

Astronomy at 100 below: Going to the ends of the Earth to see the beginnings of time

photo by
Aman
Chokshi,
PhD student
& winterover,
2022

Christian Reichardt
U. Melbourne

on behalf
of the

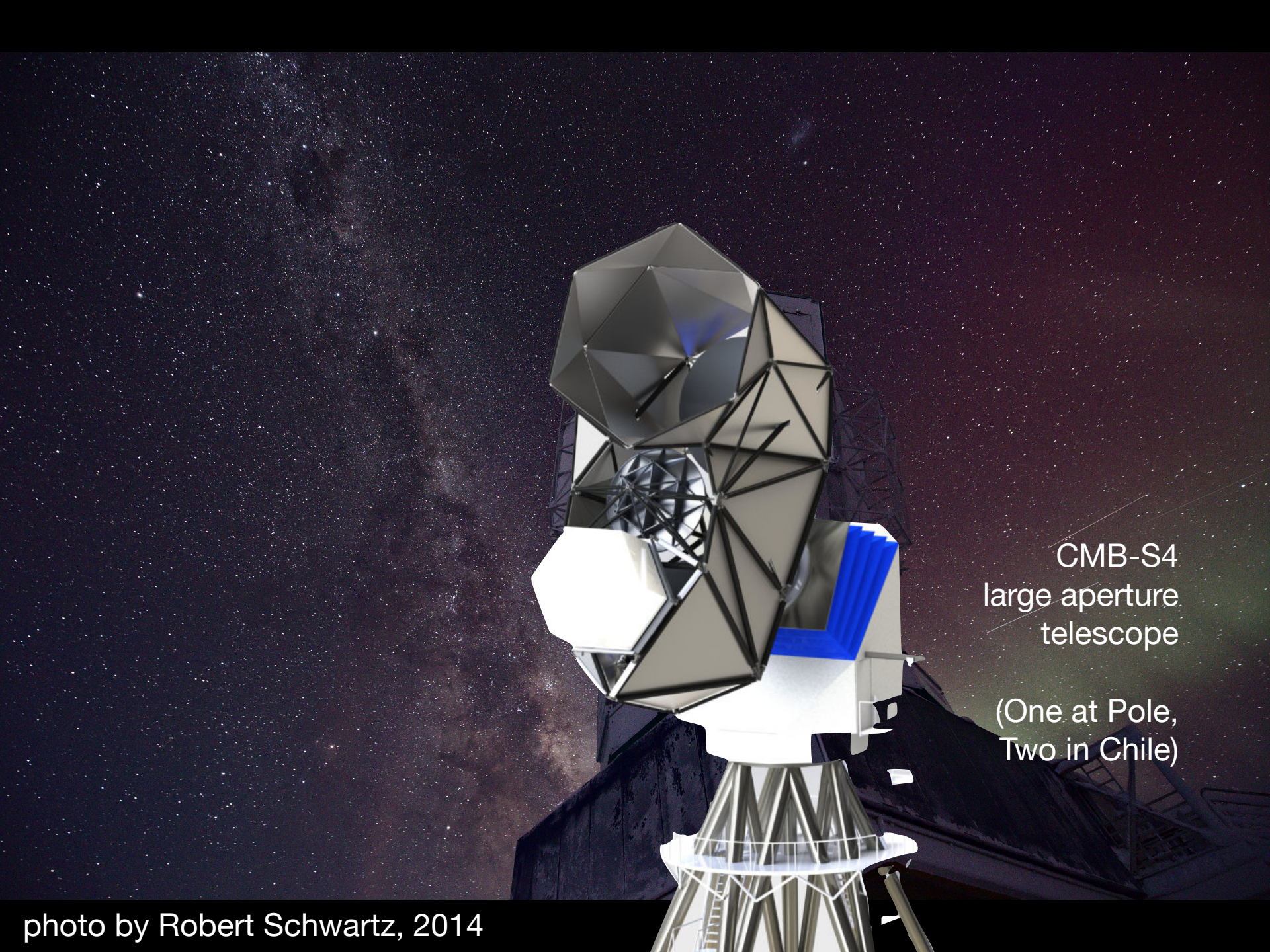


collaboration



South Pole
Telescope

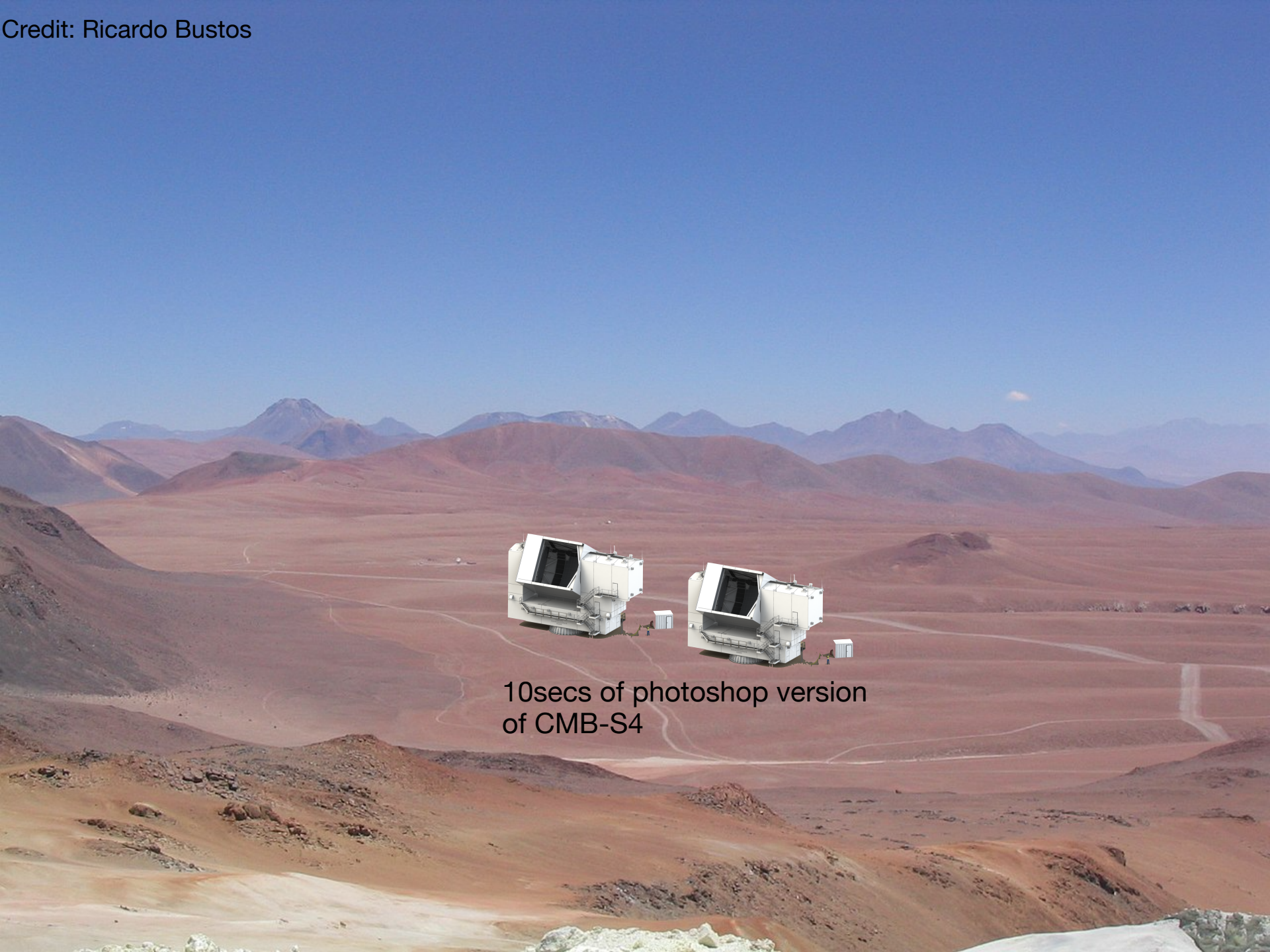
photo by Robert Schwartz, 2014



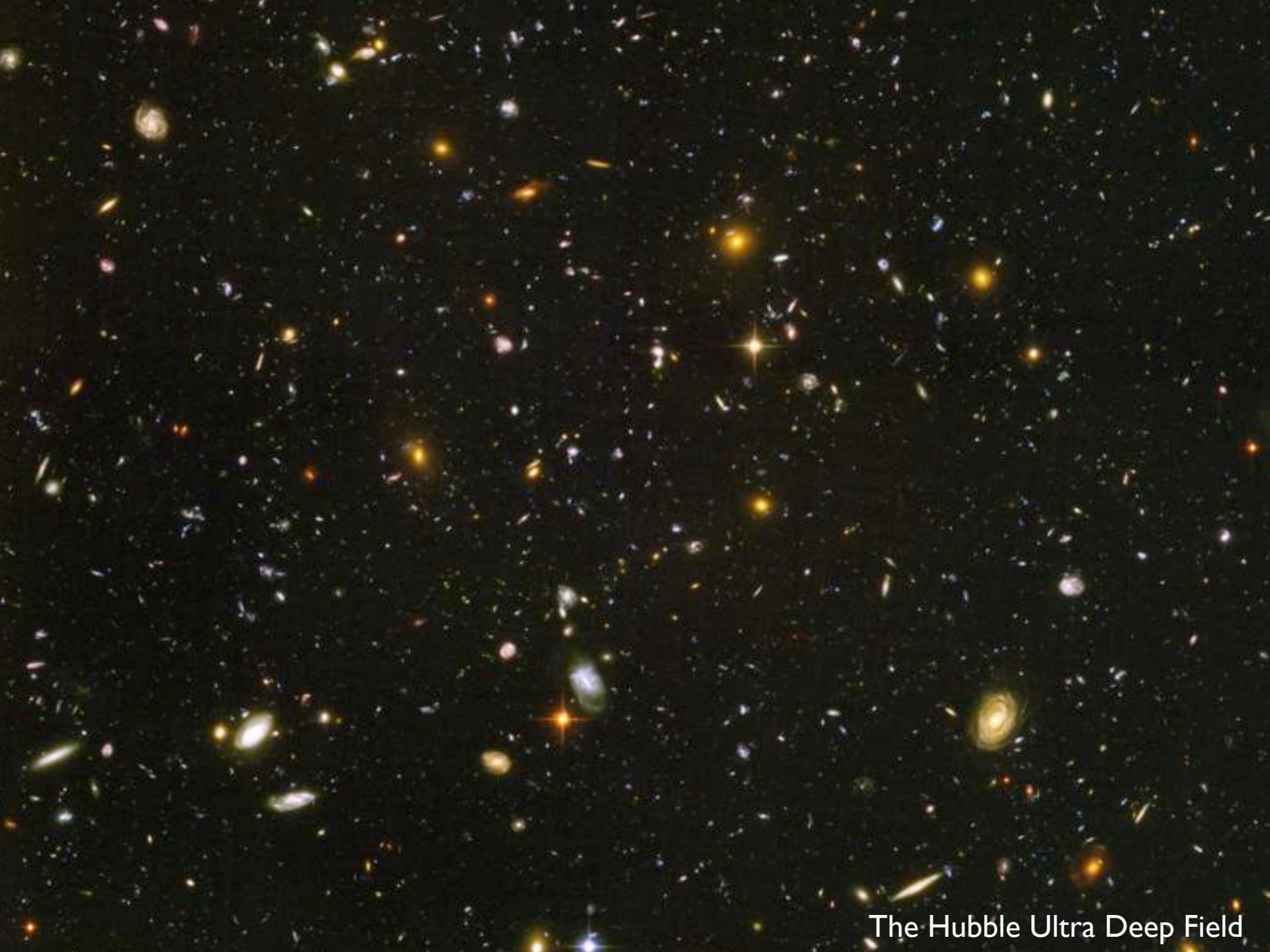
CMB-S4
large aperture
telescope

(One at Pole,
Two in Chile)

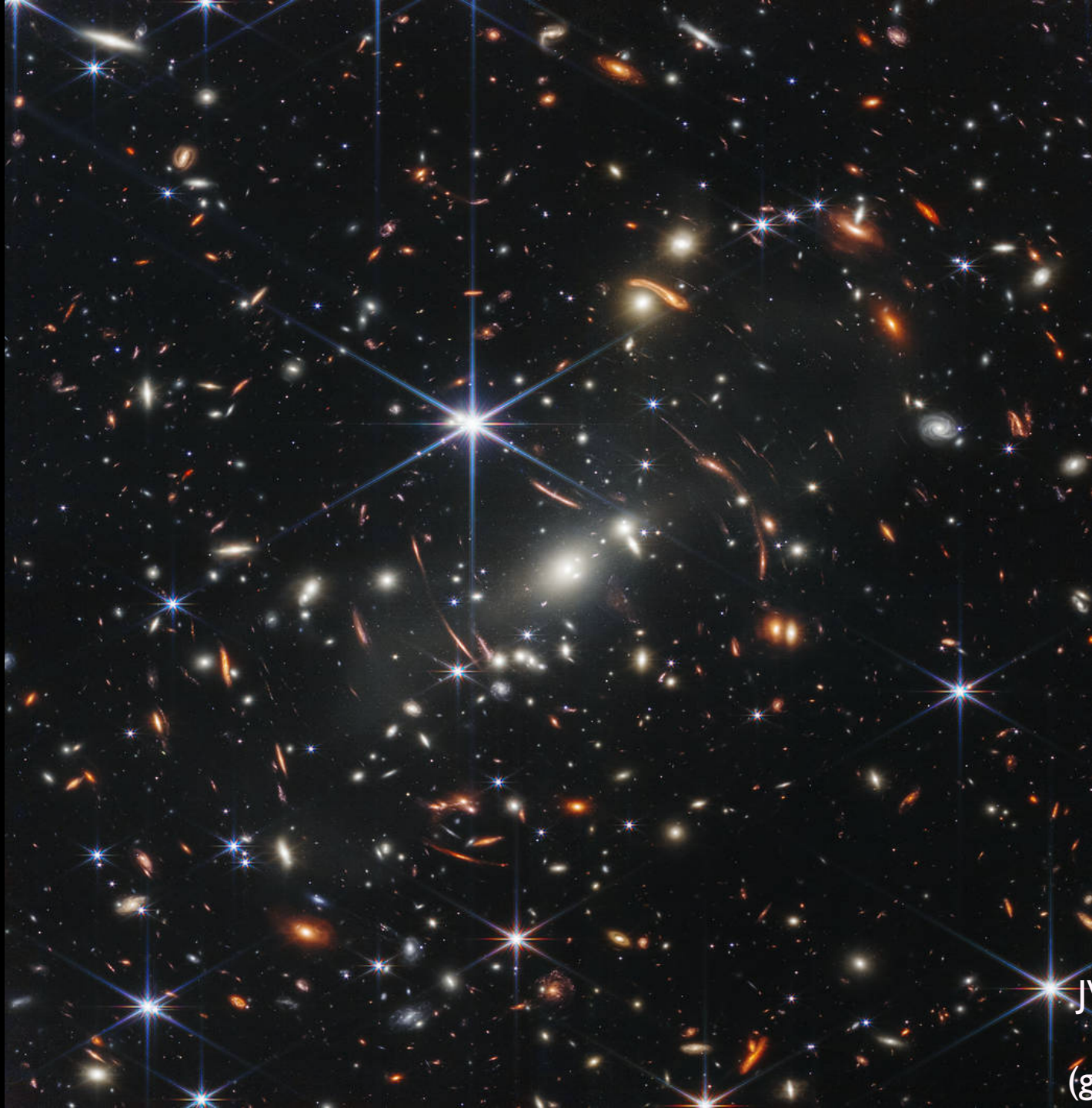
Credit: Ricardo Bustos



10secs of photoshop version
of CMB-S4



The Hubble Ultra Deep Field



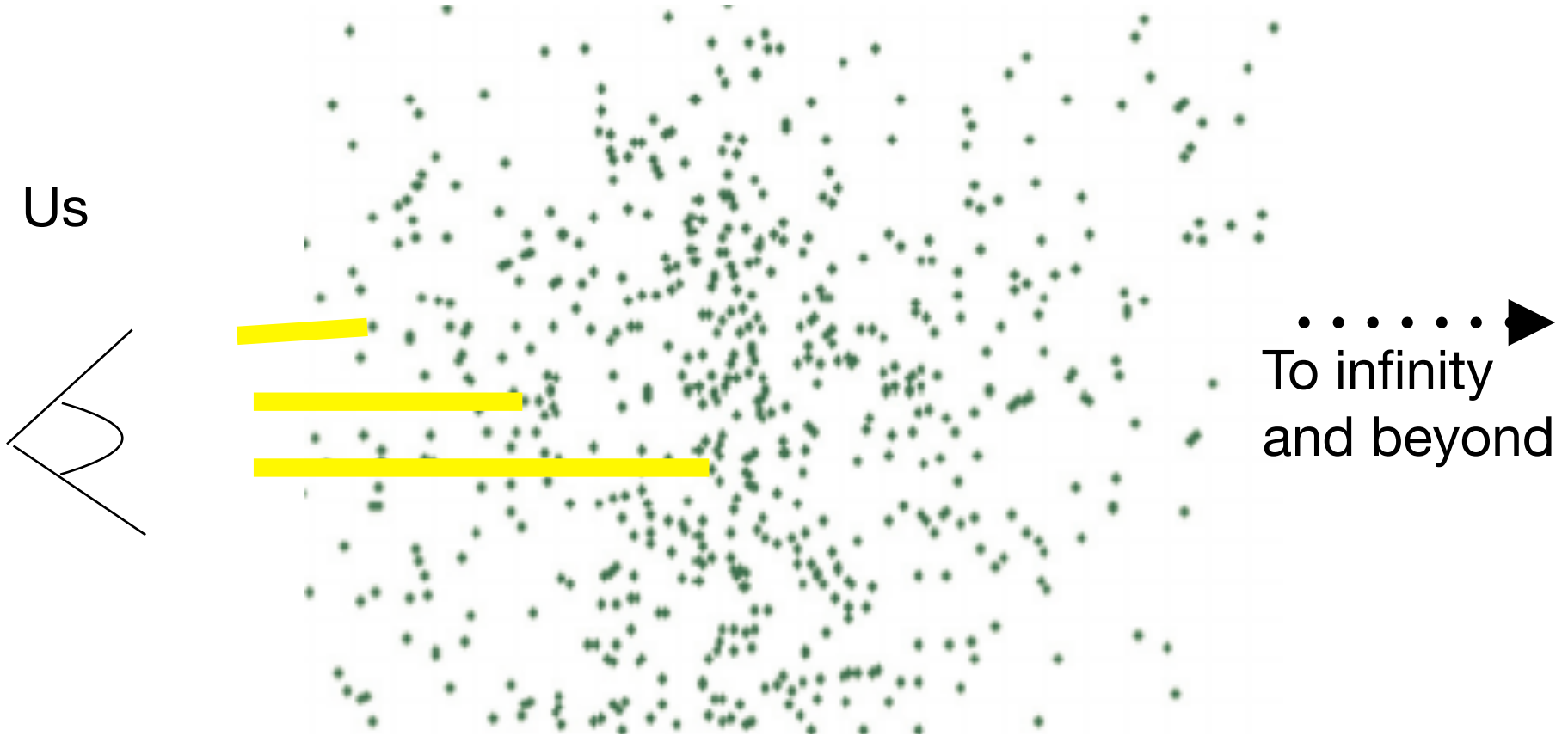
JWST - Webb
Deep Field
(galaxy cluster)



