

PBDR

Preliminary Baseline Design Report

John Carlstrom & Kevin Huffenberger



Outline

- What is the PBDR?
- Why are we preparing the PBDR?
 - History and impact of reports drafted by the collaboration
 - Purpose of the PBDR
- Timeline
- Status & plans for this workshop

Draft CMB-S4 Preliminary Baseline Design Report

CMB-S4 Collaboration March 8, 2021



What is the PBDR?

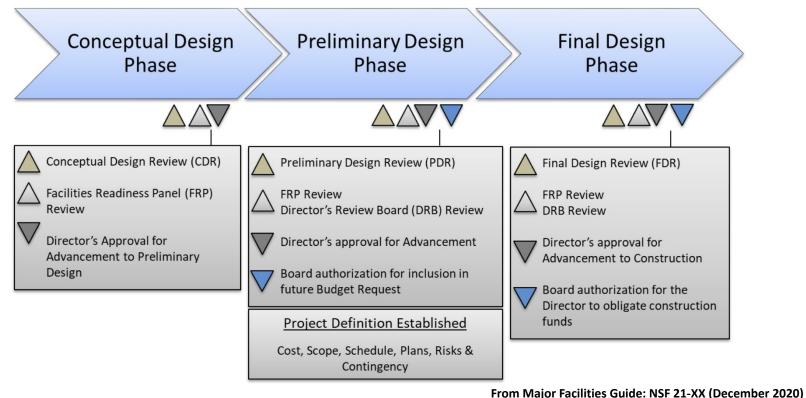
A document that gives the high-level overview of the Science Case, the Technical Design, and the Validation that shows the design will deliver the science.

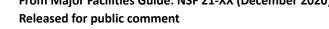
It should be readable by anyone interested in learning about CMB-S4. It will be posted to the public archive (arXiv).

While it gives the overall design, it is not the authoritative reference for design specifications, nor should it include the detailed specifications. These are covered in the various CMB-S4 project documents.

The PBDR should not have to be continually updated as the detailed specifications of the design are worked out. It should remain a viable and useful overall guide of the design as the detailed specifications and interfaces are evolved.

Preliminary Design Phase for an NSF Major Facility





Preliminary Design Phase for an NSF Major Facility

The Preliminary Design Phase further develops concepts to a level of maturity in which there are: a fully elaborated definition of the motivating research questions; a clearly defined site-specific scope; a PEP and an IMP that address major anticipated risks in the completion of design and development activities and in the undertaking of construction; and a realistic (not too low) cost estimate based on known risks that can be presented with reasonable confidence to the NSF.

PEP - Project Execution Plan IMP - NSF Internal Management Plan



Overview of the PBDR

Chapters (coordinators)

- 1. **Science Case** (SC Co-Chairs: Kevin Huffenberger & Joel Meyers)

 Overview of CMB-S4 science themes; background, context and motivation
- 2. **Science and Measurement Requirements** (Charles Lawrence & John Ruhl) Succinct summary of Level 1 science requirements and measurement requirements
- 3. **Preliminary Baseline Design** (TC Co-Chairs J. McMahon & A. Vieregg; L2 science leads) The technical design, organized by WBS elements
- 4. **Science Analysis** (Science Council)

 Overview of the Analysis from project Data Management products to final scientific results
- 5. **Project Overview** (Project Director John Corlett)

 Brief overview of the project structure and plan; does not include detailed schedule and costing

Appendices - Design Validation

- 1. **Technical Design to Measurement Requirements** (Julian Borrill)

 Demonstrate that preliminary baseline design will meet the measurement requirements
- 2. **Measurement Requirements to Science Requirements** (Science Council)

 Demonstrate that measurement requirements will meet the level 1 science requirements



Why are we preparing the PBDR?

