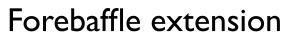
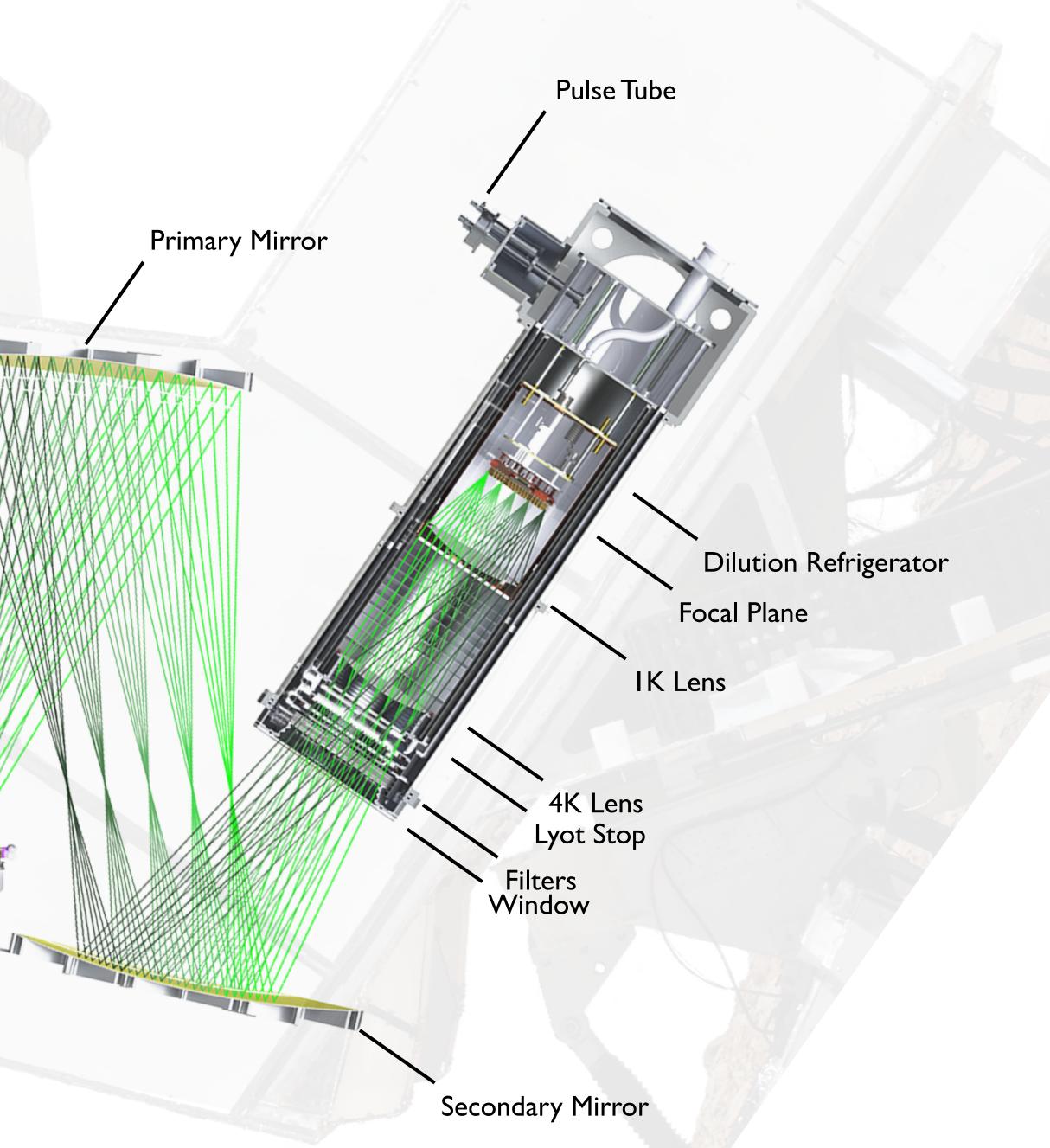
Efficacy of the Variable-delay Polarization Modulator Lessons Learned from CLASS Yunyang Li on behalf of the CLASS Collaboration