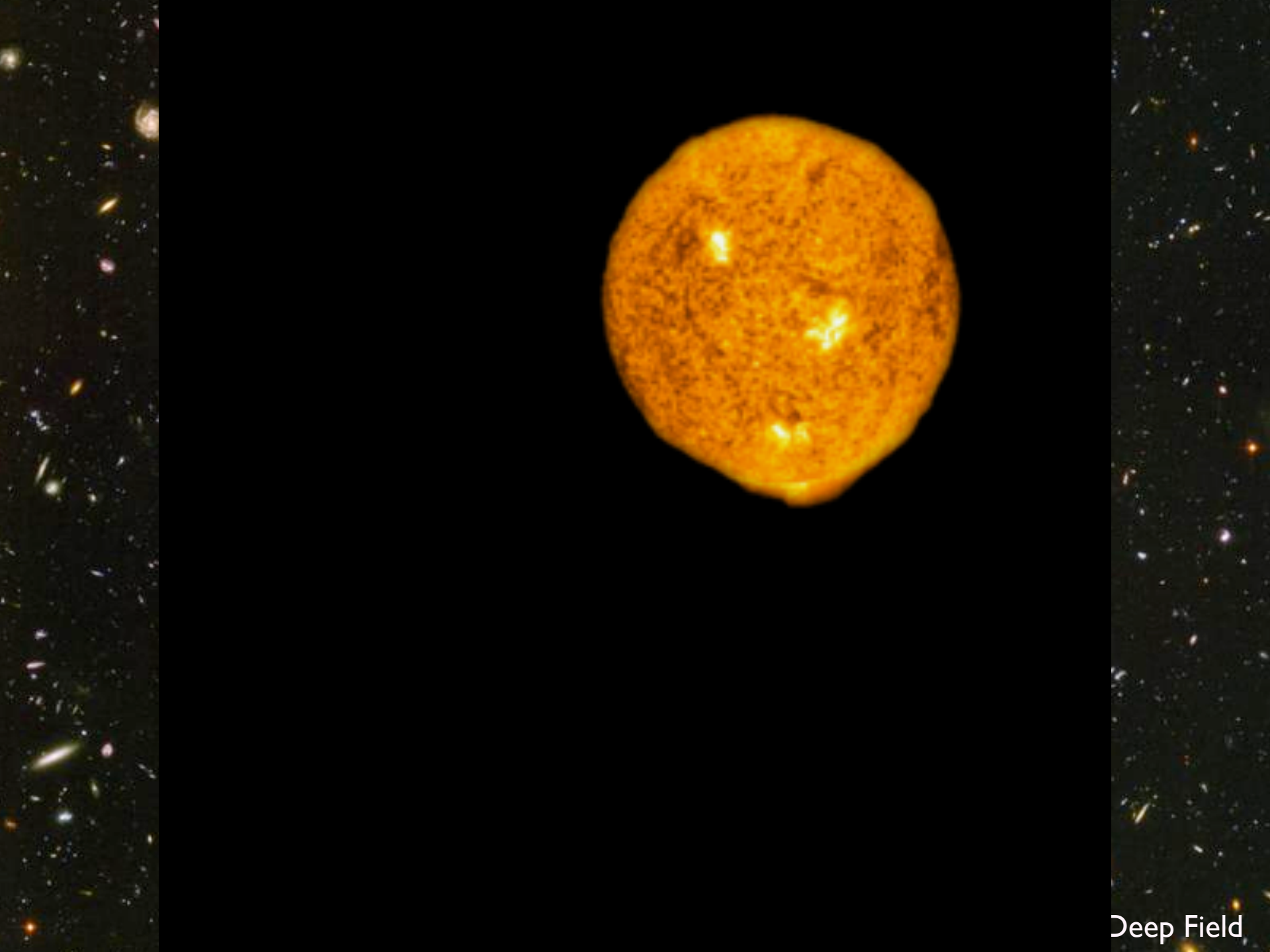
Why is most of it dark???

JWST - Webb
Deep Field
(galaxy cluster)

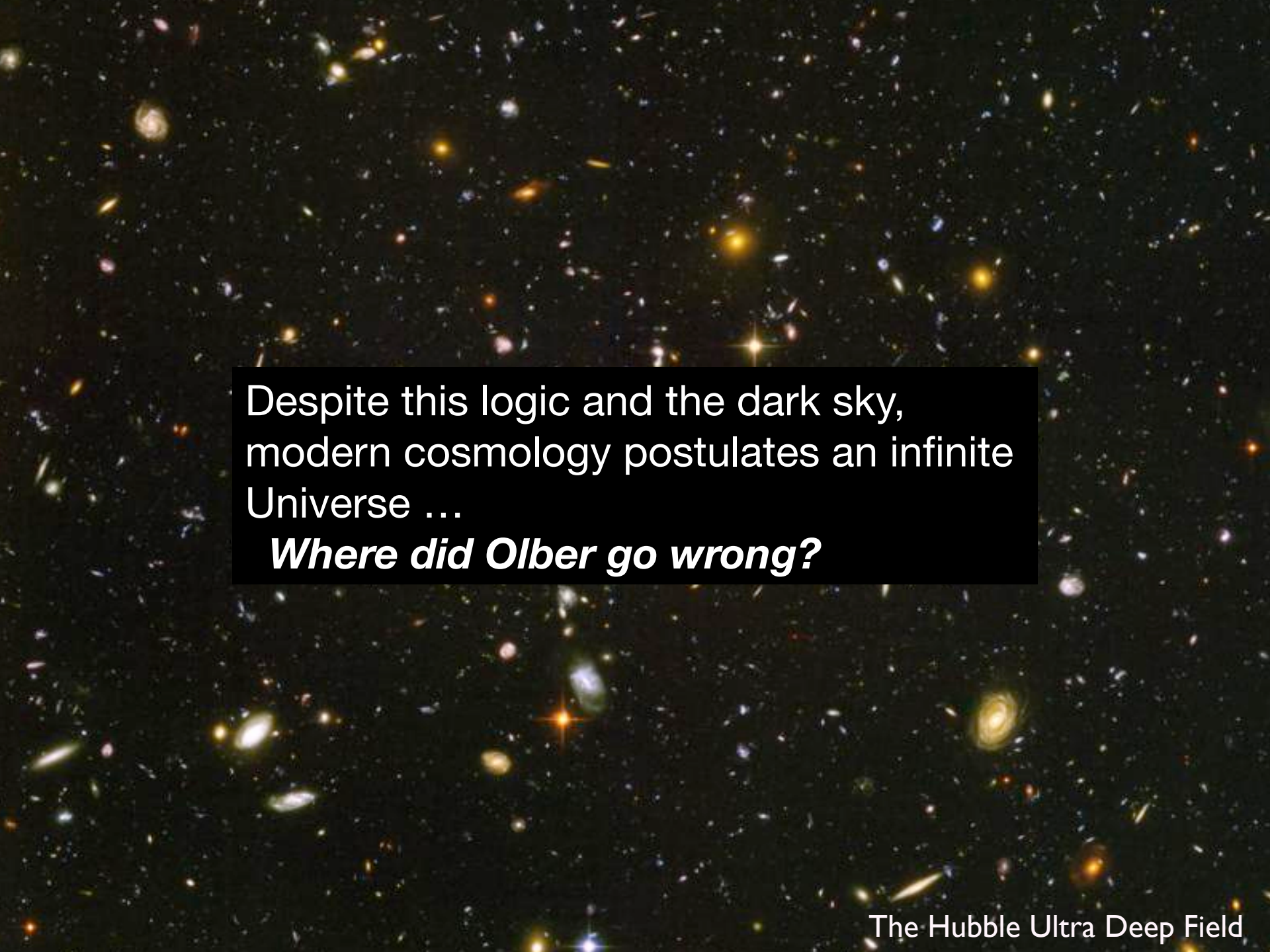
Olber's Paradox



In an infinite cloud of stars,
all rays hit a star



Deep Field

The background of the slide is a vast field of galaxies, known as the Hubble Ultra Deep Field. It shows a dense collection of galaxies in various colors (yellow, orange, blue, white) and shapes (spiral, elliptical, irregular), scattered across a dark cosmic background. The galaxies are of various sizes and orientations, creating a rich, multi-colored tapestry of the universe.

Despite this logic and the dark sky,
modern cosmology postulates an infinite
Universe ...

Where did Olber go wrong?

The background of the slide is a vast field of galaxies, known as the Hubble Ultra Deep Field. It shows a dense collection of galaxies in various colors, including yellow, orange, red, blue, and purple, scattered across a dark cosmic background. The galaxies vary in size and shape, from small, distant points of light to larger, more complex structures.

Point 1: The speed of light is finite.

**Telescopes are time machines
this image goes 13 billion years into the past**

The Hubble Ultra Deep Field

The background of the slide is a vast field of galaxies, known as the Hubble Ultra Deep Field. It shows a dense collection of galaxies in various colors, including yellow, orange, red, and blue, scattered across a dark cosmic background. The galaxies vary in size and shape, from small, distant points of light to larger, more complex structures.

Broken assumption #1: The Universe
isn't eternal.

Current age: **14 billion years**

The background of the slide is a vast field of galaxies, known as the Hubble Ultra Deep Field. It shows a dense collection of galaxies in various colors (yellow, orange, blue, white) and shapes (spiral, elliptical, irregular), scattered across a dark cosmic background. The galaxies are of various sizes and orientations, creating a rich and complex visual texture.

Broken assumption #1: The Universe
isn't eternal.

Current age: **14 billion years**

and the

Oldest observed star ~ **13 billion years.**

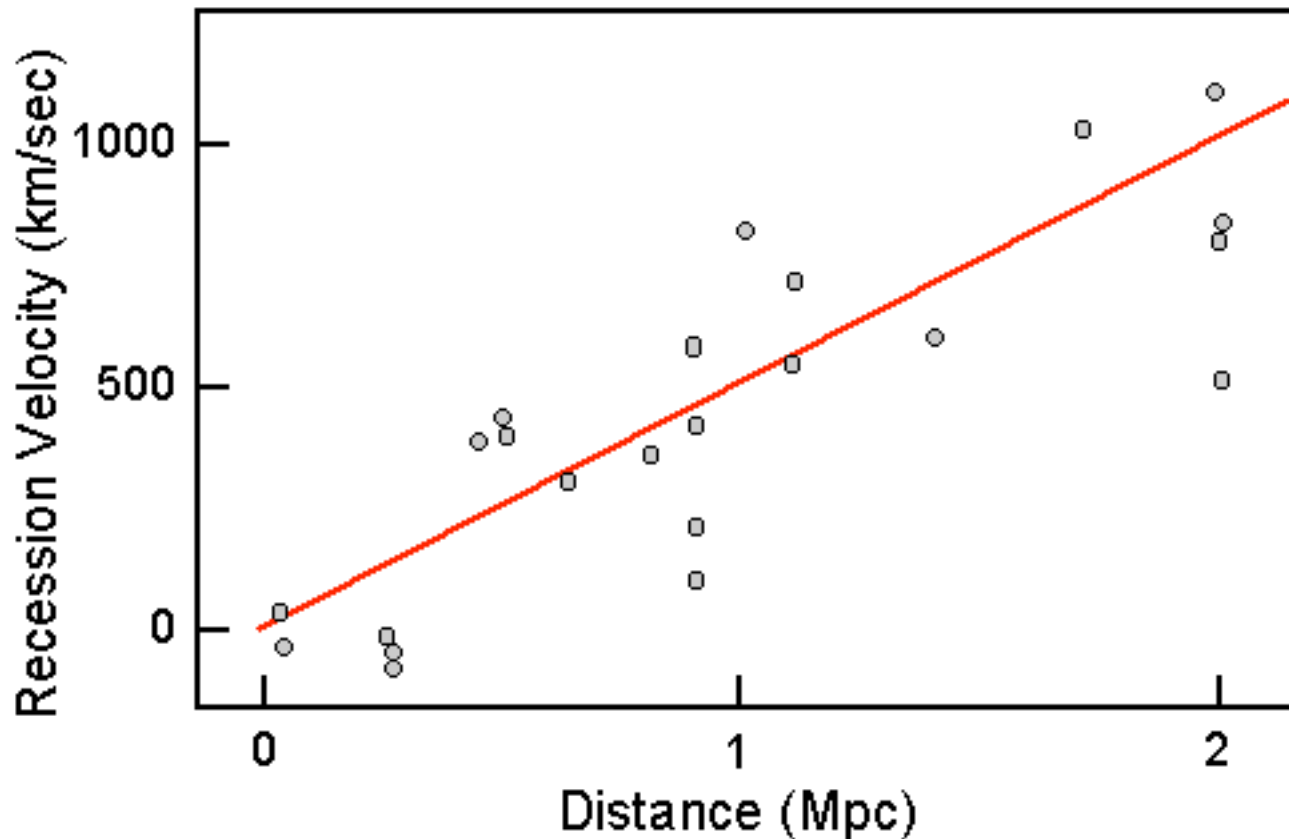
The background of the slide is a vast field of galaxies, known as the Hubble Ultra Deep Field. It shows a dense collection of galaxies in various colors, including yellow, orange, red, blue, and purple, scattered across a dark, starry sky. The galaxies vary in size and shape, from small, distant points of light to larger, more complex structures.

Broken assumption #2: The Universe
isn't static.

It's expanding.

First noticed by Edwin Hubble

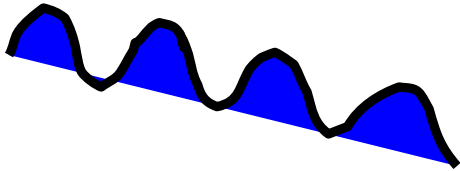
Hubble's Data (1929)



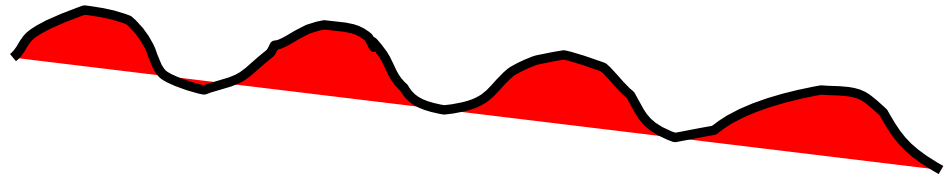
1. Everything is moving away from us...
2. Further away objects move faster

The Universe is expanding!

Cosmological redshift (z)



universe expands...



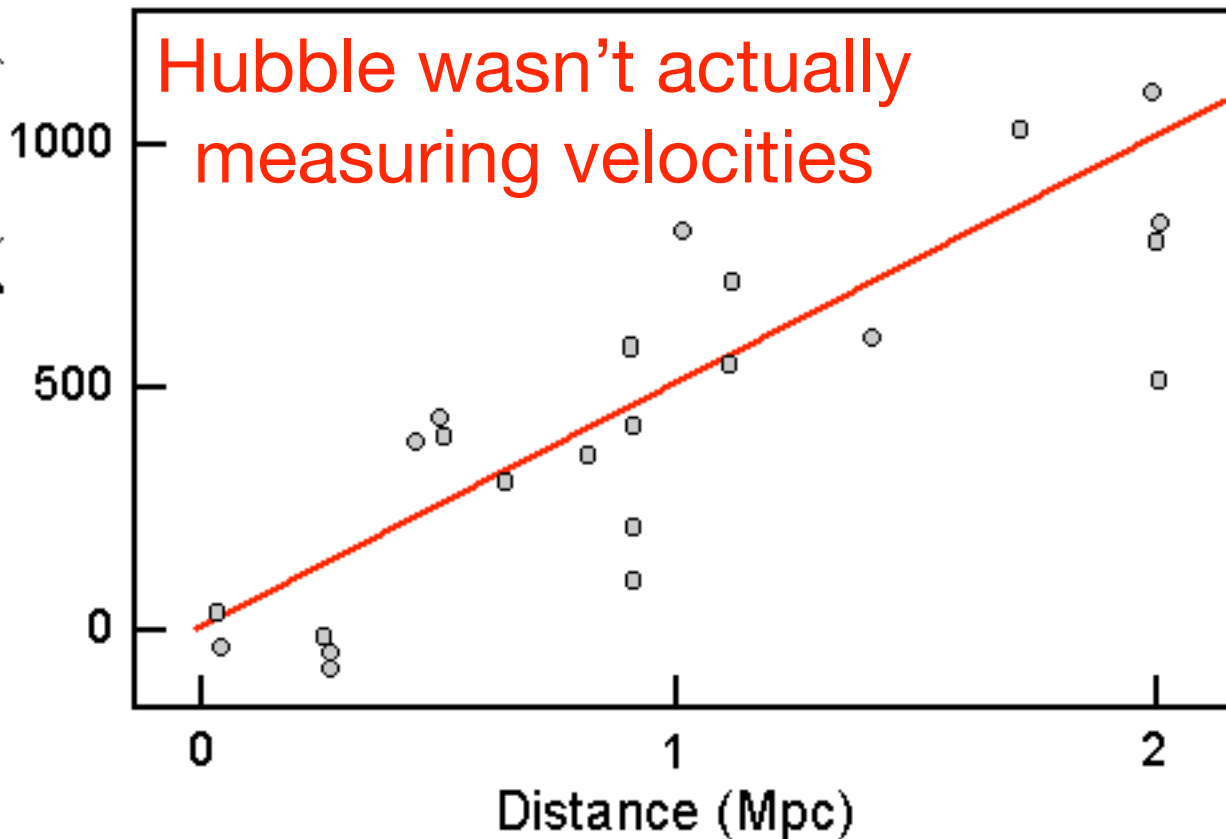
photon wavelength stretched *or*
redshifted

Expanding Universe

Hubble's Data (1929)

Redshift

Recession Velocity (km/sec)



1. A measure of how much larger the Universe is today than when the light started its journey to us

George Lemaitre first
proposed this in 1931

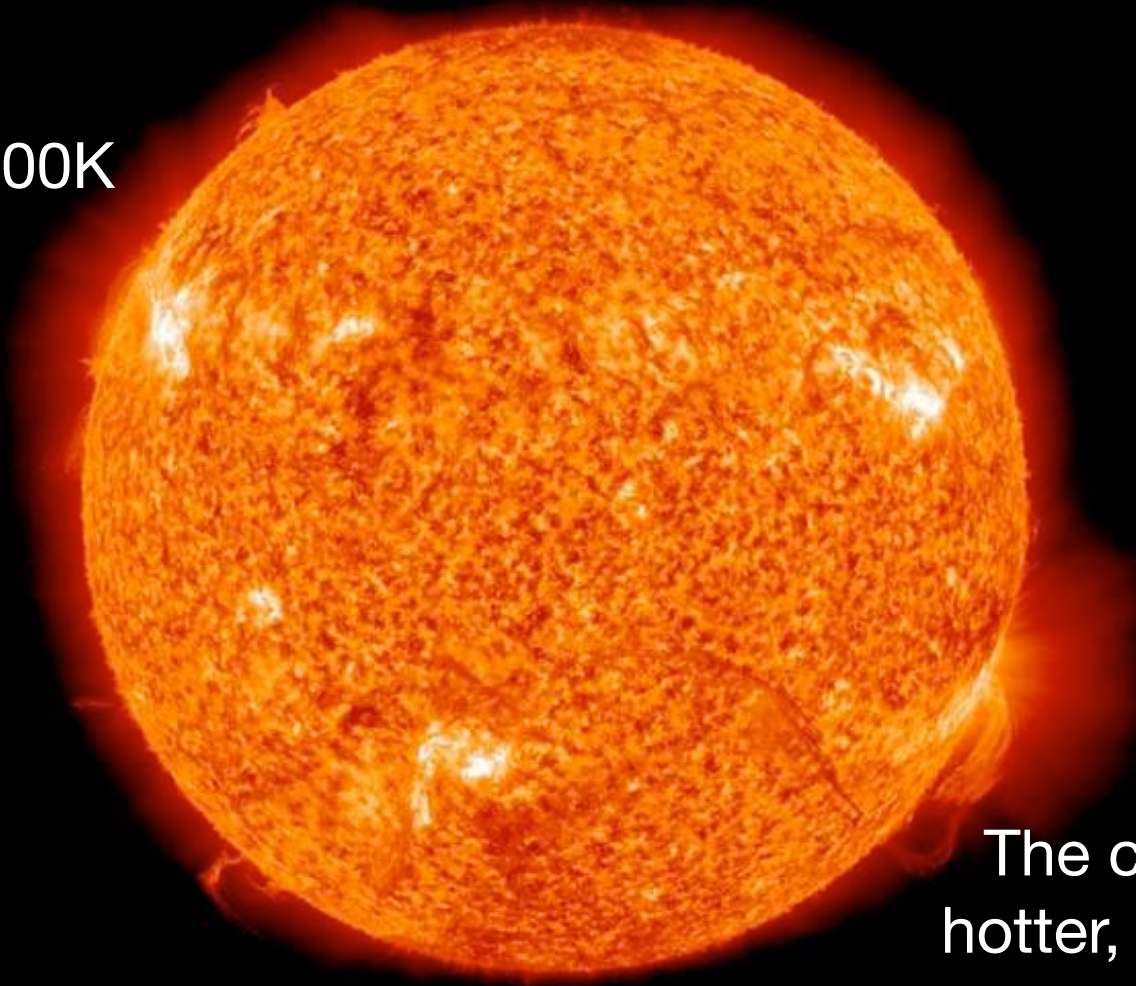


There is a beginning

- Run the clock backwards
 - Galaxies converge together
 - Universe gets smaller, denser, hotter
 - Till you hit the singularity
- The Universe has a finite age!

