

# Toward Fast BSM Physics with Bolt

*CMB-S4 Meeting 2023*

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(work w/ Zack Li)

# Motivation: Theoretical Progress

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These are familiar questions.

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Current data makes model extensions *hard!*

How do we facilitate new model development?

1. Remove standard assumptions
2. Make BSM extensions easy (Julia code, gradients)
3. Aim for (nuanced) “model discovery” with ML



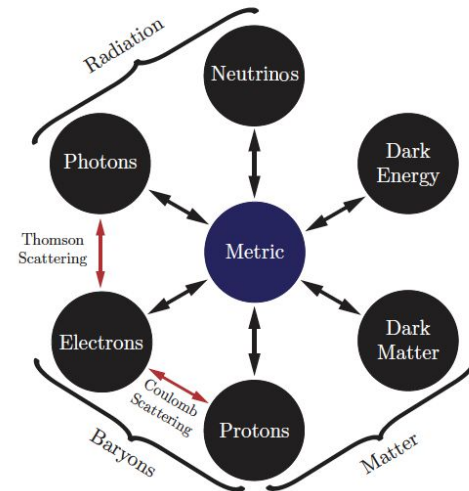
# E-B Solver - the Cornerstone of Cosmology

LSS and the CMB models evolve perturbations described by GR and the Boltzmann eqn. (Einstein-Boltzmann system)

Accurate modeling requires solving this (stiff) ODE system

Can address stiffness in two ways:

- Approximations
- Smart ODE solver



# Beyond State of the Art

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SOTA codes CAMB & CLASS are incredible feats of scientific software engineering

However:

- both rely on approximations
- neither is differentiable
- navigating the source is a time commitment

The future is differentiable and easy to modify

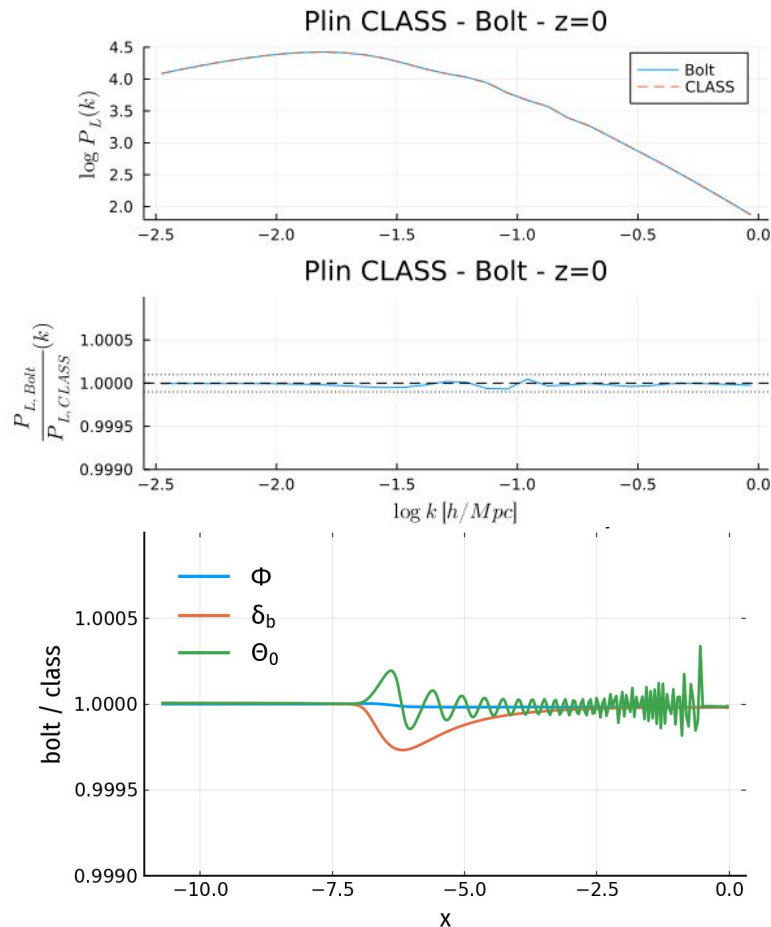
**-> Bolt.jl**

# Agreement at the 0.1% Level

CLASS/CAMB are consistent at the 0.1% level in CMB and matter power spectra

Bolt joins\* this exclusive club!

