Transients and Time-Domain Source Science with CMB-S4

Nathan Whitehorn

Michigan State University

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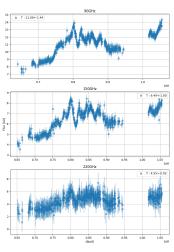




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Spring S4 Meeting - 1

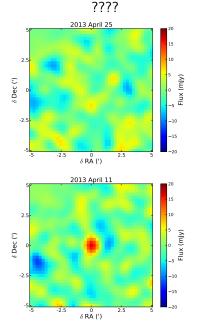
Time Variable Millimeter Sky



AGN

All data from SPT

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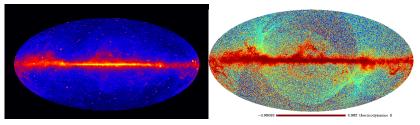


Spring S4 Meeting - 2

Things in the Millimeter Sky

- Tight connections to high-energy astrophysics: millimeter and gamma-ray skies tightly correlated
- Millimeter observations provide a different view of same processes in the same objects as high-energy probes
- Millimeter band a useful probe of otherwise hard-to-see sources – very common for follow-up, but limited fields of view from ALMA make surveys hard
- No wide-area surveys with time-domain capabilities extant or planned from 10 GHz to IR – we have a unique shot at this

Gamma Rays and Millimeter



Which one is Planck and which one Fermi-LAT?

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Capabilities of CMB-S4

Small-area survey:

- Twice daily view of 4% of the sky
- Bands from 20–270 GHz
- Prelim. map noise in central 95–150 GHz of 4 mJy
- Linear polarization sqrt(2) higher

Large-area survey:

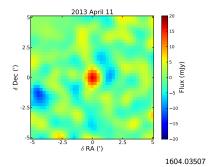
- Twice daily view of 50% of the sky
- Bands from 30–270 GHz
- Prelim. map noise in central 95–150 GHz of 7 mJy
- Linear polarization sqrt(2) higher

(LATs Only: rapidly time-variable sources are, by nature, small)

Science Targets of CMB-S4

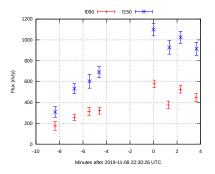
What kinds of things can we see?

- 1. GRB afterglows
- 2. AGN Flaring and Variability
- 3. Multimessenger Astronomy
- 4. Supernovae and TDEs
- 5. Stellar Flares
- 6. New and unexpected things



Things We Know from ACT

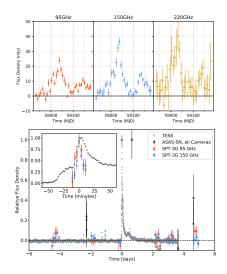
- Three stellar flares
- Last minutes
- Peak fluxes of ~ 1 Jy
- Resolution of minutes from in-observation rescan



ArXiv:2012.14347 Parallel talk by K. Huffenberger

Things We Know from SPT

- Thirteen (minutes-hours) fast stellar flares
- Two slow (weeks) extragalactic events (extreme AGN flares? something else?)
- Real-time program operating



ArXiv:2103.06166 Parallel talk by S. Guns

Unprobed but Interesting: AGN

- Huge number (\gg 1000) of AGN in S4
- CMB instruments continuously monitor all sources in the field
- Large (2x) fluctuations in temperature and *polarization* on day- to month-long timescales
- Same physics as the high-energy emission
- Questions to answer:
 - Cross-correlation with Fermi-LAT: how flares evolve at late times? Emission usually moves to long wavelengths
 - Statistics of fluctuations at long wavelength unknown
 - Major obstacle to multi-messenger correlation analyses no idea what sources do when we aren't looking at longer wavelengths
- Multiple orders of magnitude extension in number of AGN monitored at this granularity



Implications for S4

- Minimum alert rate for ACT/SPT-like sources is 500/year
- Most of these will be new (not triggered by other observatories)
- \blacktriangleright SPT and ACT analyses not yet fully optimized, factor of \sim 2 in sensitivity to S4
- \blacktriangleright Likely S4 detection rate is $\gtrsim 1000$ per year, mostly stellar flares but certainly with a large fraction of other things

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- Need to plan for:
 - Fast cadence
 - Fast response
 - Tiered interestingness of alerts: no one will follow up an alert every 8 hours

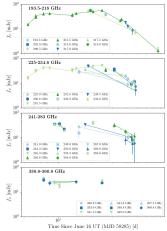
Baseline Design

- Alert pipeline, based on S3 design (baselining SPT, which has real-time alerts already
 - Straightforward application of existing, tested strategy
- Alerts based on line-generated maps, then distributed to the community
 - Maps made at pole for SPLAT, US for Chile
 - Working on distribution strategy and contents of alerts
 - Do we include AGN flares? Which ones?
- Open question: we now know there are *lots* of sub-day sources we could see. Do we need to revisit planning for short-scale sources?

Times central.

- ▶ 14:30: Kevin Huffenberger on ACT Results
- ▶ 14:50: Sam Guns on SPT Results
- ► 16:10: **Tarraneh Eftekhari** on Theory

Tantalizing Hints of New Discoveries



- Extraordinarily bright/weird/crazy supernova-esque object (potentially CCSN)
- Found by chance in optical survey
- Are there more like it? No good way to tell.
- CMB-S4 would see this at 10 - 20σ every day for three weeks anywhere in the survey region

AT2018cow 1810.10880 (Ho+ 2018) Huge discovery potential in time-domain—and it comes almost for free

Summary

- Breaking open a fundamentally new view of the sky, especially with time-domain information
- Strong complementarity to other 2020s planned projects (LSST, LIGO, SKA, upgraded IceCube, etc.) using other wavelengths and messengers
- Similar optimization to N_{eff} —science comes at little cost
- At least daily revisits of the field and large primary mirrors are key—faster is better
- Few glimpses of this unexplored territory—we know there are things there waiting!