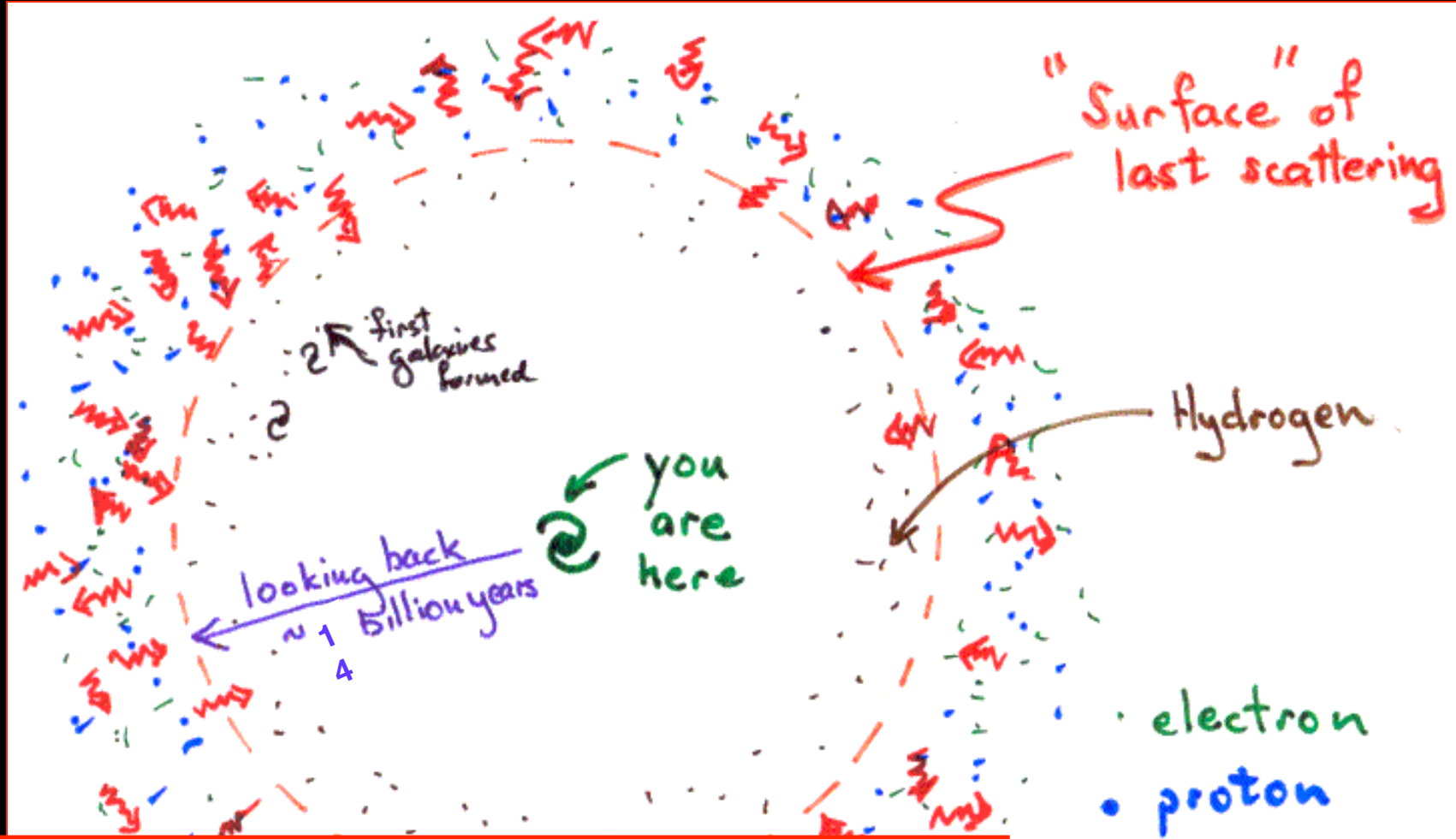
Can we see the beginning?

Our sun:
We see the
surface ~ 5500K



The core is 3000x
hotter, but shielded
from view

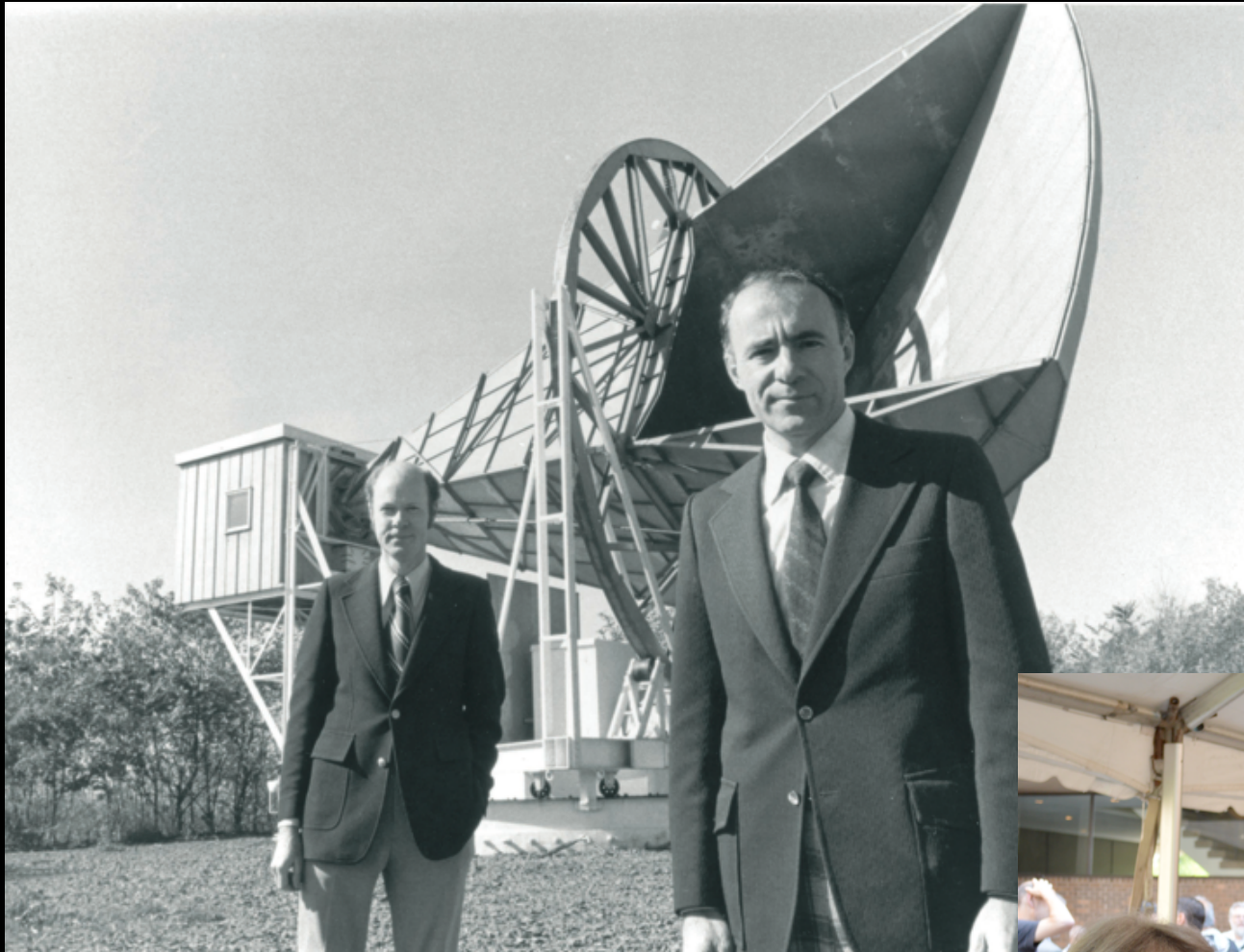
We are inside a glowing sphere



We should see the hot plasma!
Expected temperature (back then):
half the Sun's

First predicted
in 1948

Discovery of the Cosmic Microwave Background



Arno Penzias & Robert Wilson in front of the 20ft Bell Labs antenna used to discover the microwave background in 1965

**direct evidence
for a Hot Big
Bang**

1978 Nobel prize



**Bob Wilson @ anniversary
2015**

A better place to look

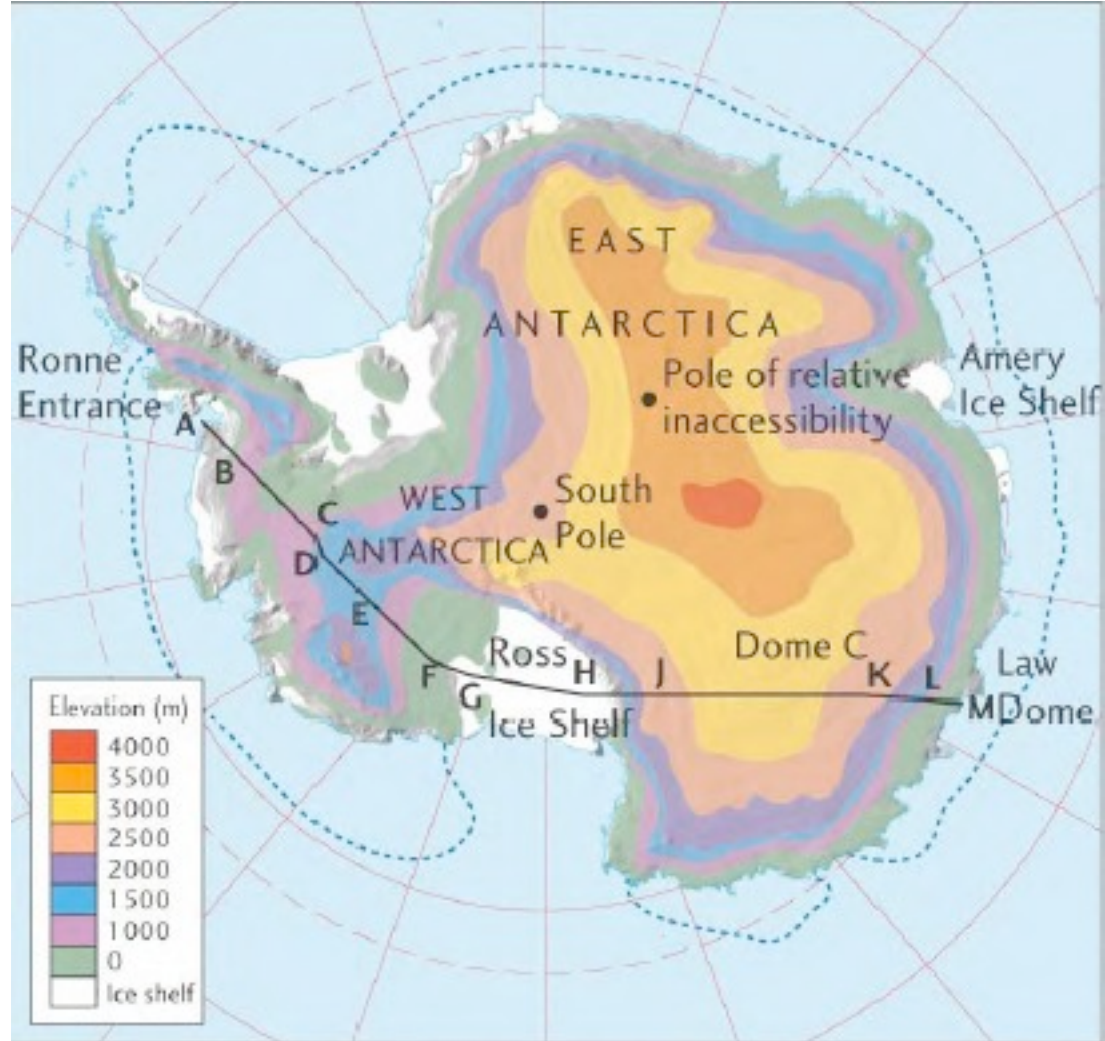


South Pole Station

Why Antarctica?

1. Height

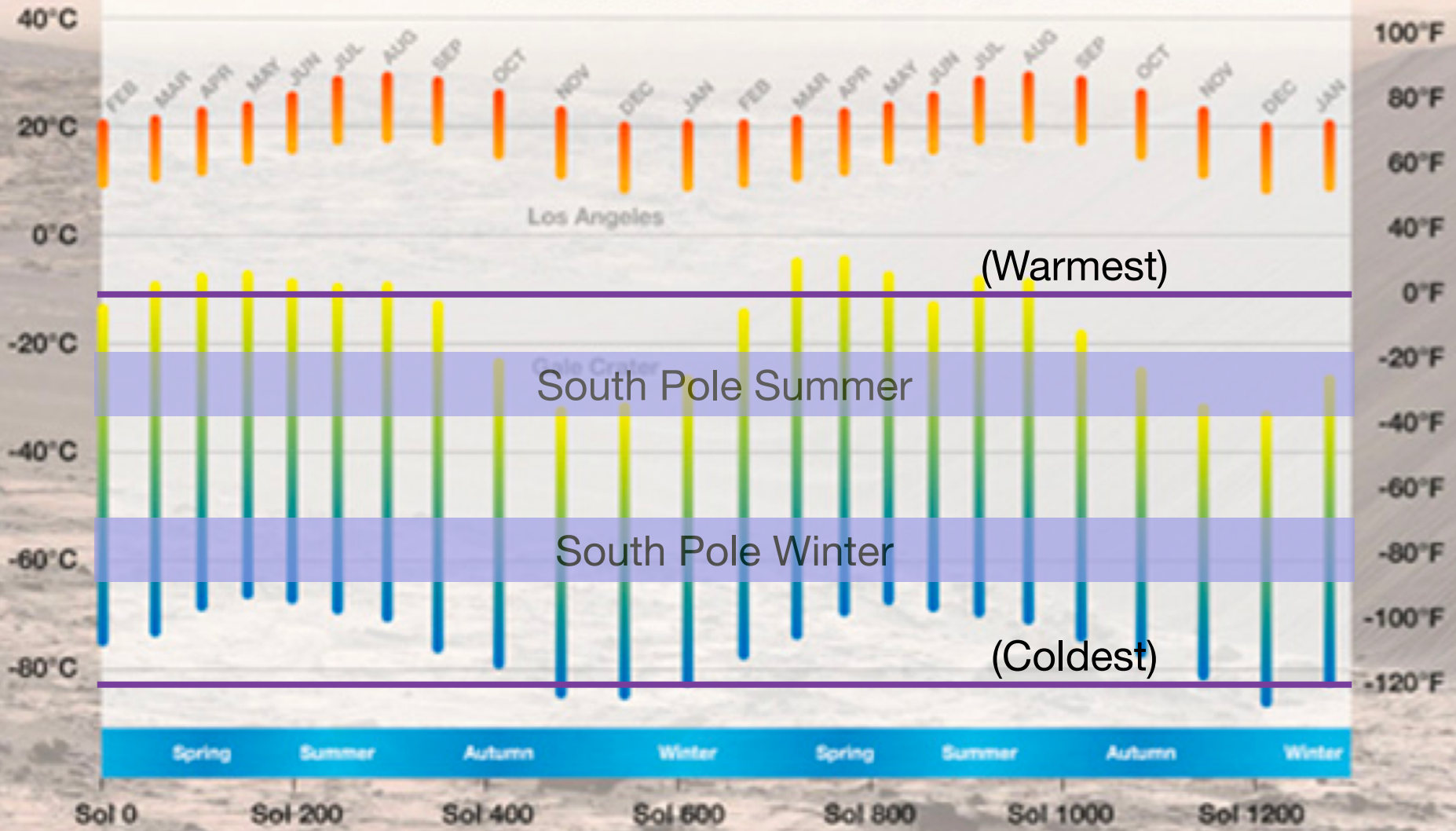
- Highest point:
 - 4.9 km
- Lowest:
 - -2.6 km (+4 km of ice)
- The ice is *really thick*. Plateau averages 3km (SLAC is 86m above sea level)



Reason 2: Martian Temperatures

Credit: NASA/JPL

Seasonal Temperature Ranges at Gale Crater
(with temperatures in Los Angeles at equivalent seasonal points)



the cold means its **Very Dry!**

Precipitation:

Outback deserts: <15 cm/year

Sahara: 10 cm/year

South Pole: 7cm/year



Effectively a **desert** despite sitting on 2km of ice
which matters because

Water is terrible for microwave astronomy

Why is water bad for microwave astronomy?



Remember how your microwave works...

