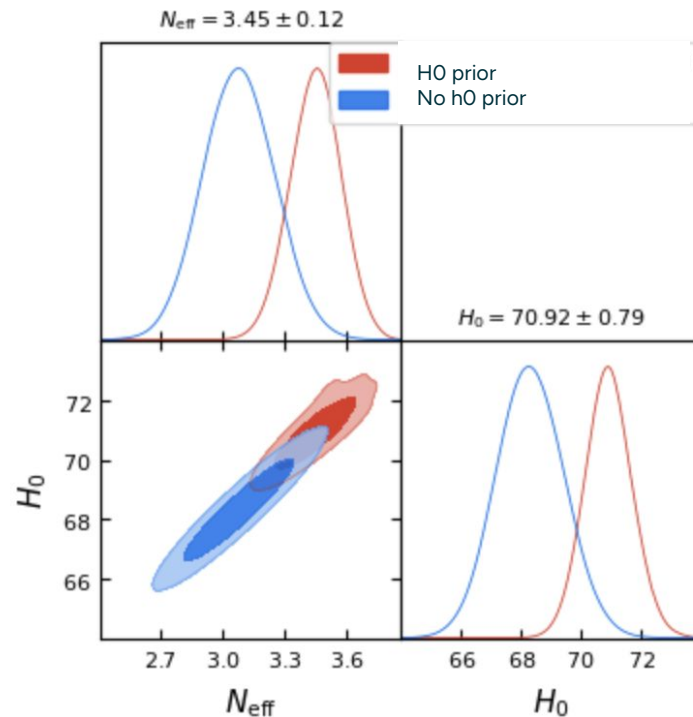
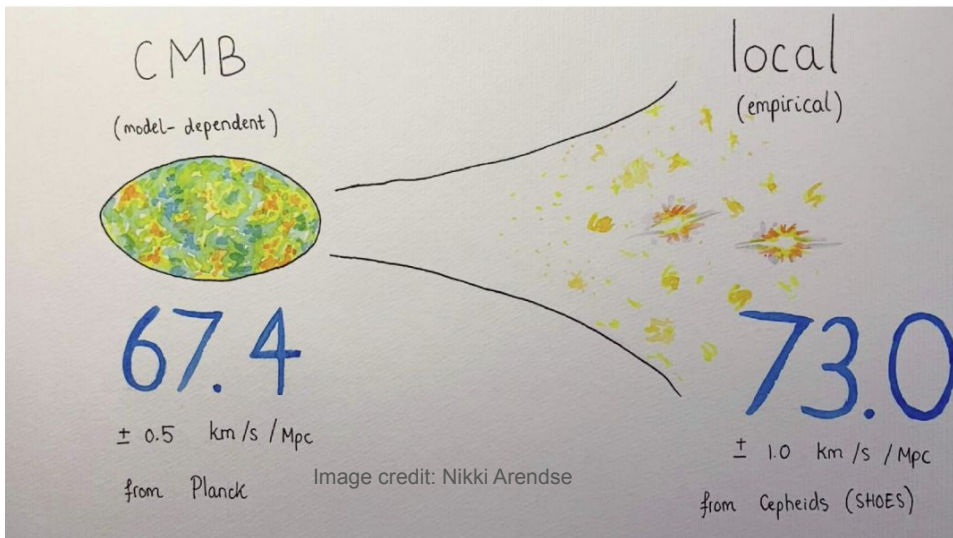
**([garciaeh@uci.edu](mailto:garciaeh@uci.edu))**

# Backup slides

**Helena Garcia Escudero**

**([garciaeh@uci.edu](mailto:garciaeh@uci.edu))**

$N_{\text{eff}}$  significantly different from 3.044 = signature of new physics!