- We have benefited greatly by documenting the case for CMB-S4 as it evolves
 - Snowmass papers provided case for input to P5, and led to P5 recommendation
 - Science Book and Instrumentation Books led to joint agency Concept Definition Task Force, (CDT).
 - The acceptance of the CDT report by the AAAS led to the establishment of the Integrated Project Office and the Science Collaboration
 - The Decadal Survey Report (DSR) with its Reference Design was essential to making our case to the public and the Decadal Survey Committee and it was crucial for preparing the many requests for information from the Decadal Survey
 - We now need to prepare the document that presents the <u>Preliminary Baseline Design</u>
- The PBDR is needed for the upcoming NSF PDR and DOE CD-1 reviews.
 It will be particularly helpful, by providing the reviewers with a coherent overview of the entire project. It is a much needed guide to CMB-S4.



Timeline for the PBDR

Now

- Science Case is ready for collaboration reading and comments
- Science Goals and Level 1 Science Requirements ready for reading and comments

At this workshop

- Collaboration update and input to the <u>Measurement Requirements</u> and the validation that the Measurement Requirements are sufficient to meet the Science Requirements
- Collaboration update and input to the technical <u>Preliminary Baseline Design</u>
- Collaboration update and input on the plan to validate the technical design, by showing it will be able to provide the Measurements Requirements
- All major design decisions that cross WBS elements are identified and made, or a path to their resolution is mapped out and scheduled

By end of April, for input to the NSF MSRI-1 review to be held in May

 All sections completed and ready to provide to review panel, except the <u>Technical Design to</u> <u>Measurements Requirements</u> validation section for which we will present its scope and the schedule for its completion.

By end of Summer

 PBDR complete in advance of Director's review and posted prior to NSF PDR and DOE CD-1 (tentatively) planned for end of CY 21

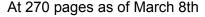


PBDR Status

- Lots of recent progress!
- Some sections much more advanced than others.
- The rest of this presentation briefly reviews the PBDR sections and highlights the workshop sessions dedicated to updating the collaboration and, most importantly, working with the collaboration on the remaining issues of the Preliminary Baseline Design Report. This is the work of the workshop.

Draft CMB-S4 Preliminary Baseline Design Report

CMB-S4 Collaboration March 8, 2021





Chapter 1: Science Case

1	Scien	ce Case (Kevin Huffenberger and Joel Meyers)	1
	1.1	Introduction	1
	1.2	Summary of Science Themes	2
	1.3	Primordial gravitational waves and inflation	10
	1.4	The dark Universe	25
	1.5	Mapping matter in the Cosmos	50
	1.6	The time-variable millimeter-wave sky	65

Status:

- Ready for collaboration reading and comments; required some updating from DSR
- Related workshop sessions:
 - This session! Next slides review the Science Case.
 - Tuesday's plenary/parallel on Science Case Beyond the Design Drivers (Joel Meyers & Kevin Huffenberger)

Chapter 1: Science Case

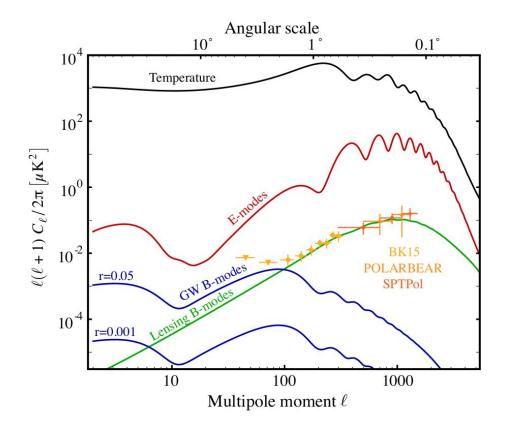
1.2 Summary of Science Themes

Scientific Program of CMB-S4:

- 1. Primordial gravitational waves and inflation
- 2. The dark Universe
- 3. Mapping matter in the cosmos
- 4. The time-variable millimeter-wave sky

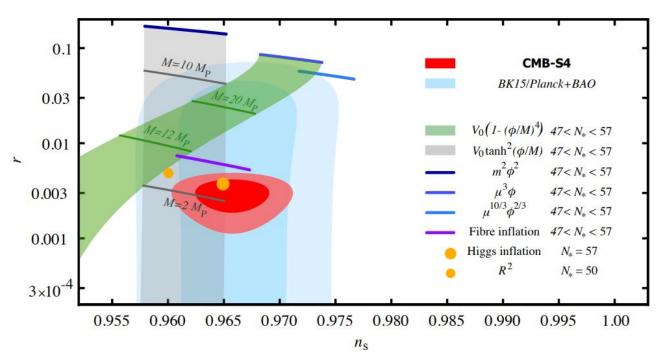
Chapter 1.3: Primordial Gravitational Waves and Inflation