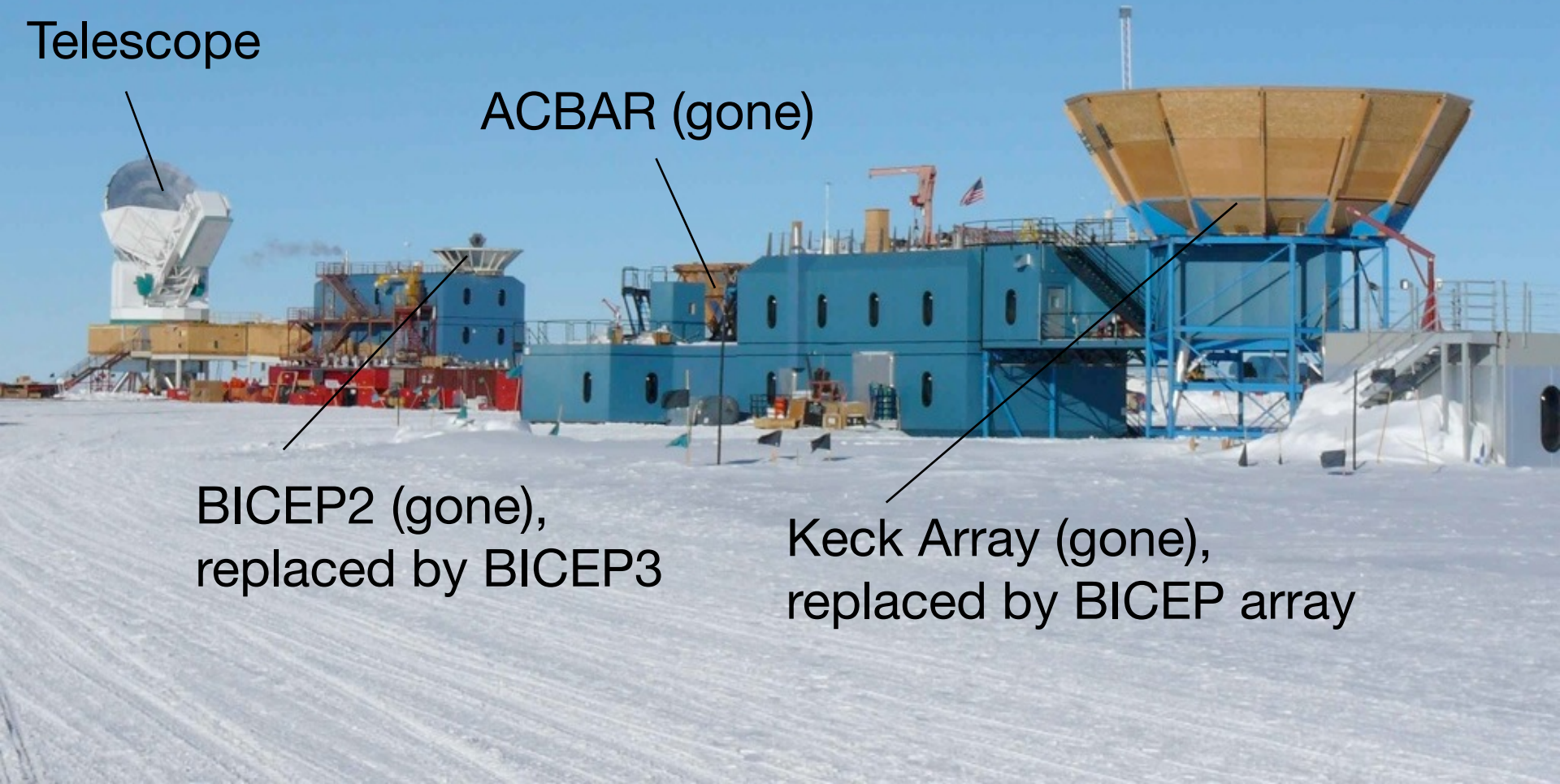
CMB at Pole

South Pole
Telescope

ACBAR (gone)

BICEP2 (gone),
replaced by BICEP3

Keck Array (gone),
replaced by BICEP array



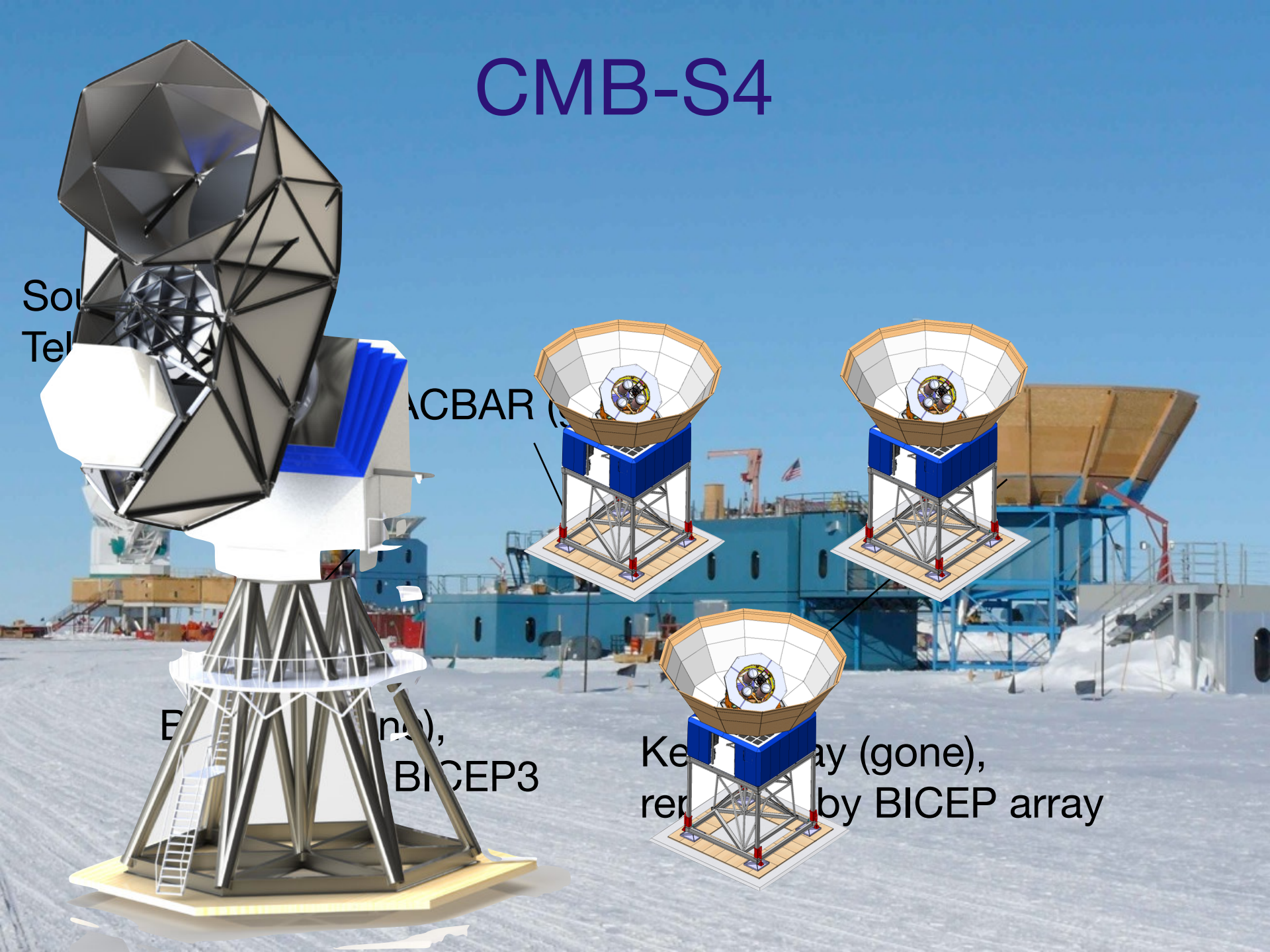
CMB-S4

Sol
Tel

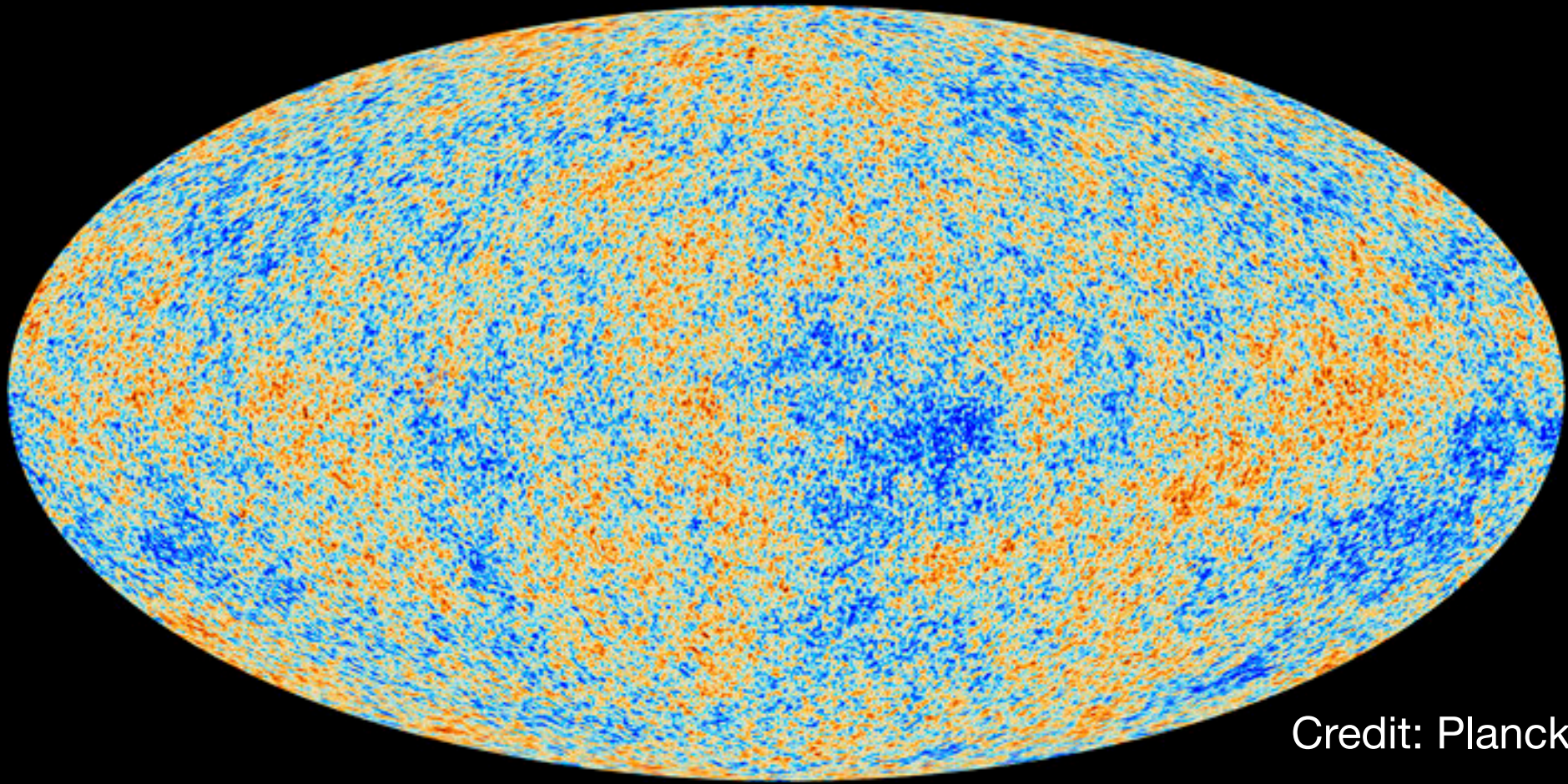
ACBAR

BICEP3

Key array (gone),
replaced by BICEP array



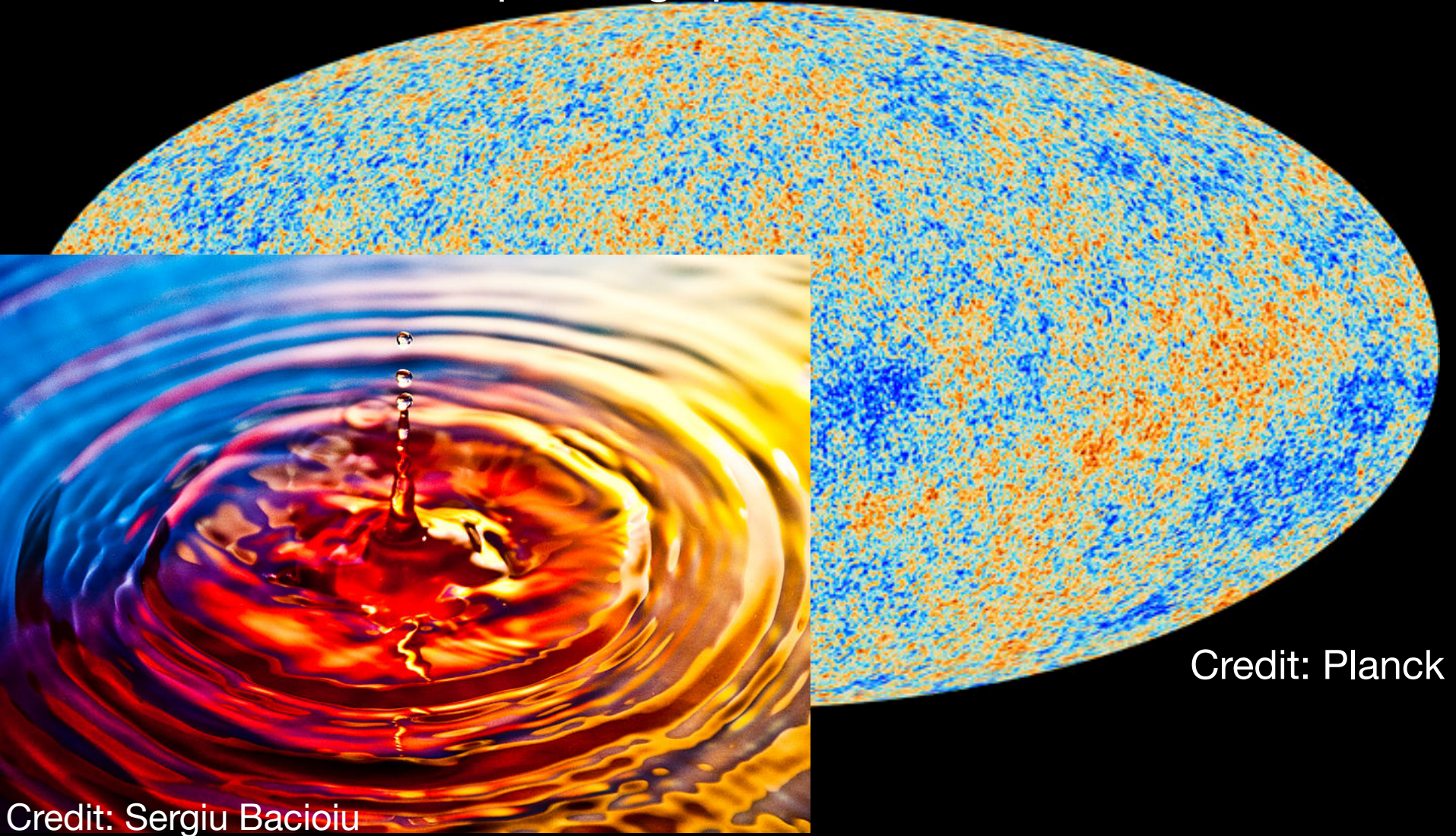
The goal: exquisite maps of the
Universe just after the Big Bang



Credit: Planck

What are we looking at?

Lots of expanding spherical sound waves



Credit: Planck

Credit: Sergiu Bacioiu

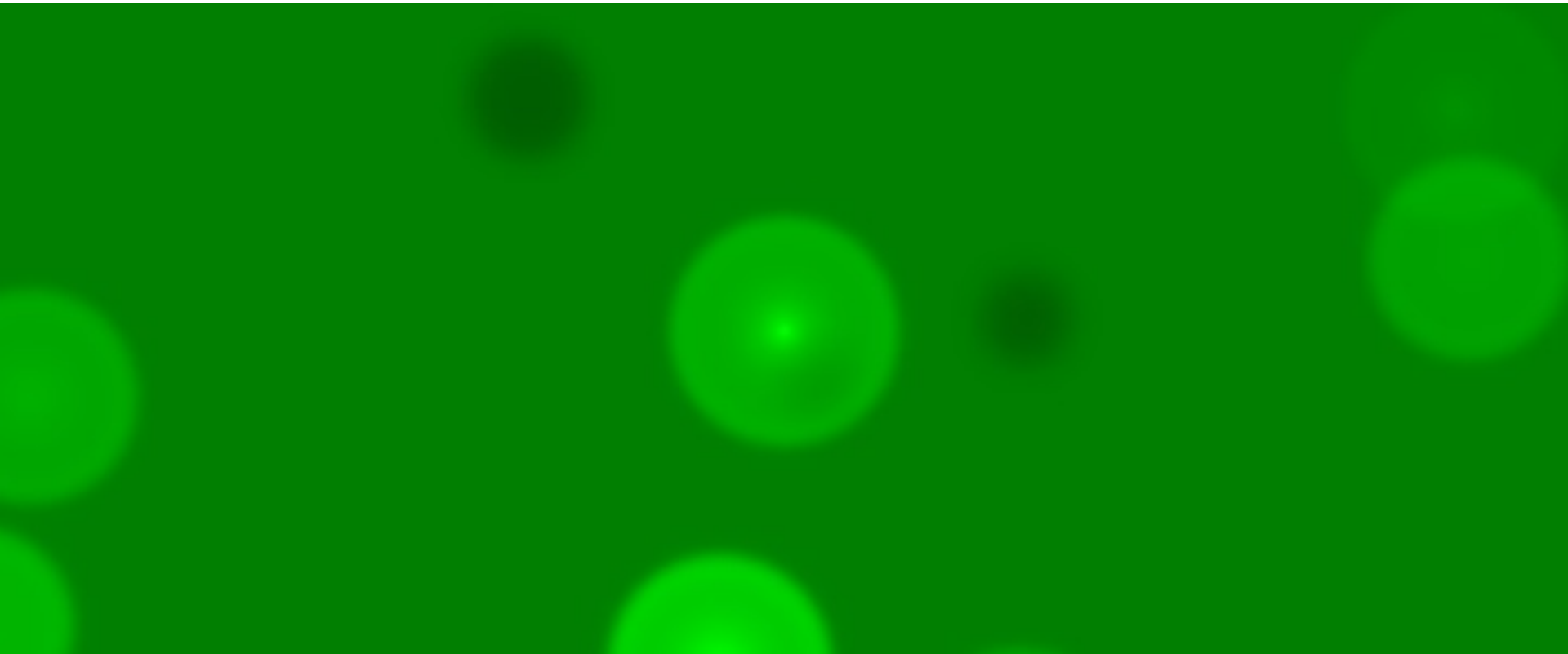
Adding “pebbles”

