#### **Dark Matter-Dark Radiation Interaction**

CMB-S4 Collaboration Meeting, Snowmass Session

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# Phenomenology of dark matter-dark radiation (DR) interaction

#### Dark acoustic oscillation (DAO)

In the early Universe...



# Coupling dark matter to light relativistic species

• Example 1: Dark matter interacting with a massless photon.

$$\mathcal{L}_{\rm int} = -(D^{\mu}\chi)^{\dagger} D_{\mu}\chi - m_{\chi}^{2}\chi^{\dagger}\chi, \quad \text{where} \quad D_{\mu} = \partial_{\mu} - ig_{\chi}\tilde{A}_{\mu}.$$

• Example 2: Dark matter interacting with a massless neutrino via a massive mediator.

$$\mathcal{L}_{\text{int}} = -g_{\chi}\phi_{\mu}\bar{\chi}\gamma^{\mu}\chi - \frac{1}{2}g_{\nu}\phi_{\mu}\bar{\nu}_{\text{s}}\gamma^{\mu}\nu_{\text{s}} - \frac{1}{2}m_{\phi}^{2}\phi_{\mu}\phi^{\mu} - \frac{1}{2}m_{\chi}\bar{\chi}\chi$$
  
...and many more!

Hofmann et al. 2001; Chen et al. 2001; Bœhm et al. 2002; Green et al. 2004; Bertschinger 2006; Bringmann & Hofmann 2007; van den Aarssen et al. 2012 and many more.

DR

## Impact on CMB power spectra: TT



Cyr-Racine et al. (2013)

## Impact on CMB power spectra: EE



Cyr-Racine et al. (2013)

## Sensitivity of Planck data to DM-DR interaction



Cyr-Racine et al. (2013), see also Archidiacono et al. (2019)

## Prospects for DM-DR interaction with CMB-S4



# Conclusions

- DM-DR collisions at early times leave distinct imprints on the primary CMB and through CMB lensing.
- CMB already constraints  $f_{int} < 5\%$  for DM-DR interaction of SM electromagnetic strength.
- Future constraints on DM-DR interaction will be lensing dominated.
- CMB-S4 is likely to push the allowed fraction to sub-percent level, or to an actual detection.

# Thank you!