We agree on the perturbations for individual k-modes too



\*see also, Moser++22

# A Need for Differentiable Models

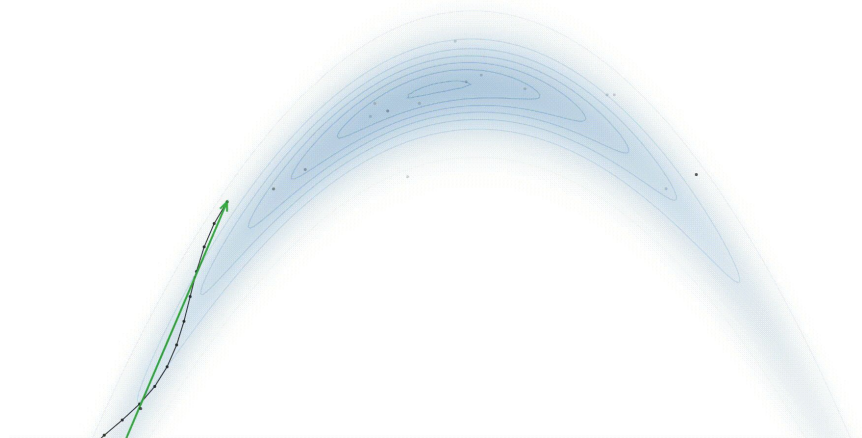
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More parameters means high-dimensional inference

Gradient-based methods make this tractable

Must use model gradients, even when not analytic!

Can make an approximation\*,  
or use a differentiable model\*\*



\*e.g. Mancini++21, DeRose++21, Colas++20, Wibking++20

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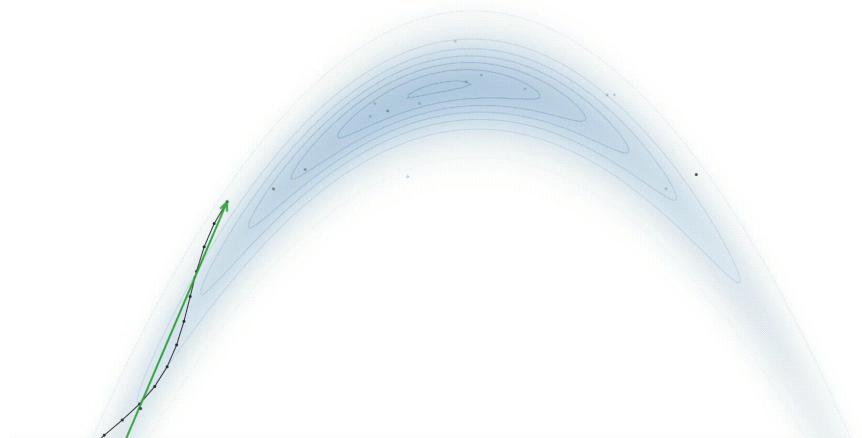
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# Gradients on Computers

## Symbolic Differentiation

(aka Diff by grad student...or Mathematica)

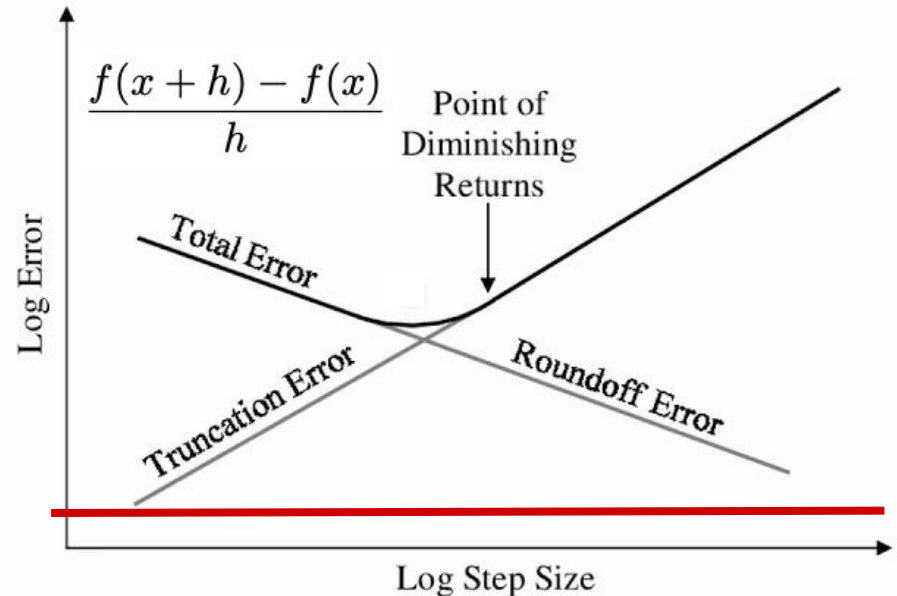
Doesn't work for complex models

## Finite Difference

(frequently used for  
Fisher forecasts)

But what is the step?!