1



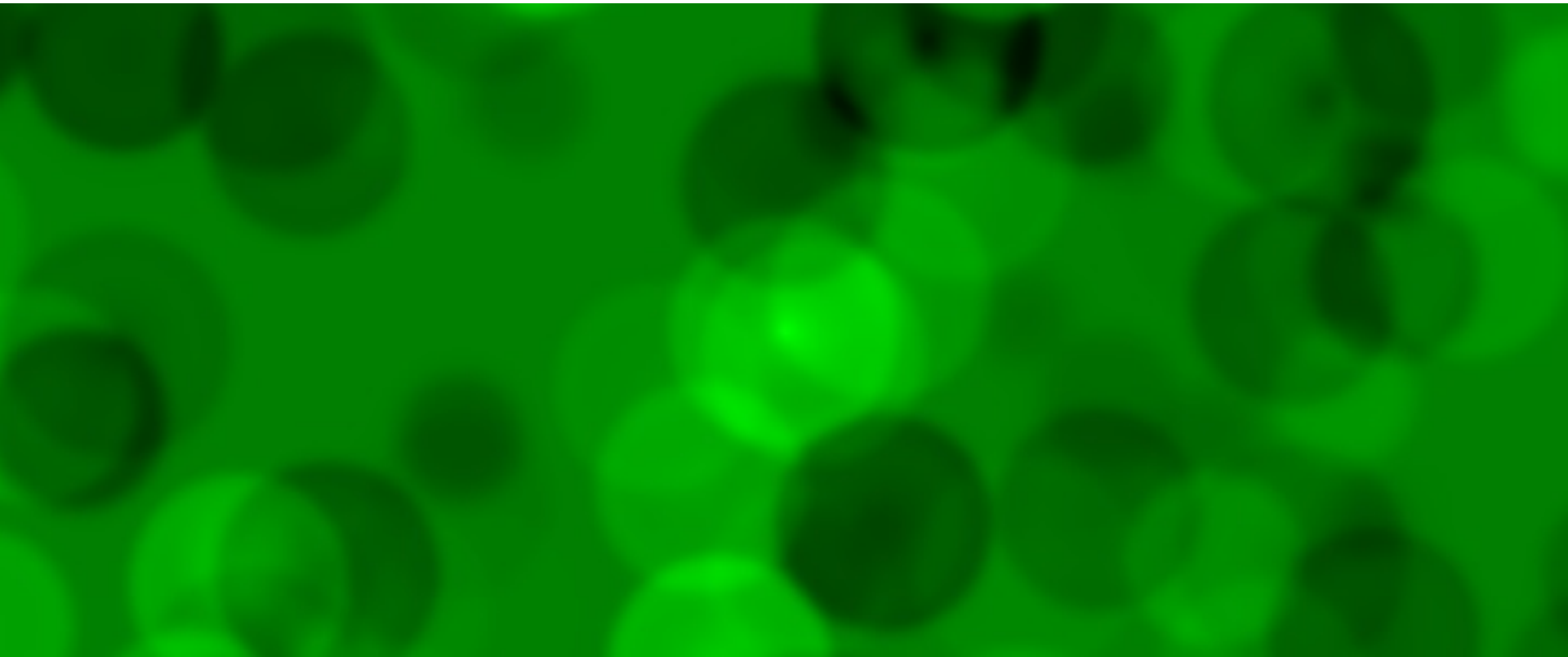
Adding “pebbles”

10



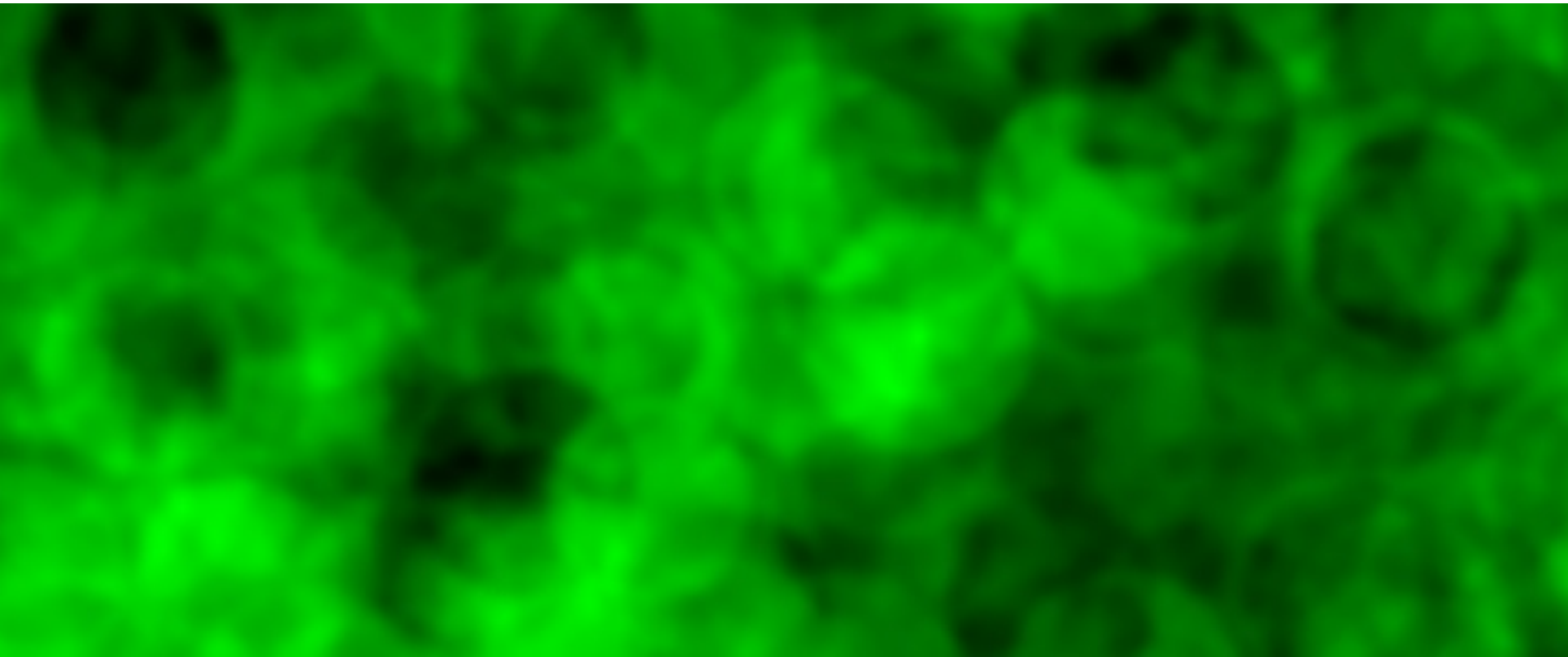
Adding “pebbles”

100



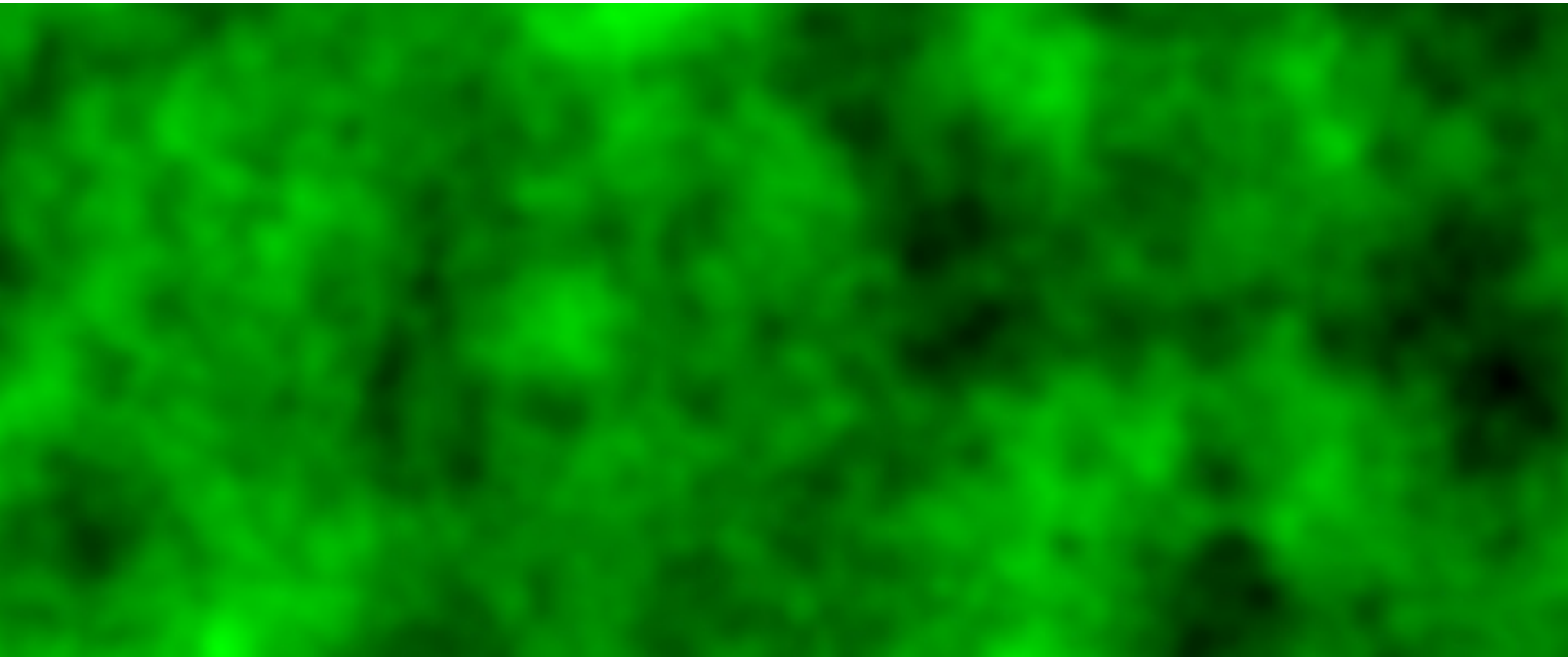
Adding “pebbles”

1000



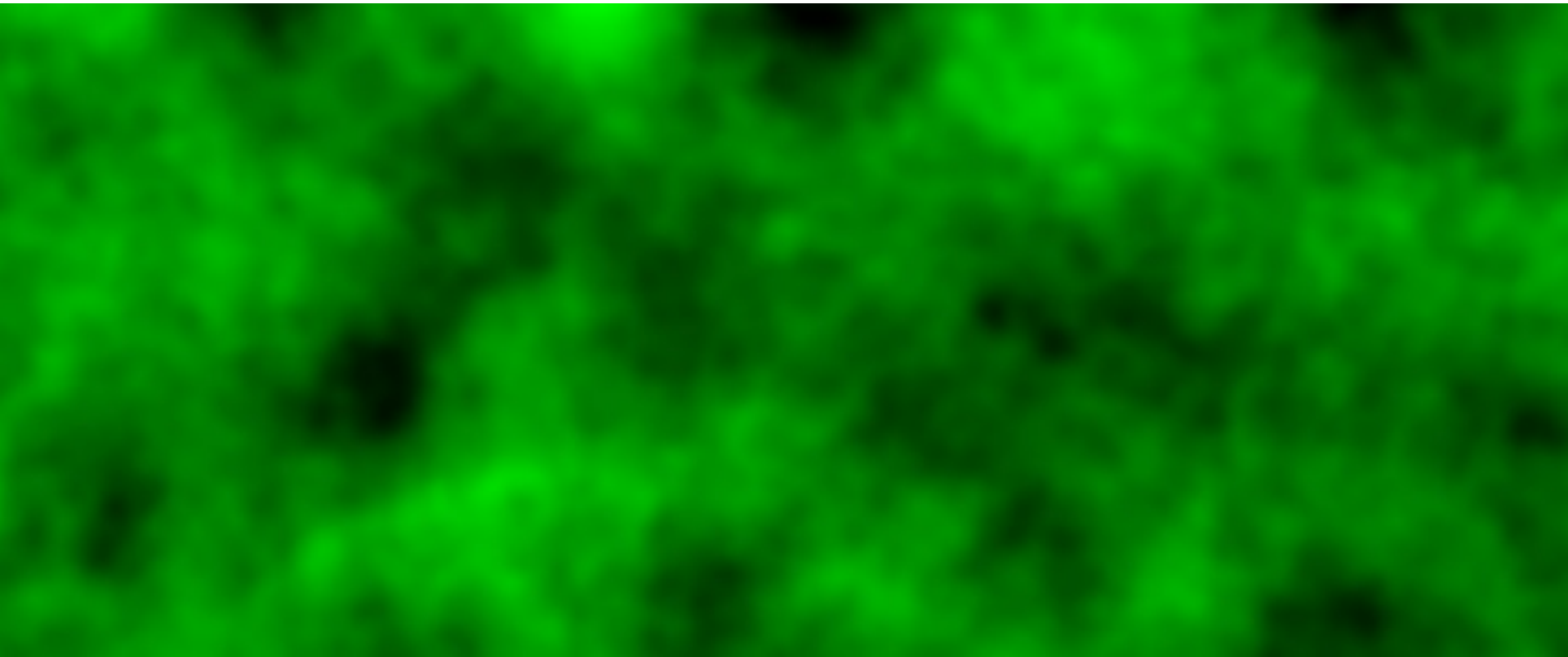
Adding “pebbles”

10,000



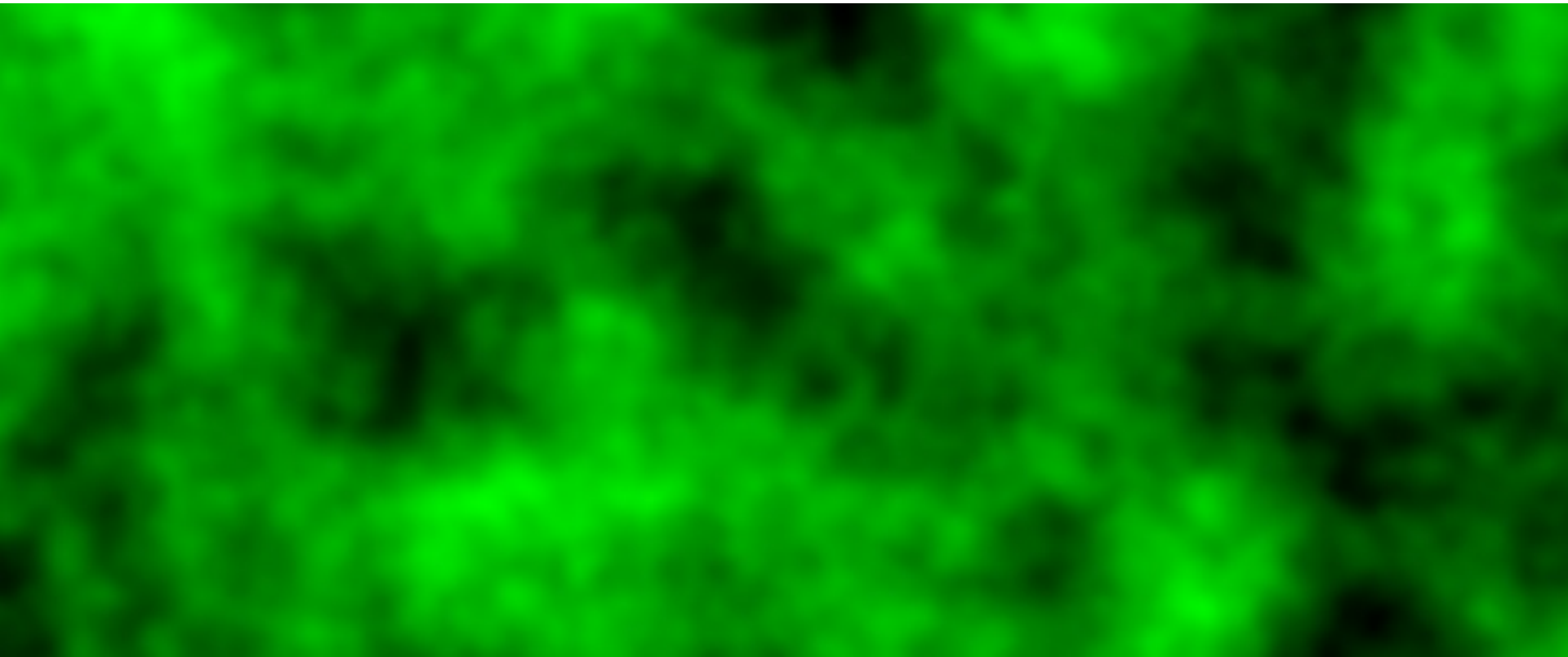
Adding “pebbles”

100,000



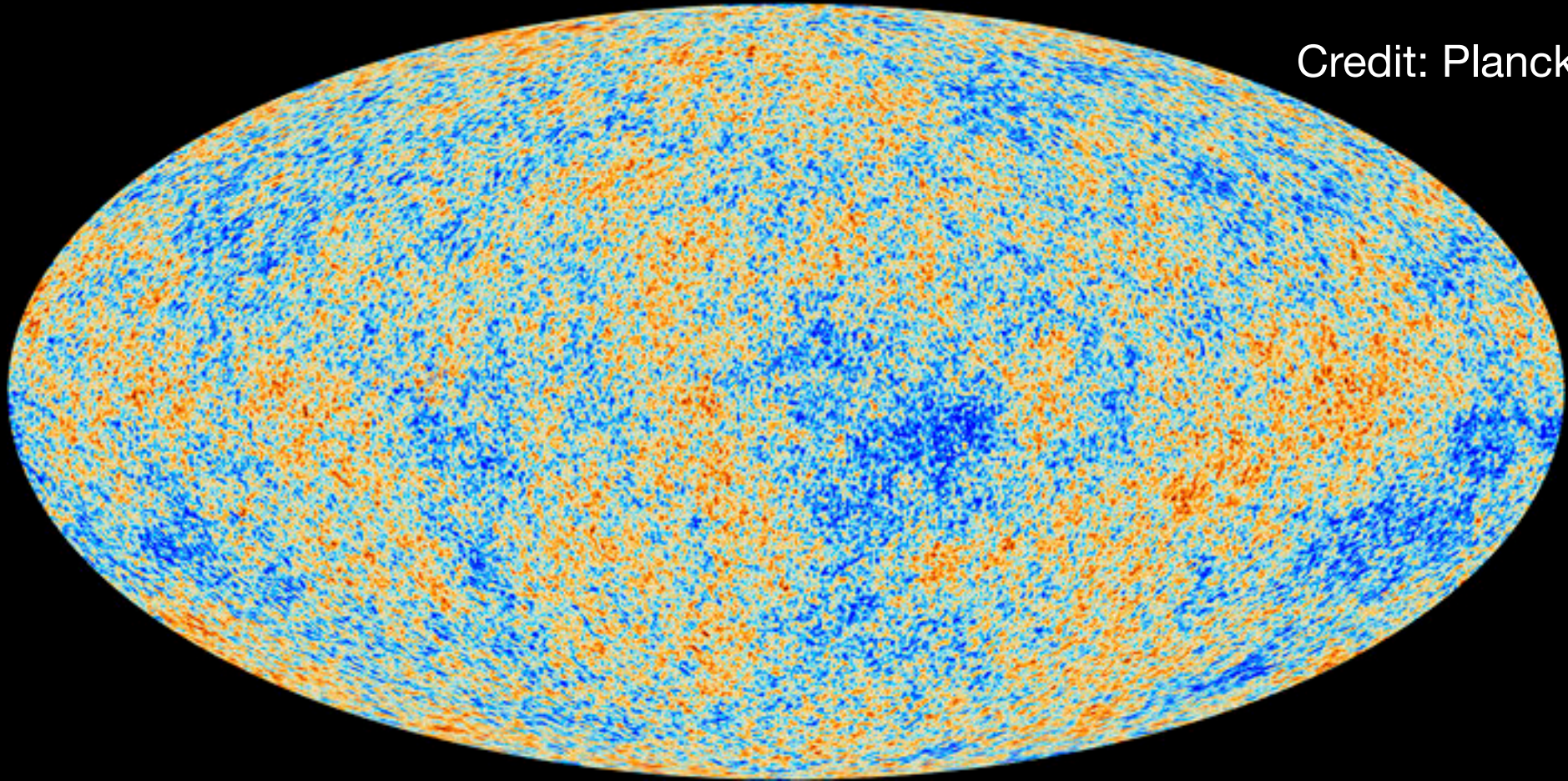
Adding “pebbles”

490,000



We know how sound works!