 CMB-S4 will reach the sensitivity necessary to detect gravitational waves from a large class of models: those which naturally explain the deviation from scalar spectral invariance



Chapter 1.3: Primordial Gravitational Waves and Inflation

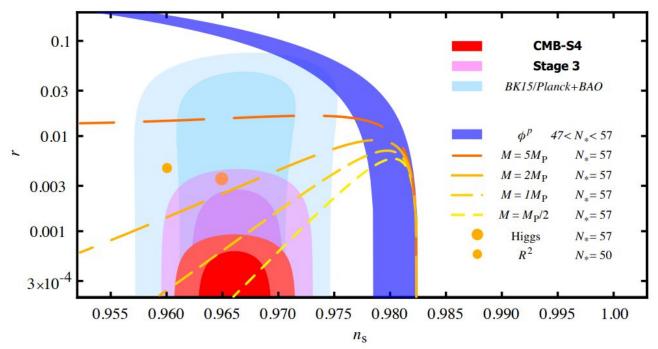
 CMB-S4 will reach the sensitivity necessary to detect gravitational waves from a large class of models: those which naturally explain the deviation from scalar spectral invariance



Forecasted constraints in n_s - r plane assuming r = 0.003, compared to predictions from several inflationary models

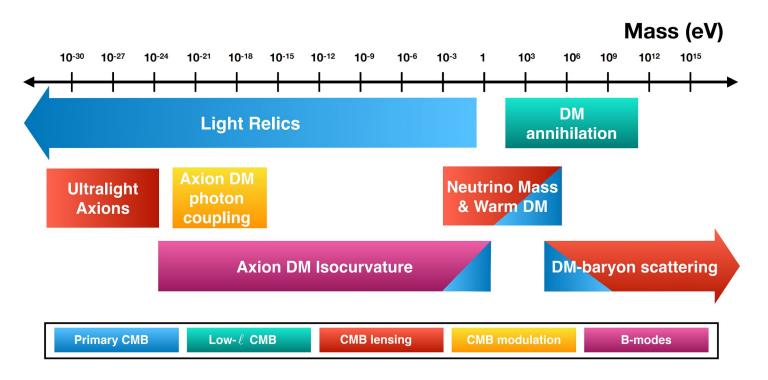
Chapter 1.3: Primordial Gravitational Waves and Inflation

- Even an upper limit on the amplitude of primordial gravitational waves at CMB-S4 sensitivity would also be extremely informative
- A non-detection of primordial gravitational waves would rule out a large class of inflationary models
- Other tests
 (non-gaussianity, etc.)
 of inflation picture