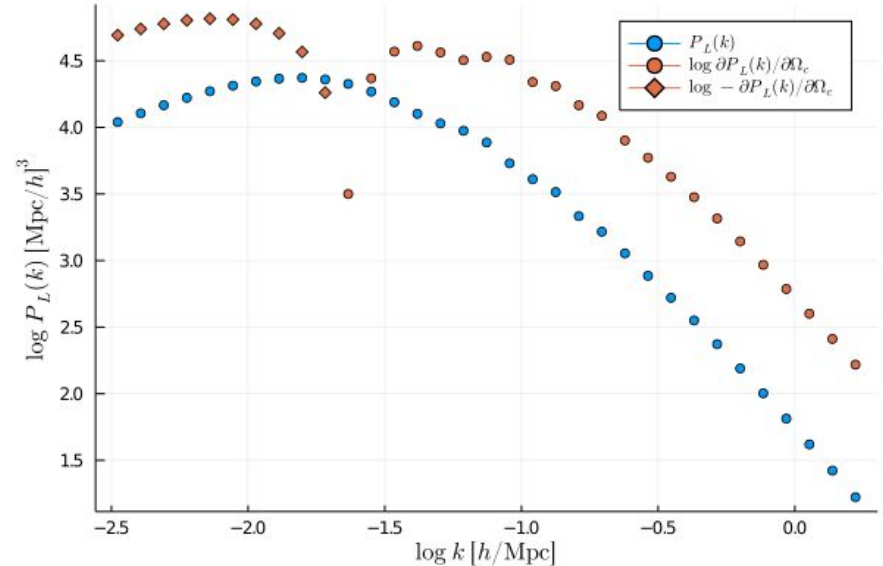
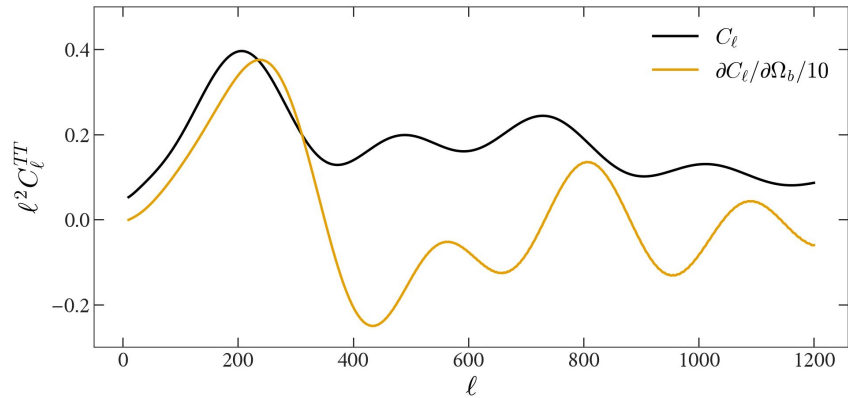
A better way?