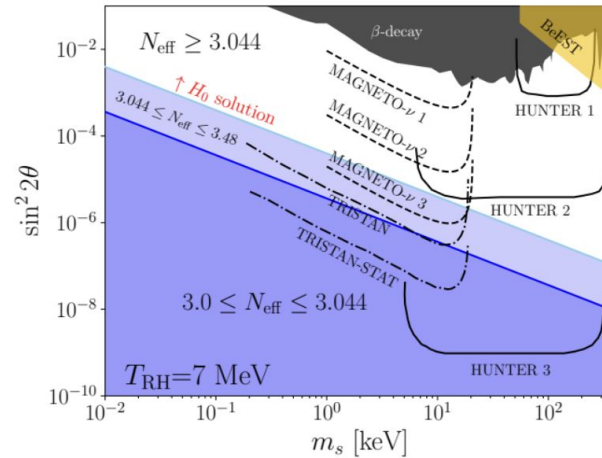
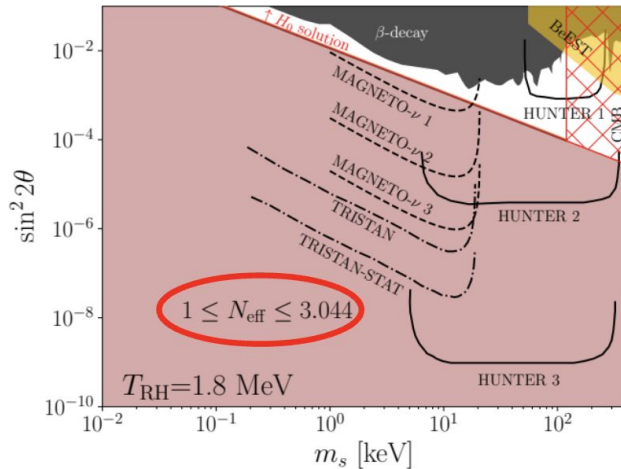


- Explore alleviation cosmological tensions

arxiv: 2107.10291

# Negative neutrino mass anomaly

LRT Universes might also be driving a  $<$  neutrino masses



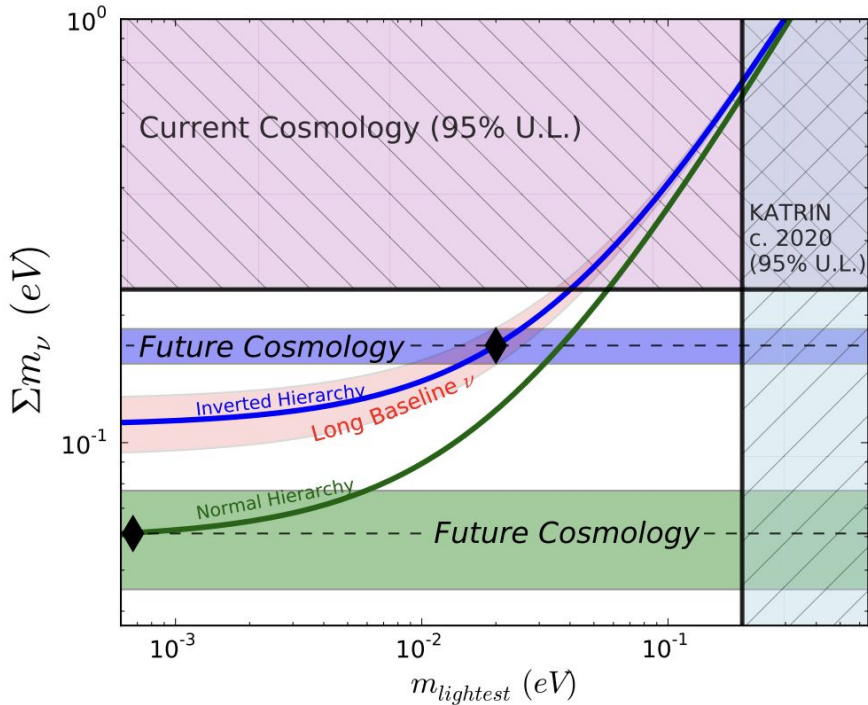
-LRT universes + new neutrino physics  $\Rightarrow$  broad new **observationally-consistent** & novel cosmologies!

-DR can be a combination of thermal + non thermal radiation  $\Rightarrow$  **novel CMB signature!**

- $\uparrow N_{eff}$   $\Rightarrow$  alleviation of the  $H_0$  tension!

<https://arxiv.org/pdf/2309.11492.pdf>

# Conclusions and upcoming future



Combined probes CMB-S4 experiment +  
BAO precision neutrino cosmological  
parameters (within LCDM):

$$\sigma \left( \sum m_\nu \right) = 16 \text{ meV},$$
$$\sigma (N_{\text{eff}}) = 0.020,$$