Credit: Planck



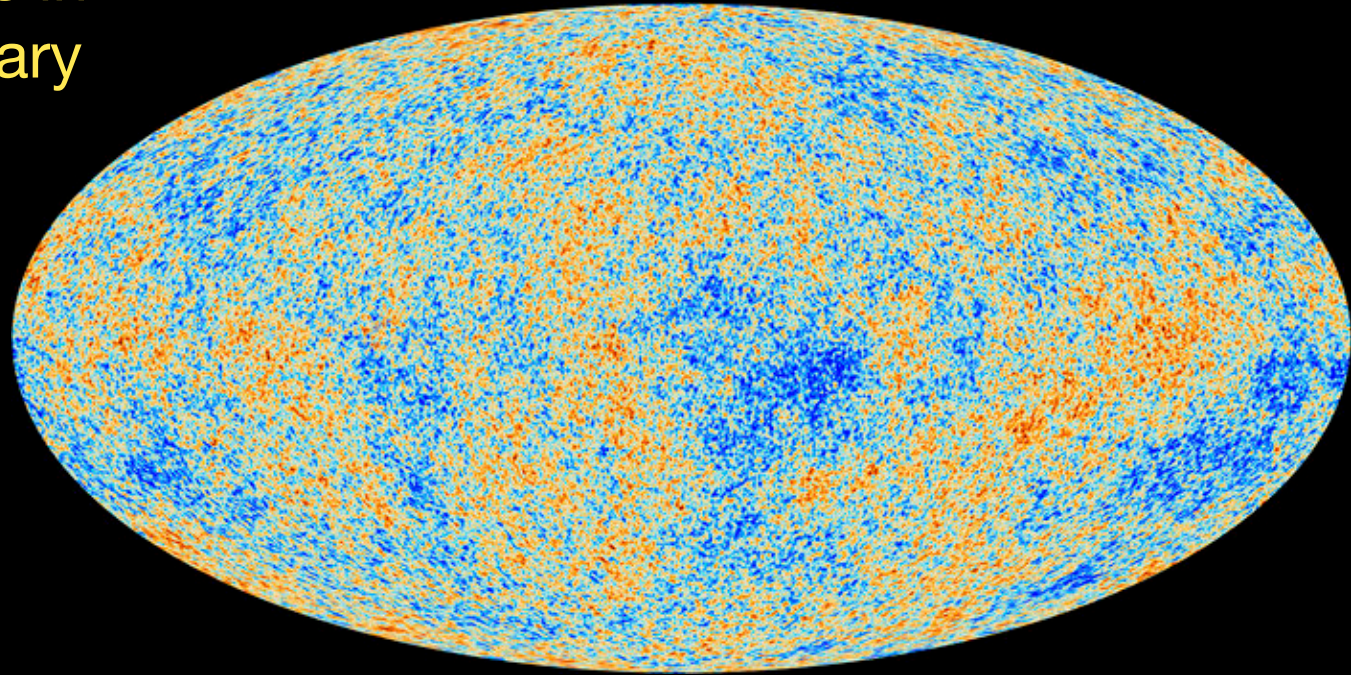
Thus, we can use the patterns to learn about:

what is the Universe made of?

what are the 'pebbles' - what caused the Big Bang? What is dark matter?

Follow up: The Smoothness problem

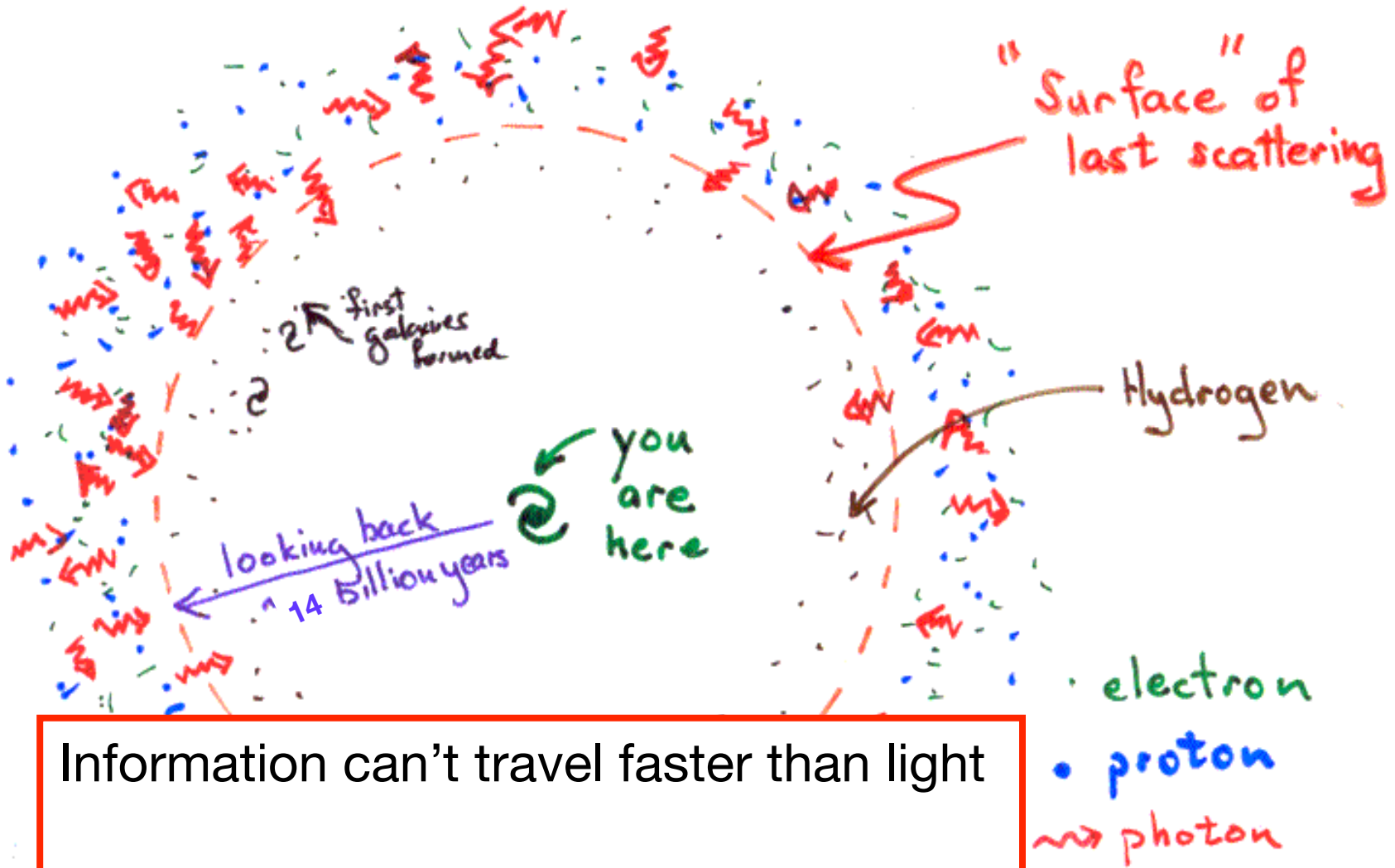
The temperatures in
this room likely vary
by 1°C
ie 1 part in a 300



How can you do
better across the
entire Universe?!?

Observed temperature
variations in CMB:
1 part in 100,000

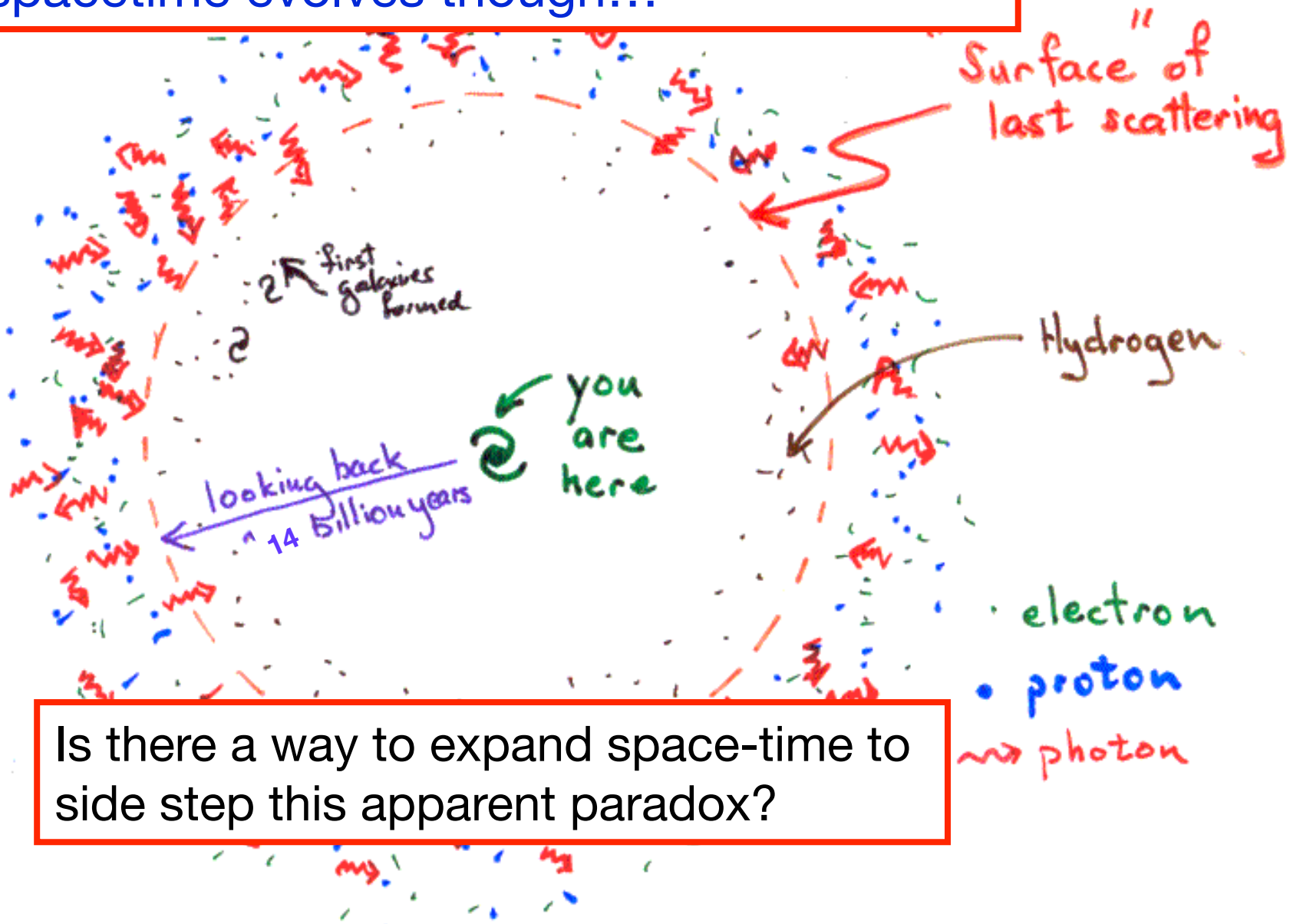
Is special relativity wrong???



Information can't travel faster than light

Light hasn't had enough time to get from 1 side to the other...

Special relativity says nothing about how spacetime evolves though...



Is there a way to expand space-time to side step this apparent paradox?

Can we see Inflation?

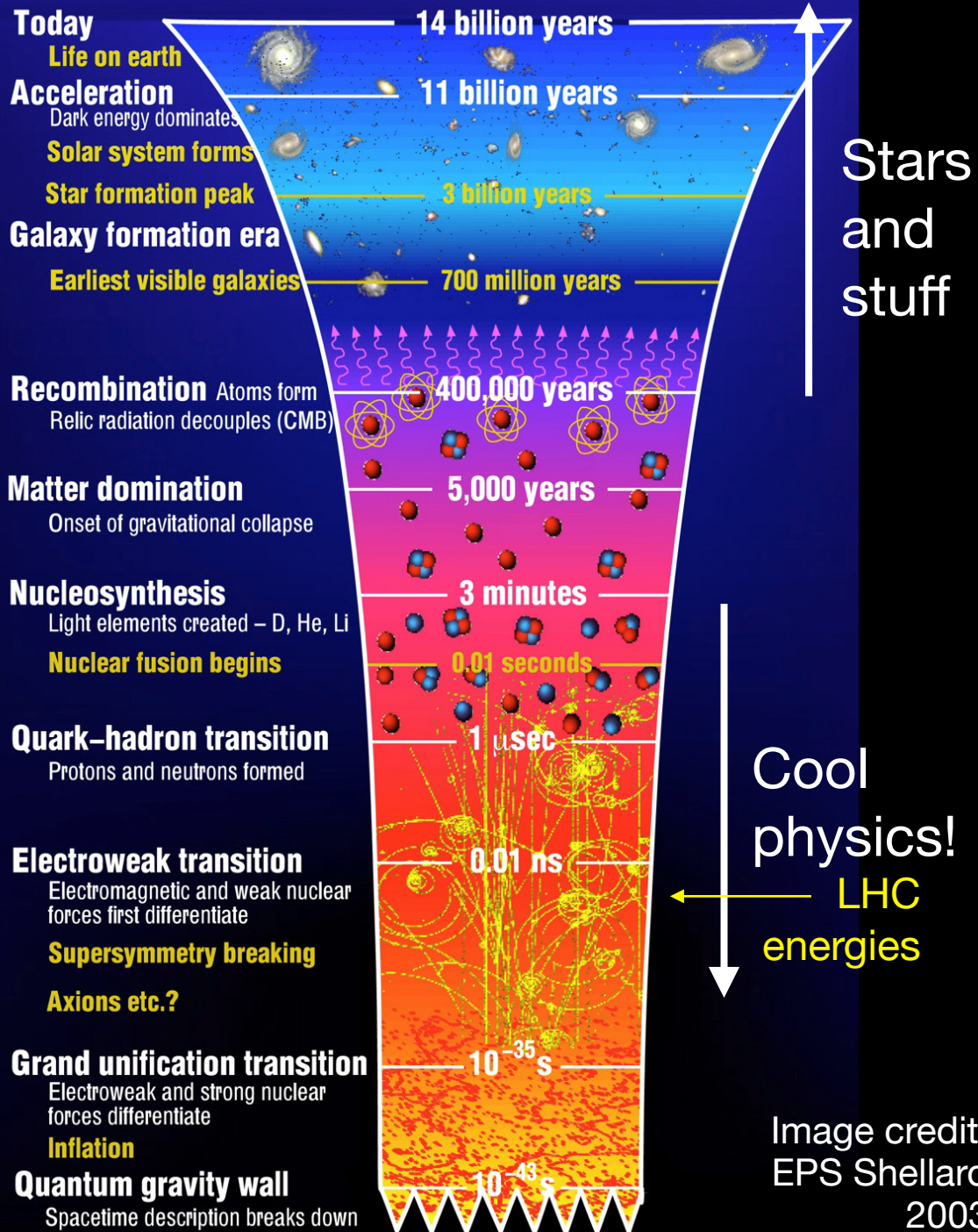
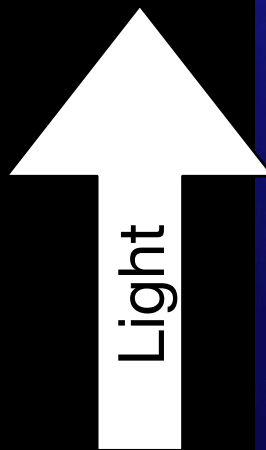