# Gradients for the CMB and LSS

*Exact* gradients of the CMB anisotropy, linear matter power spectra

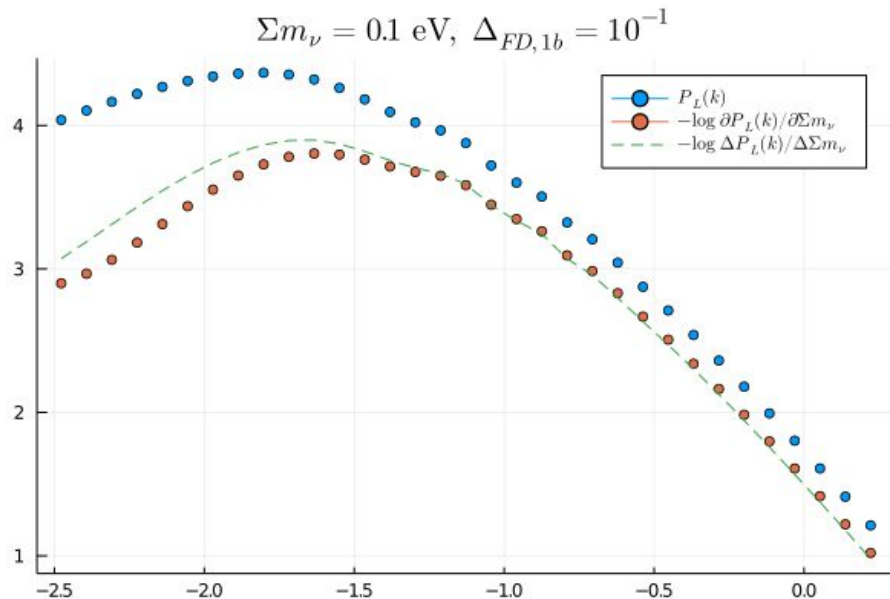


# AD >> FD

Finite difference (FD) gives  
gradient approximations

Well-known that FD does not  
perform well for neutrino mass

*Exact* AD derivatives solve this  
problem!



# Bolt.jl Design Recap

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Goals:

1. Differentiability
2. Promote rapid physics prototyping

Design aspects:

1. Julia & automatic differentiation (AD)
2. Minimal approximations

# Bolt - Model Extension “Live” Demo

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Add dark radiation with an interaction a la Lesgourgues++2015\*:

$$\dot{\delta}_{\text{dm}} = -\theta_{\text{dm}} + 3\dot{\phi}$$

$$\dot{\theta}_{\text{dm}} = -\frac{\dot{a}}{a}\theta_{\text{dm}} + a\Gamma(\theta_{\text{dr}} - \theta_{\text{dm}}) + k^2\psi$$

$$\dot{\delta}_{\text{dr}} = -\frac{4}{3}\theta_{\text{dr}} + 4\dot{\phi}$$

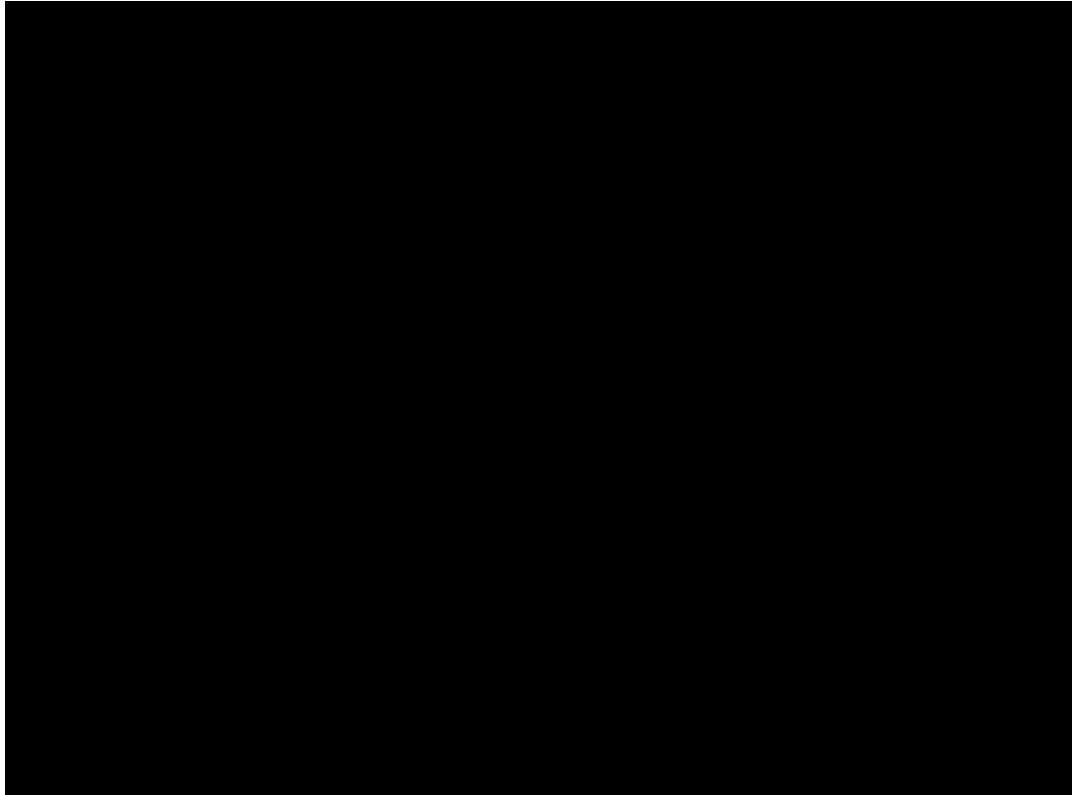
$$\dot{\theta}_{\text{dr}} = k^2\frac{\delta_{\text{dr}}}{4} + k^2\psi + \frac{3}{4}\frac{\rho_{\text{dm}}}{\rho_{\text{dr}}}a\Gamma(\theta_{\text{dm}} - \theta_{\text{dr}})$$