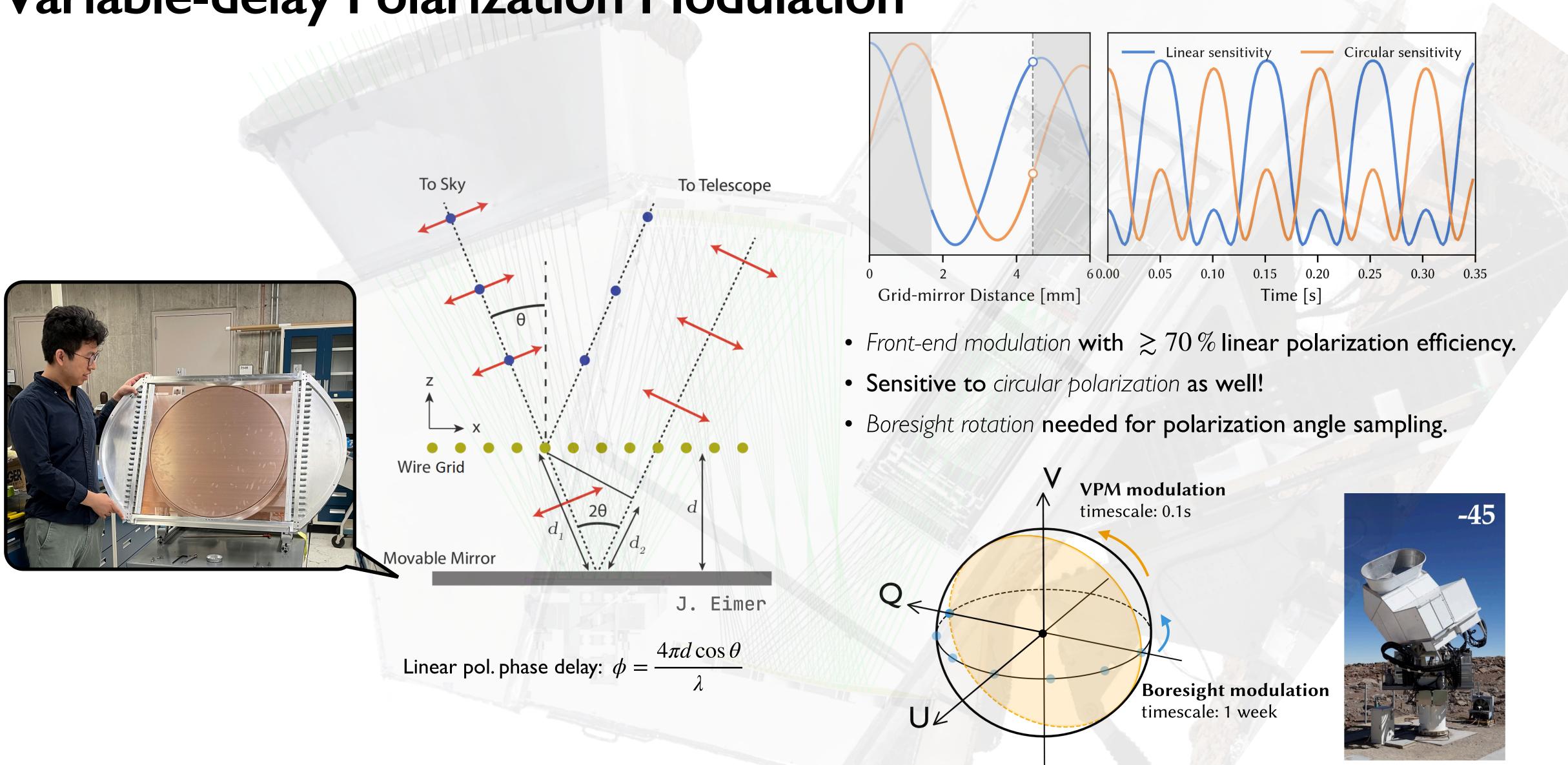




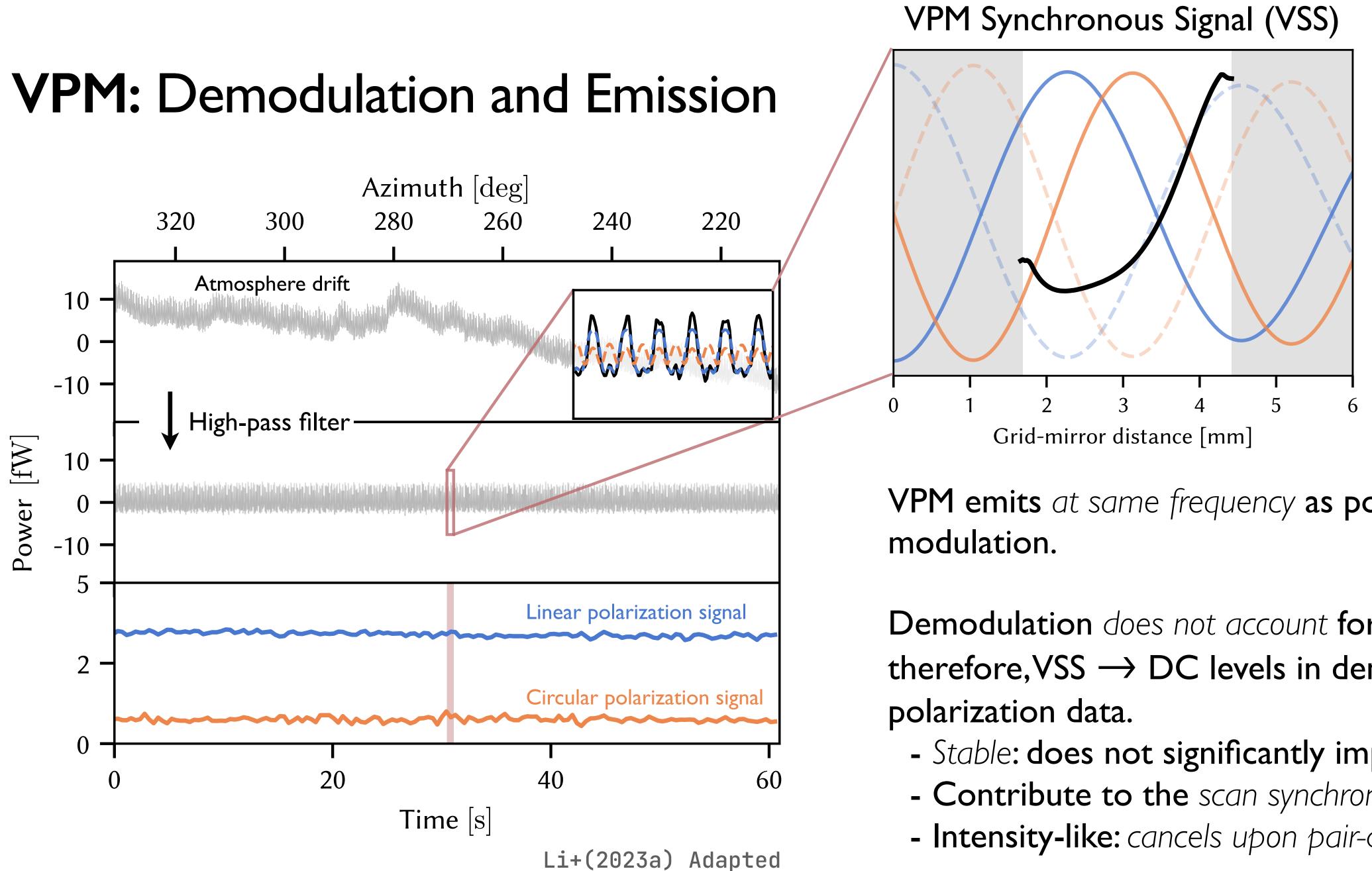




Variable-delay Polarization Modulation



hear pol. phase delay:
$$\phi = \frac{4\pi}{--}$$



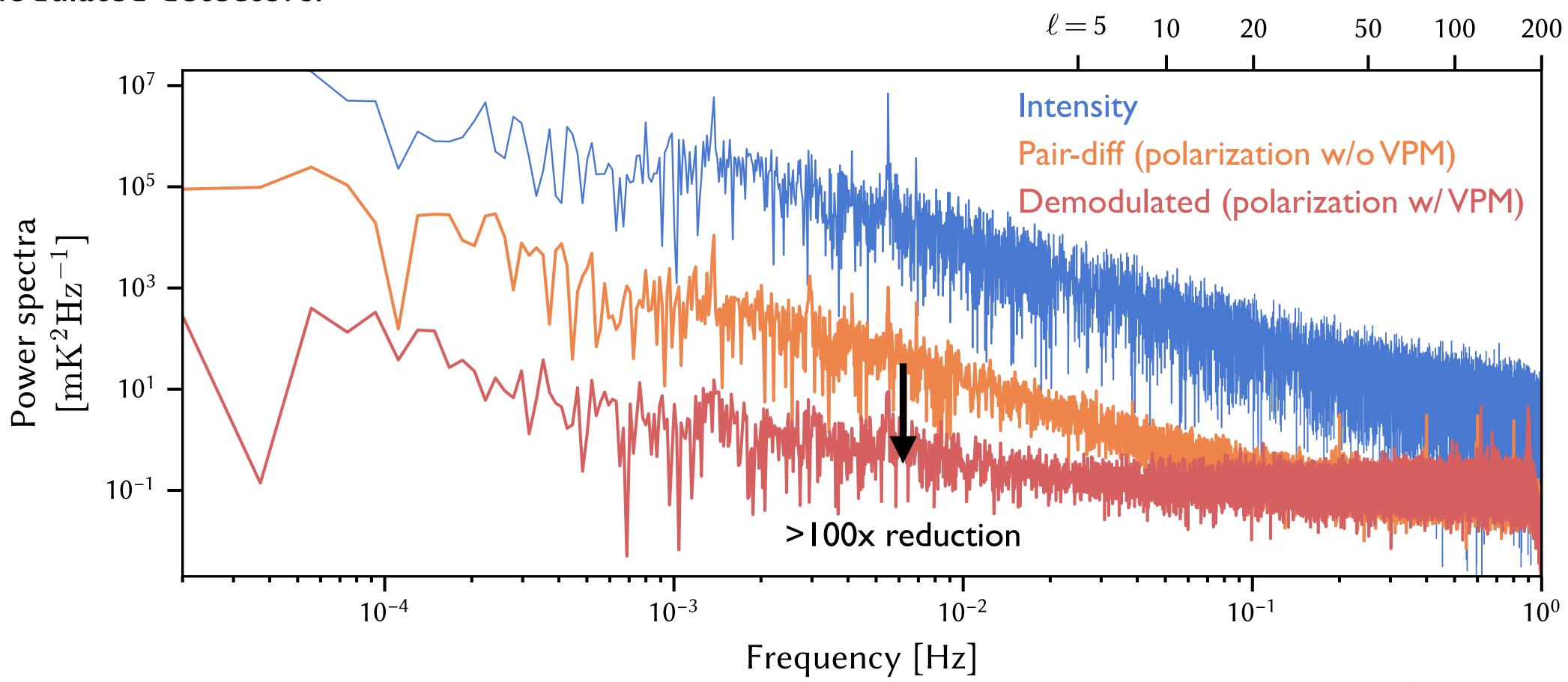