Image credit:
EPS Shellard
2003

~400,000 years
post Big Bang



Early Universe
is opaque to
light

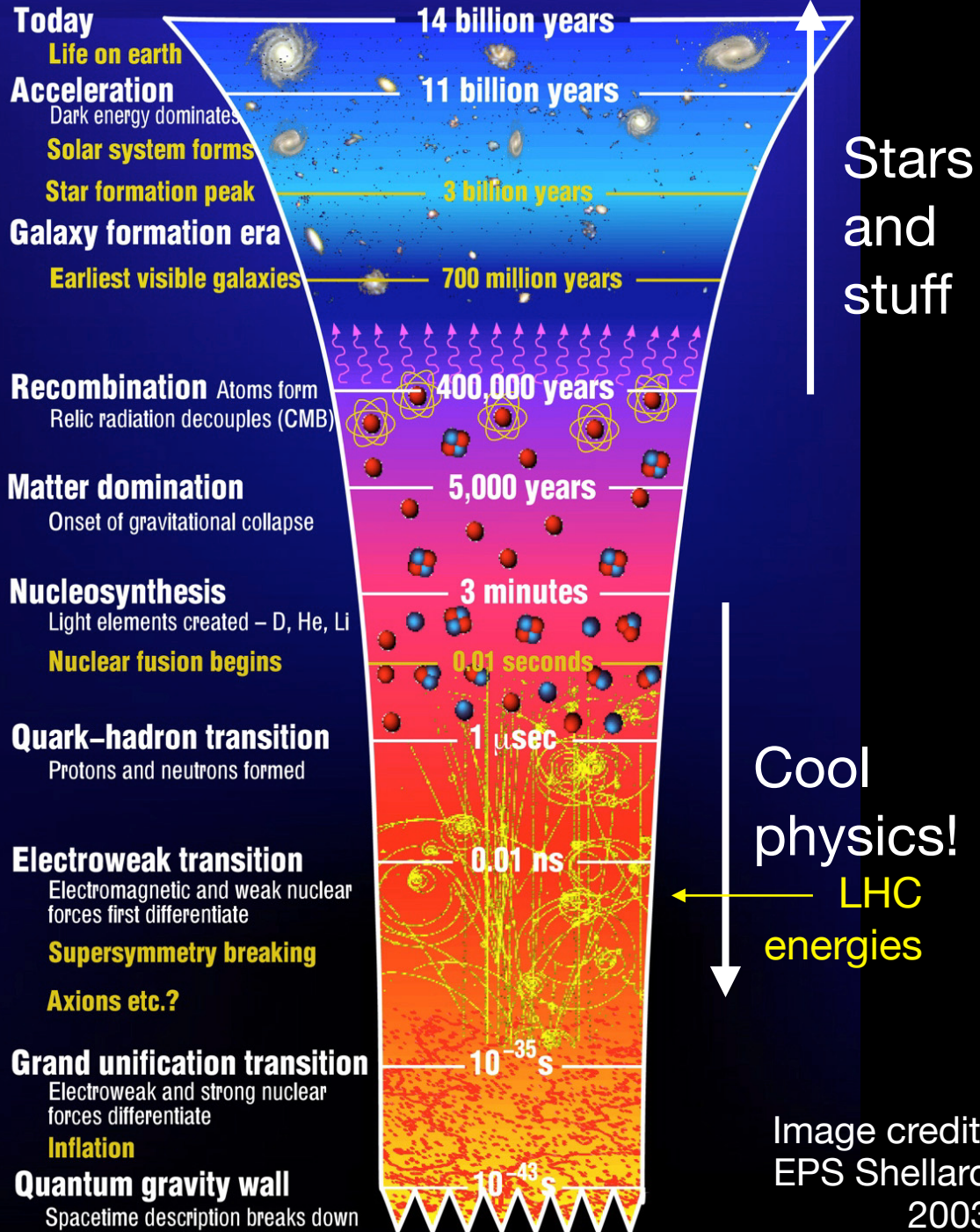


Image credit:
EPS Shellard
2003

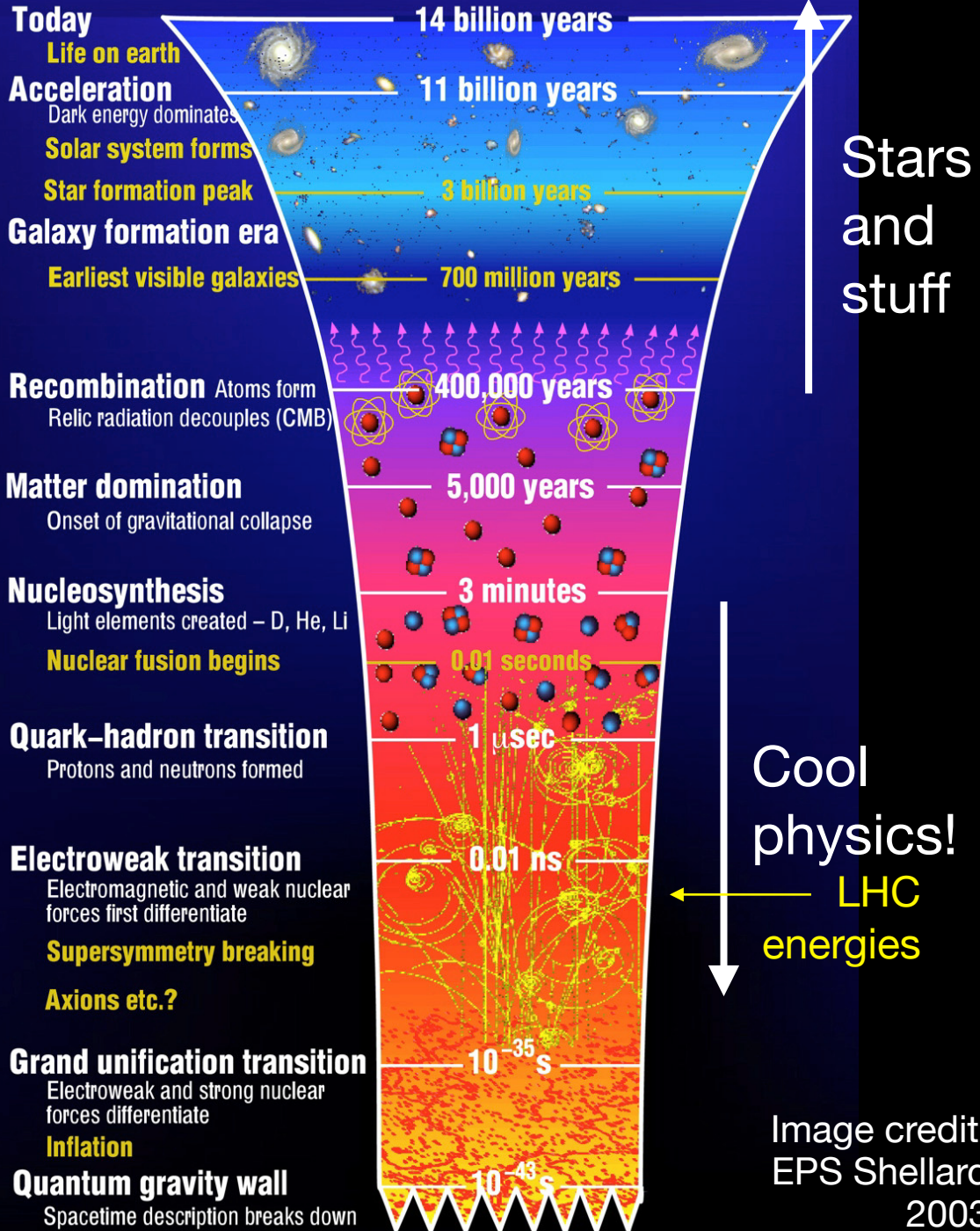
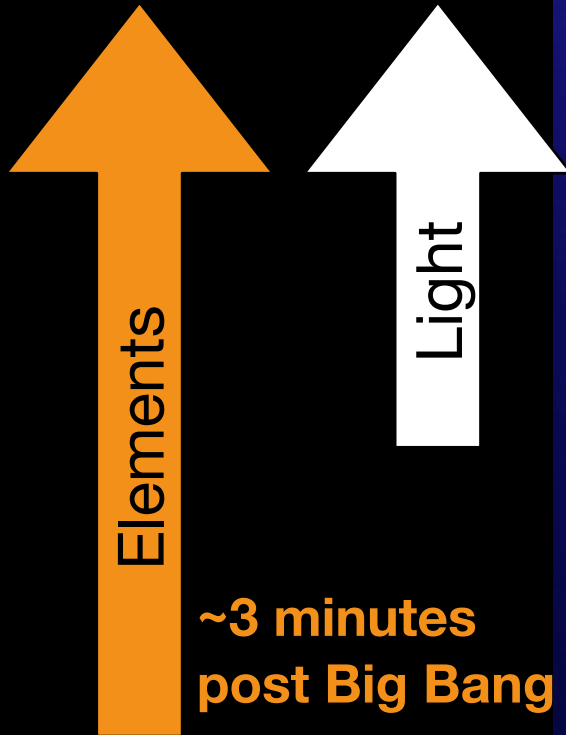
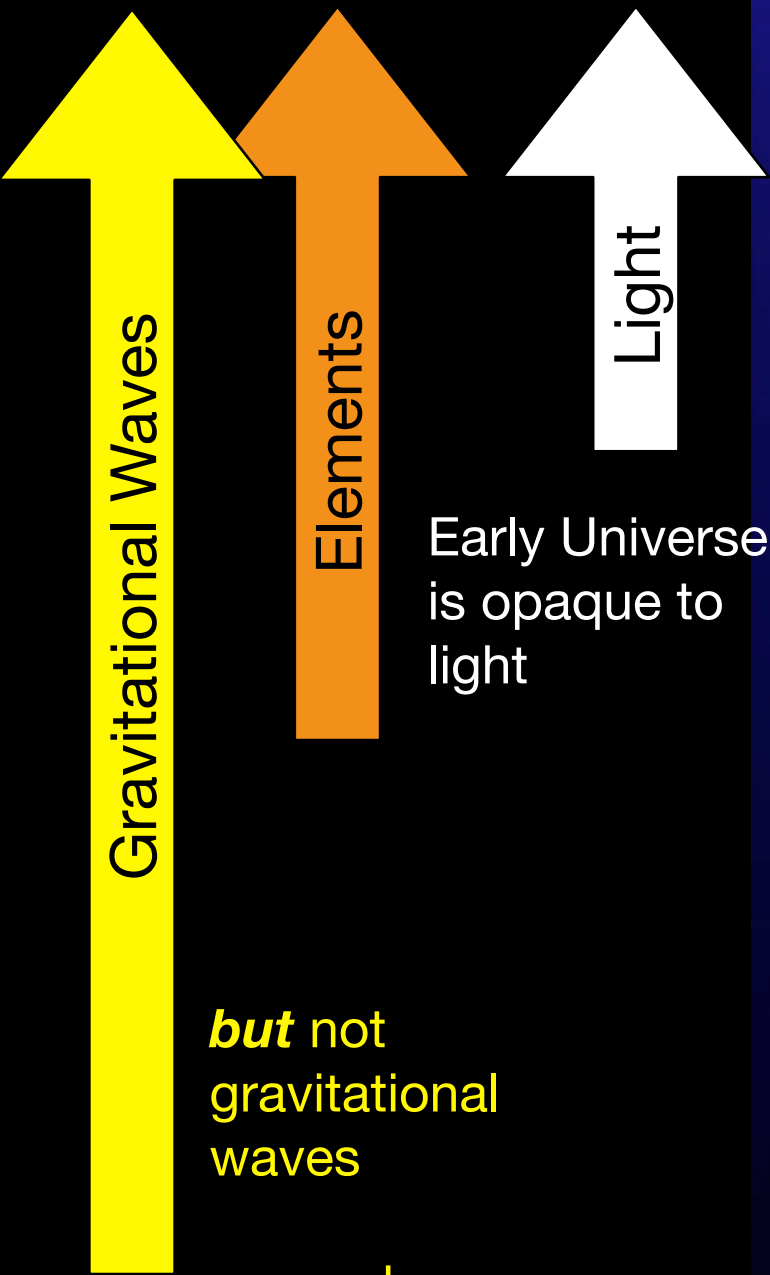


Image credit:
EPS Shellard
2003



Early Universe is opaque to light

but not gravitational waves
so maybe

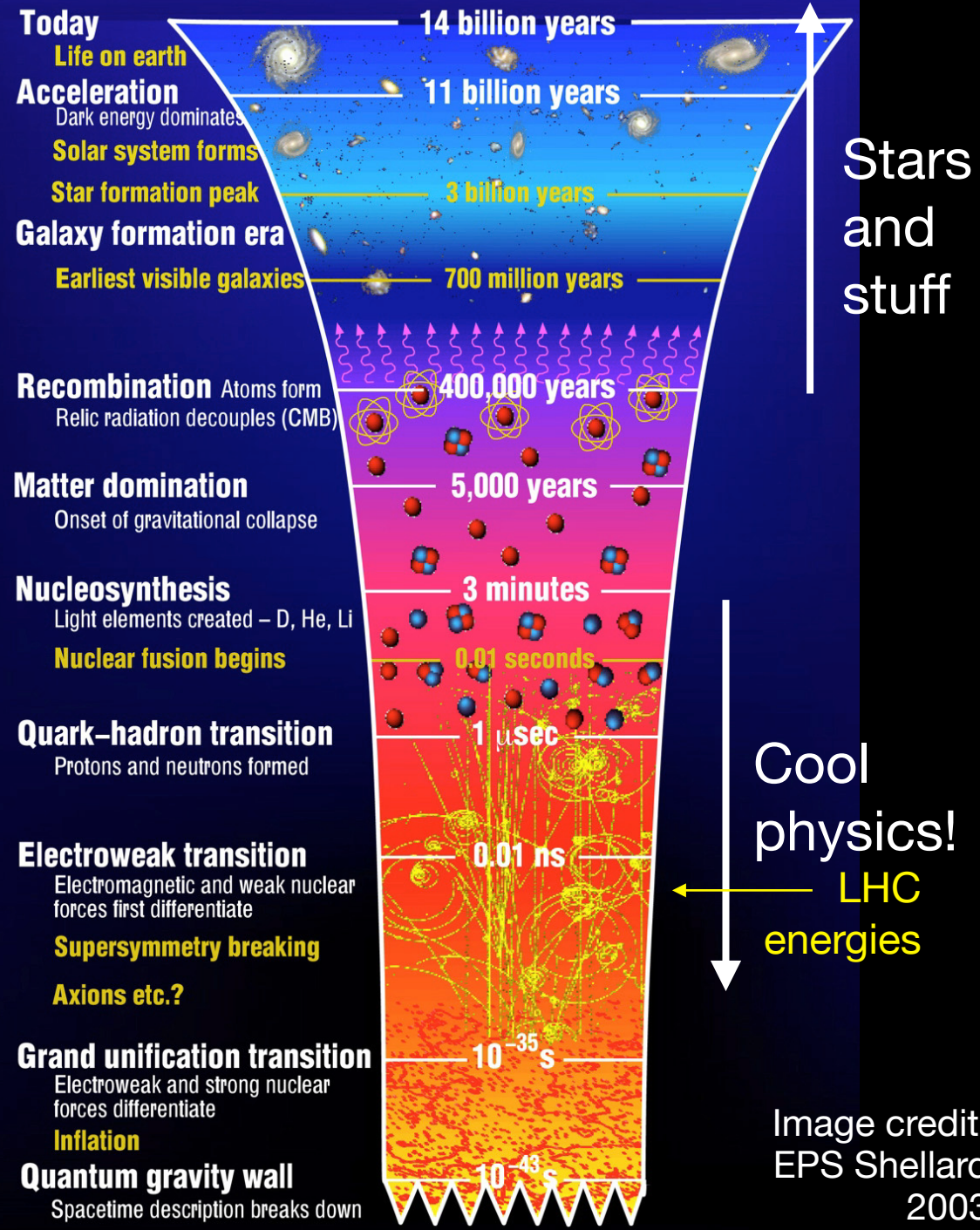
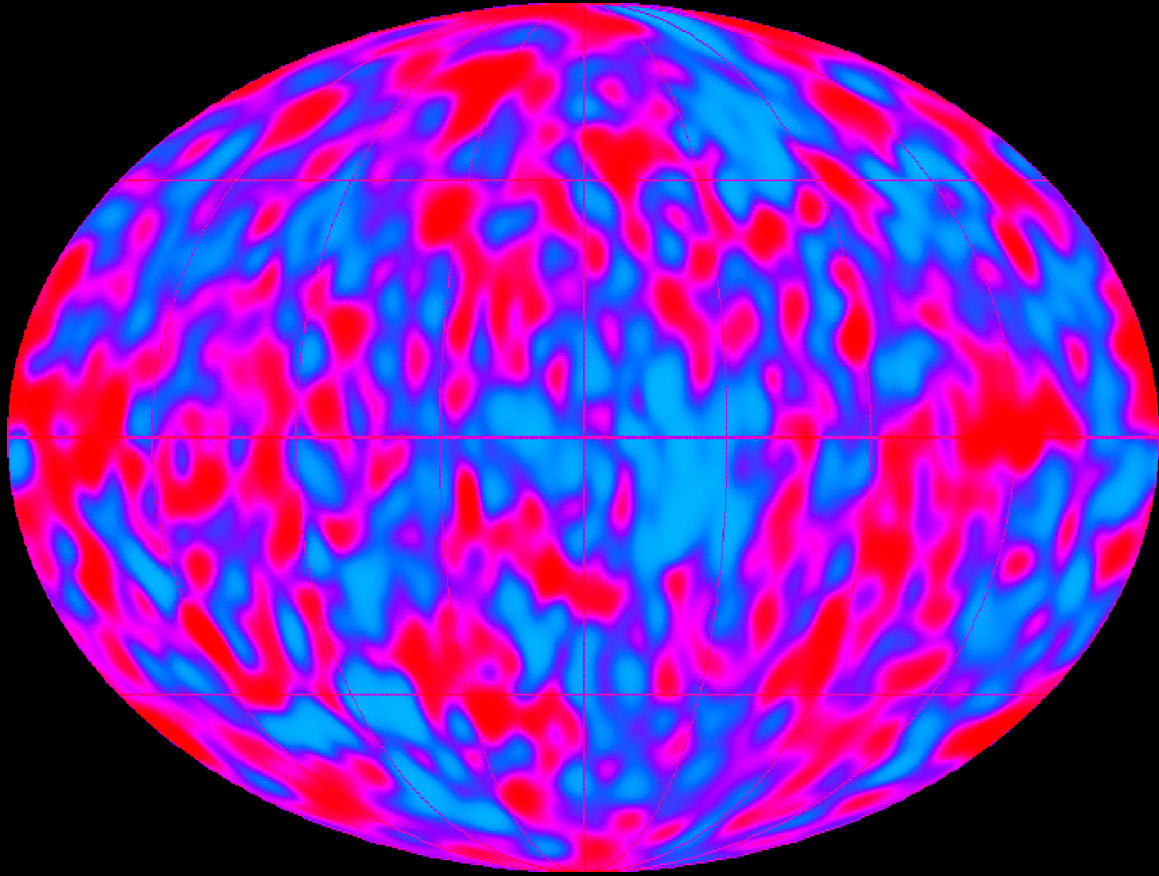


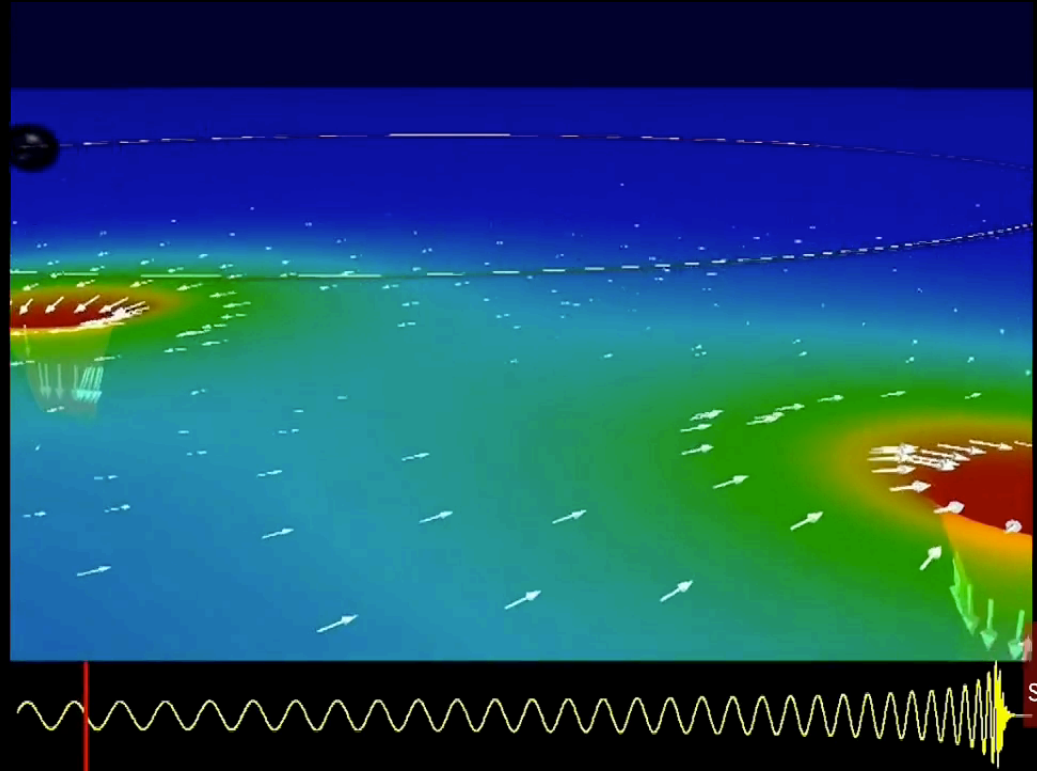
Image credit:
EPS Shellard
2003

Will there be gravitational waves?