\*Thanks to Ben Wallisch for pointing me to this paper

# Bolt - Model Extension “Live” Demo

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A few minutes and 17 lines to add DR and run HMC!



# A new workflow for linear cosmology?

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How do we learn unknown physics?

Background has been explored, but now perturbations!

Embed neural network *inside* ODE function:

$$\delta'_i(k, u, x) = NN_1^{(\theta)}(u, x)$$

$$v'_i(k, u, x) = NN_2^{(\theta)}(u, x)$$

As flexible as you desire! (or dare)



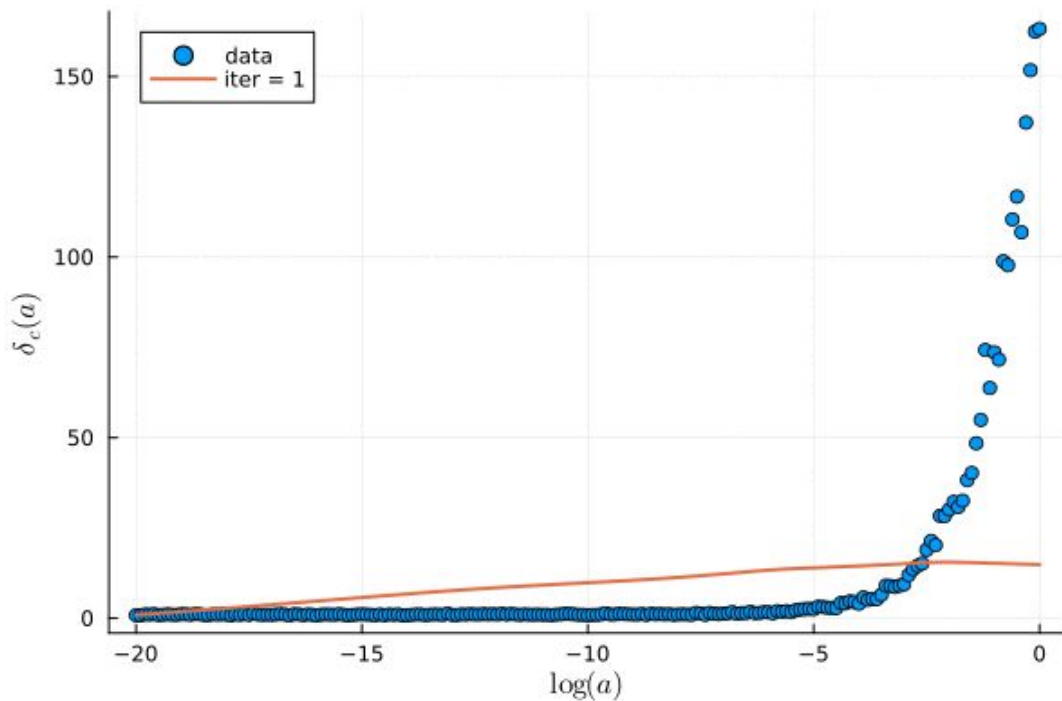
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Pretend we “forgot” CDM linear theory - can it be learned?

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**Yes!\***

# What's next? - NN Uncertainty!

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NNs in ODEs a step forward in flexibility

But does **not** tell you **where** to focus model-building efforts

Workflow goal is to guide **human** model building

Luckily, we can obtain uncertainty estimates for neural network predictions - current work!

# Summary

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Bolt **differentiably solves** the stiff E-B system for the **first time**

Leverage this to **relax standard assumptions**

Lack of approximations and readability helps **quickly try new physical models**

Flexible neural network models will **guide model builders** toward **missing physics**

Differentiability indispensable for high-dimensional inference