VPM emits at same frequency **as polarization**

Demodulation does not account for the VSS, therefore, VSS \rightarrow DC levels in demodulated

- Stable: does not significantly impact 1/f.
- Contribute to the scan synchronous noise.
- Intensity-like: cancels upon pair-diff.

VPM: Results

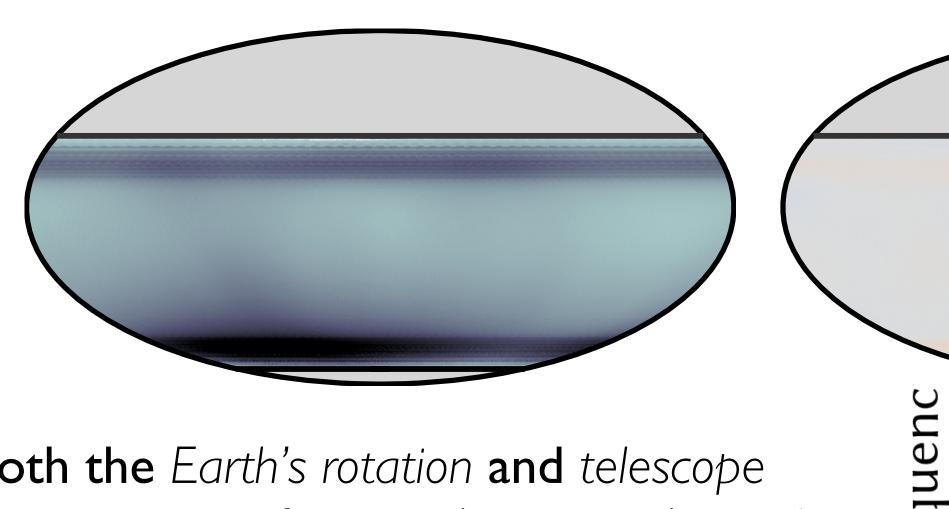
Modulation reduces noise power at <10 mHz by over two orders of magnitude compared to pair-differencing unmodulated detectors.