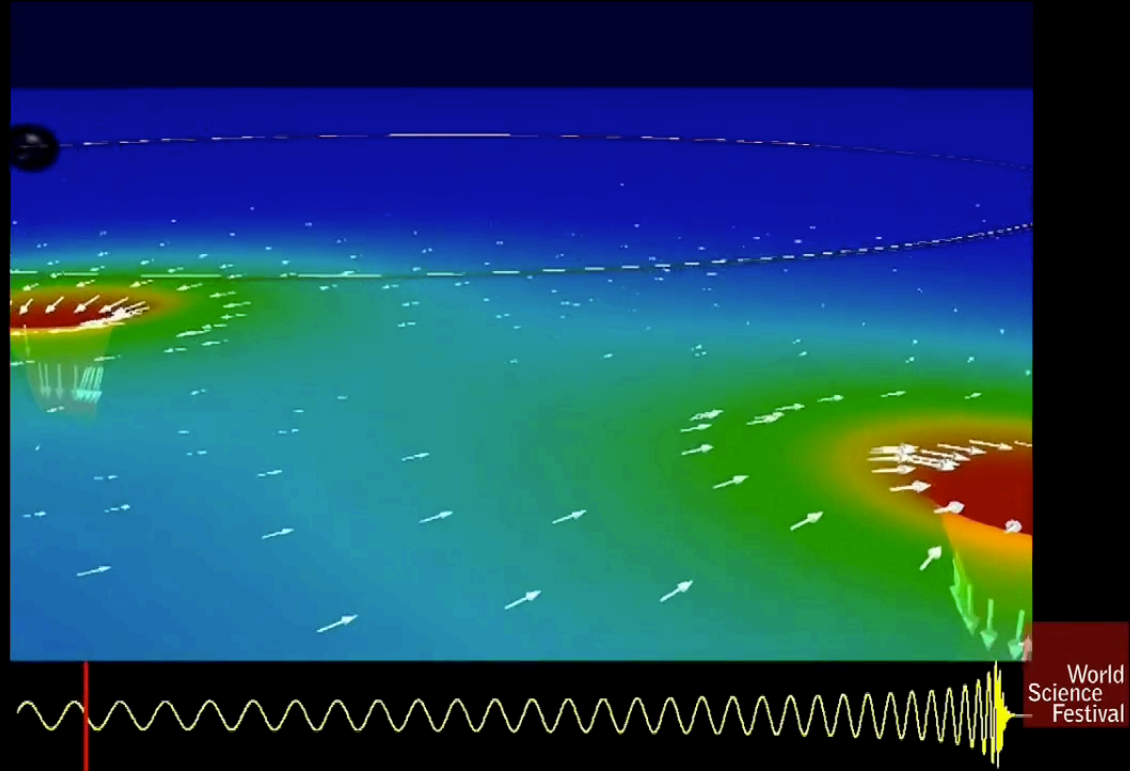
Will there be gravitational waves?

- G. Waves come from accelerating masses



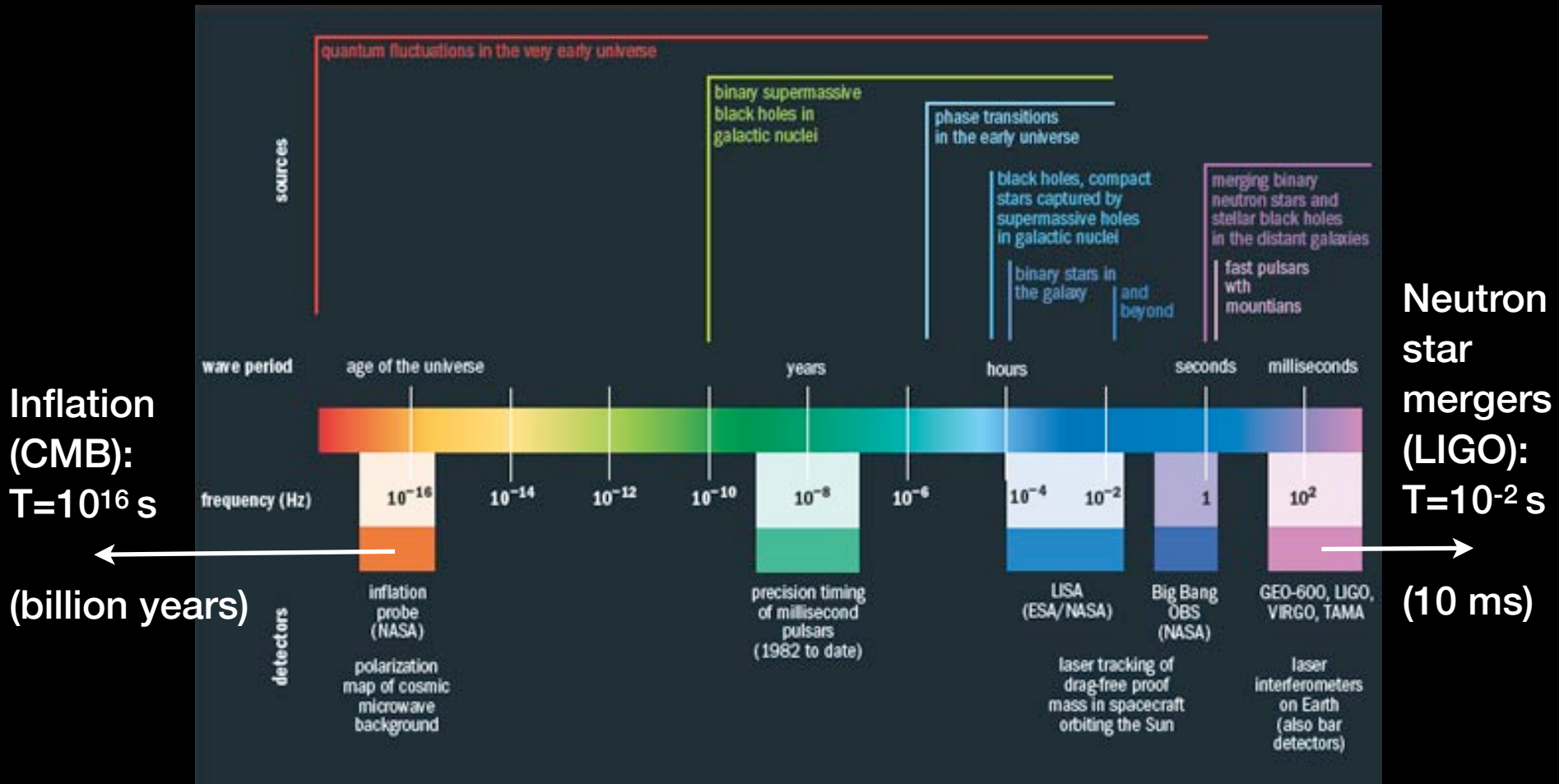
Will there be gravitational waves?

- G. Waves come from accelerating masses
- Plenty of acceleration...
- Almost perfectly homogenous ...
 - But still quantum fluctuations

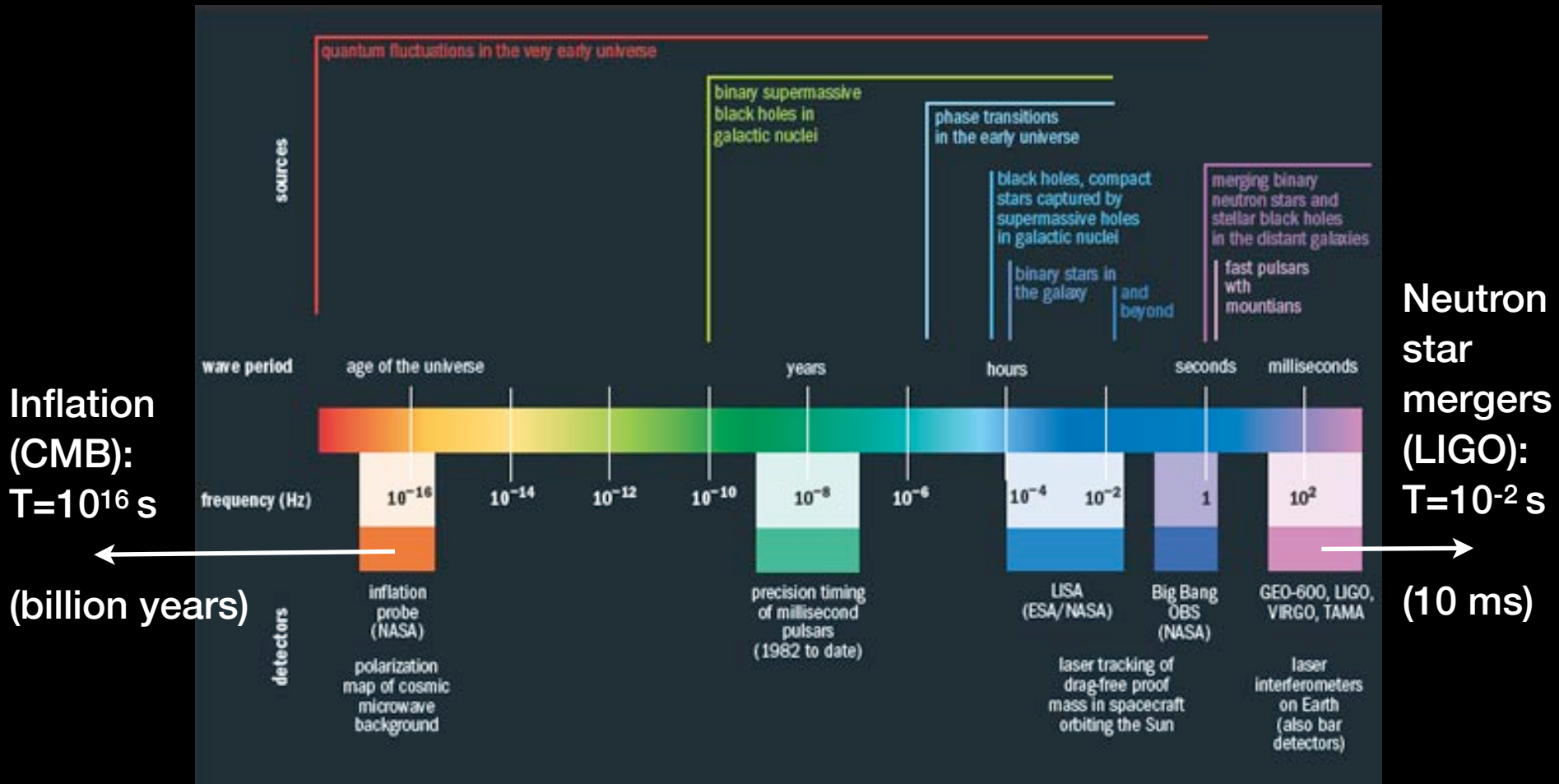


So yes

Wavelengths comparable to the size of the Universe

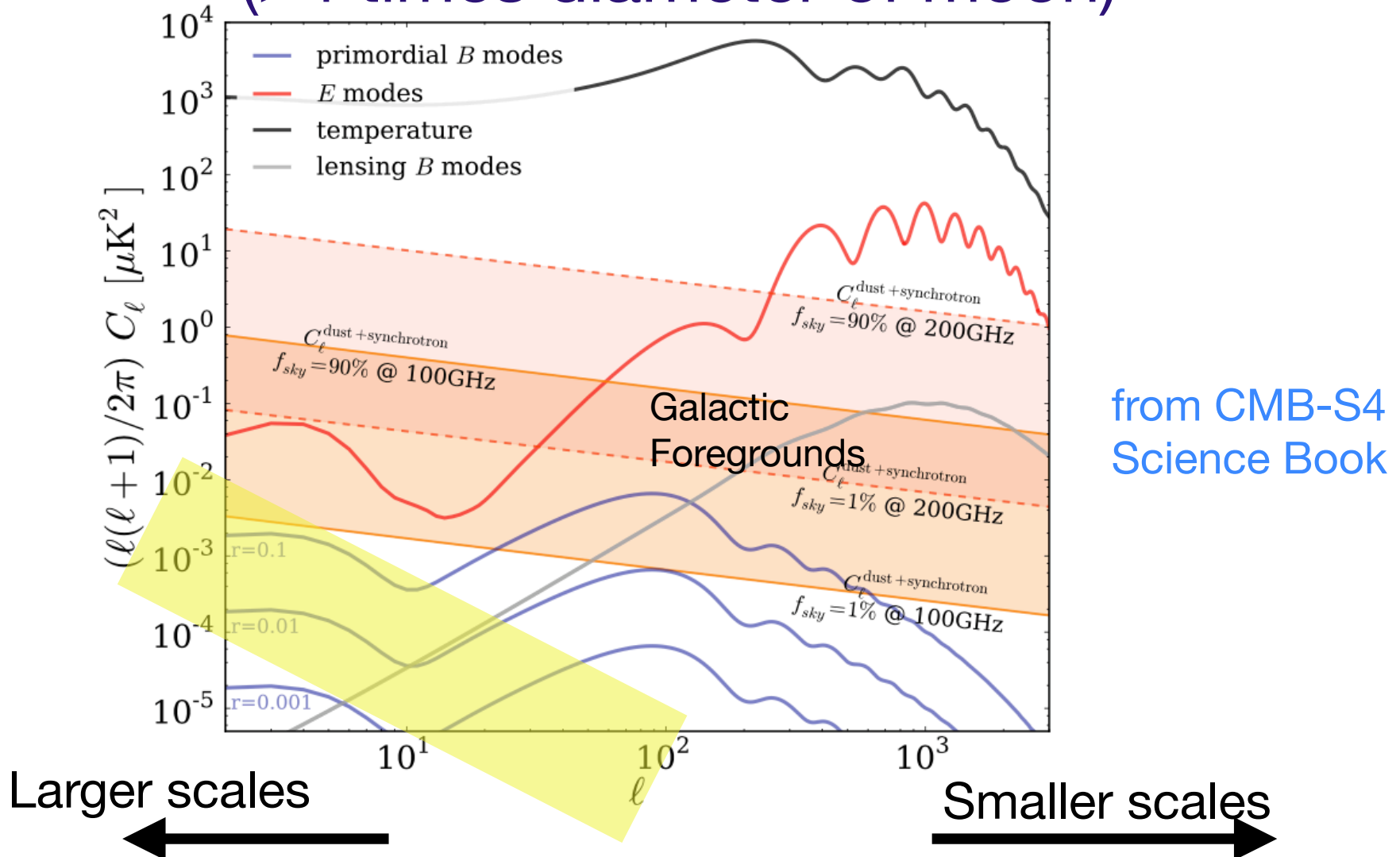


A million trillion times larger than LIGO



We can't build a big enough tunnel...

G. Waves show up as polarized power at large scales (>4 times diameter of moon)



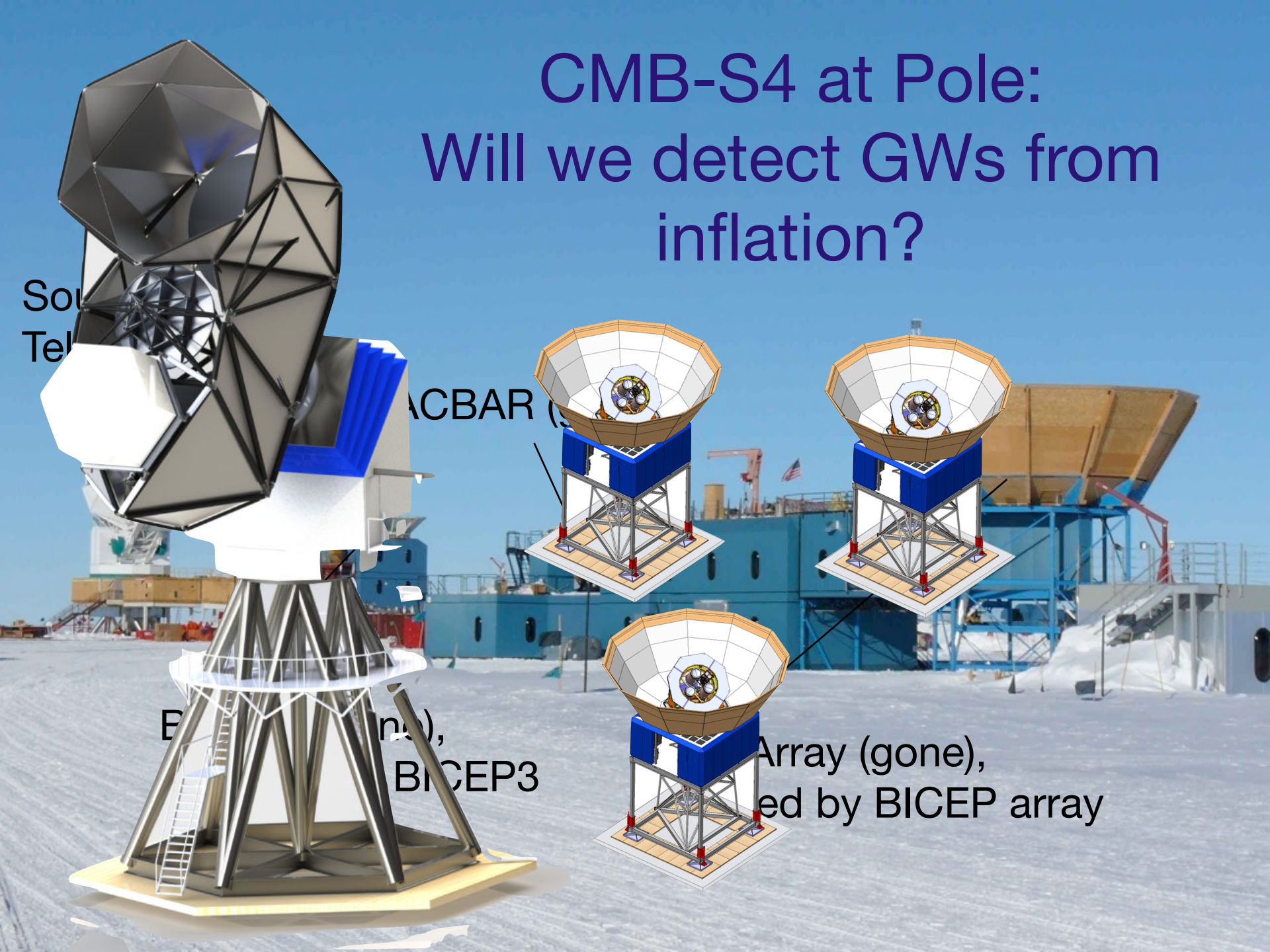
CMB-S4 at Pole: Will we detect GWs from inflation?

Sol
Tel

ACBAR (S)

EPIC (In),
BICEP3

Array (gone),
ed by BICEP array



Searches underway for gravitational waves from inflation

Movie Credit:
Robert
Schwartz

Questions?