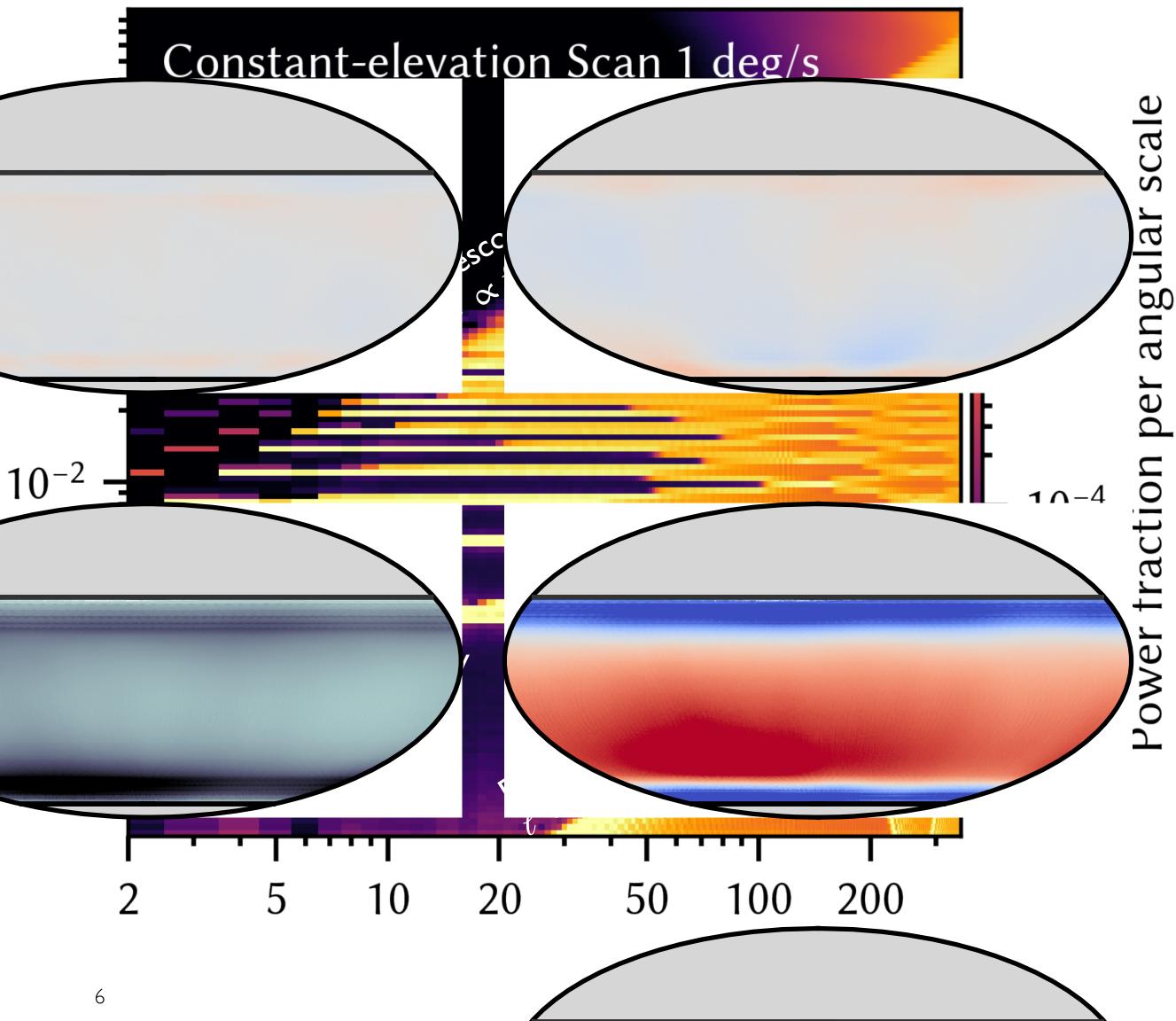
Harrington+(2021) Adapted

Mappings between Fourier & Harmonic modes

Continuous constant-elevation scans

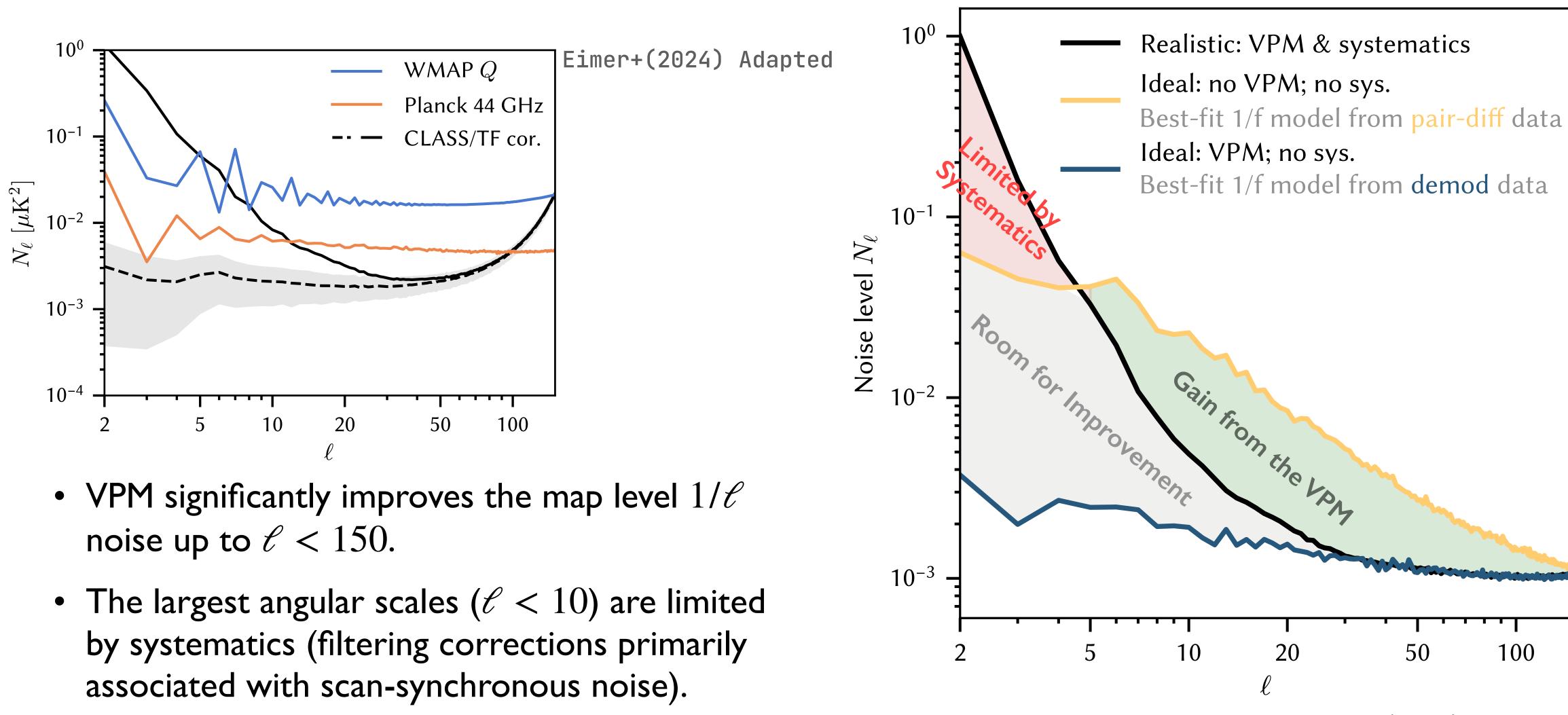


- Both the Earth's rotation and telescope scanning contribute to large angular scc
 (l > 20)
- Most of the largest angular scales (*l* > comes from the scanning frequency ar its harmonics.
- Systematics at these frequencies are more concerning than just 1/f.





Breakdown of the 40 GHz sensitivity



Cleary+(2024) Adapted





Summary

- VPM modulates linear (70% efficiency) and circular polarization.
- VPM increases low frequency mapping speed by over two orders of magnitude.
- VPM emission leaks into demodulated polarization signal
 - Does not significantly contribute to 1/f noise.
 - Contributes part of scan synchronous noise (mostly canceled with pairs).
- After I/f noise is mitigated, residual scan synchronous noise is the primary challenge to largest angular scale measurement.
 - Paper in preparation on 2016-2023 data on characterization of possible sources.
 - 2023-2024 higher elevation scans to probe ground pickup.
 - New RHWP modulator just installed to investigate modulator systematics.
 - Further study of atmospheric polarization.