Forecasted constraints in n_s - r plane assuming r = 0

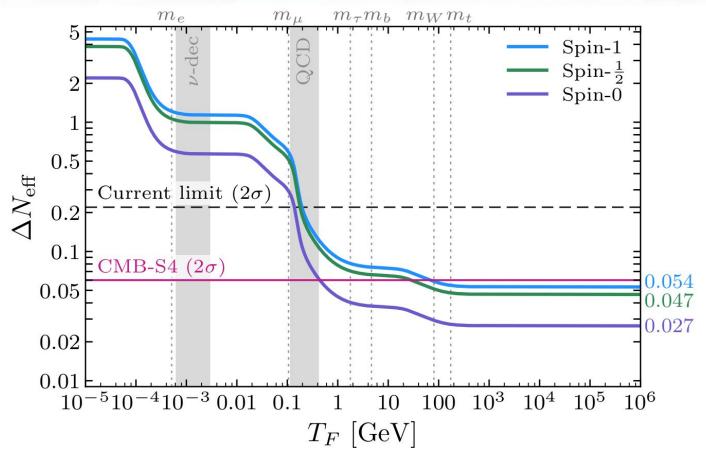
Chapter 1.4: The Dark Universe



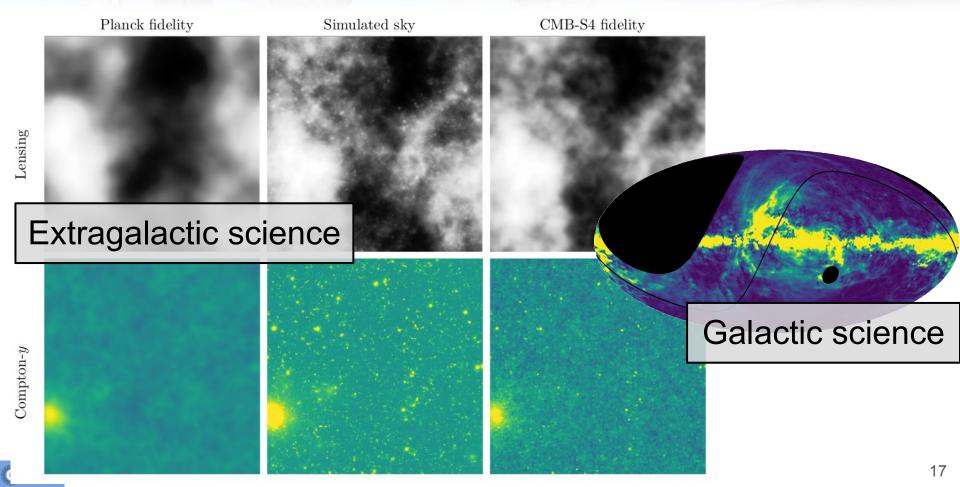
CMB-S4 will probe dark sector physics across a very wide range in mass



Chapter 1.4: The Dark Universe

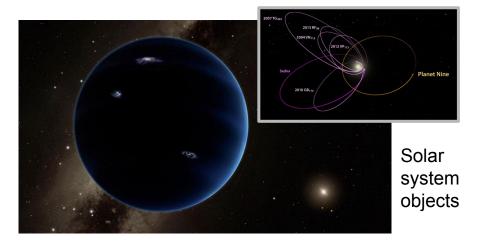


Chapter 1.5: Mapping Matter in the Cosmos

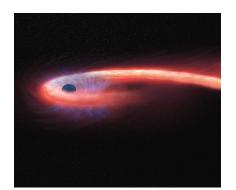


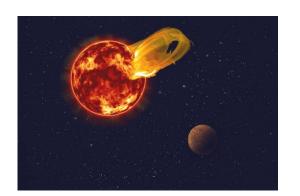
Chapter 1.6: The Time-Variable mm-wave Sky

Gamma ray bursts



Tidal disruption events





Stellar flares

Chapter 2: Science and Measurements Requirements

Four science themes get condensed into four

	science goals, the most important/challenging. These are design drivers.
2	Science and Measurement Requirements (Charles Lawrence and John Ruhl)

				_	
iwrence	and	John	Ruhl)		

	These are de	sign anv	ers.				
Scien	ce and Measurement	Requirements	(Charles	Lawrence	and Je	$ohn \ I$	Ruhl)
2.1	Introduction						

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2.1	Introduction		 		 					ě			• € 5
2.2	Definition of CMB-S4: Science	Goals	 		 		 ٠			٠		•	

2.1 Introduction		 	
2.2 Definition of CMB	-S4: Science Goals	 	

2.2	Definition of CMB-S4: Science Goals	
2.3	Science Requirements	•
2.4	Measurement Requirements	

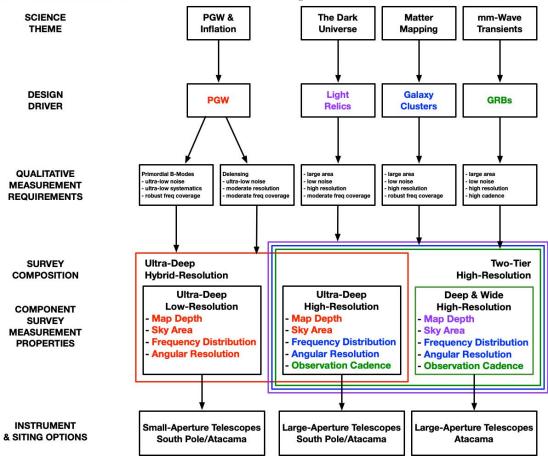
Status:

Flowdown of requirements

Chapter 2.3: Science Requirements

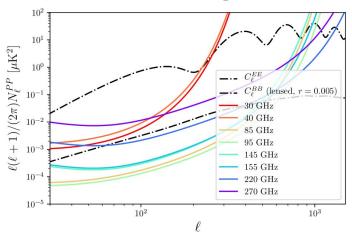
	Formal Science Goal	Science Requirement	Science Theme
PGW	Test models of inflation by measuring or putting upper limits on r , the ratio of tensor fluctuations to scalar fluctuations.	Placing upper limit on r of r \leq 0.001 at 95% confidence if r = 0, or by measuring r at a 5σ level if r > 0.003.	PGW & Inflation
Light Relics	Determine the role of light relic particles in fundamental physics, and in the structure and evolution of the Universe.	ΔNeff ≤ 0.06, 95% confidence	Dark Universe
Galaxy Clusters	Measure the emergence of galaxy clusters as we know them today. Quantify the formation and evolution of the clusters and intracluster medium during this crucial period in galaxy formation.	 (1) Detect at ≥ 5σ all galaxy clusters at z ≥ 1.5 with an integrated Compton Y_{SZ,500} ≥ 10⁻¹² over 50% of the sky. (2) Detect at ≥ 5σ all galaxy clusters at z ≥ 1.5 with an 	Mapping matter
		integrated Compton $Y_{SZ,500} \ge 5$ × 10^{-13} over 2.8% of the sky.	
GRBs	Explore the mm-wave transient sky. Measure the rate of mm-transients for the first time. Use the rate of mm-wave GRBs to constrain GRB mechanisms. Provide mm-wave variability and polarization measurements for stars and AGN.	Detect γ-ray-burst afterglows brighter than 30 mJy at 93 and 145 GHz	Transient sky

Chapter 2.4: Measurement Requirements



Chapter 2.4: Measurement Requirements

Noise curves: e.g. Ultra-deep, low-res



- Measurement requirement noise curves specified Chapter 2.
- Logic of PBDR design validation. Need to show:
 - Design exceeds measurement requirements Chapter 3 and A.1
 - Measurement requirements meet science requirements A.2

Flowdown plenary 1st session Tuesday: science

→ measurement

→ technical

Chapter 3: Preliminary Baseline Design

3	\mathbf{Preli}	minary Baseline Design (Abby Vieregg and Jeff McMahon)	77	
	3.1	Overview (Abby Vieregg and Jeff McMahon)	77	
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	3.6	Small Aperture Telescopes (John Kovac)	124	
	3.7	Data Acquisition & Control (Laura Newburgh)	142	
	3.8	Data Management (Julian Borrill)	150	
	3.9	Sites (Kam Arnold & Amy Bender)	167	

3.1 PBD Overview

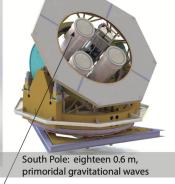
Status:

- Draft is in progress. Being updated from the DSR; defines the overall technical design and will define the surveys.
- Related workshop sessions:
 - Essentially all technical sessions!
- For survey definition:
 - Tuesday morning "flowdown"
 Plenary (J. Ruhl)
 - Thursday Technical to Measurement Plenary/Parallel sessions (Reijo Keskitalo)

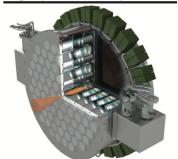
Large Aperture Telescopes



Small Aperture Telescopes

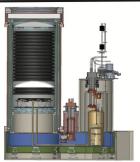


Large Aperture Telescope Receiver



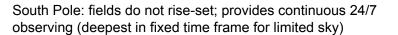
one receiver per telescope each with 85 independent optical paths each feeding one detector array

Small Aperture Telescope Receiver



Three receivers per mount each with one optical path that feeds 14 detector arrays

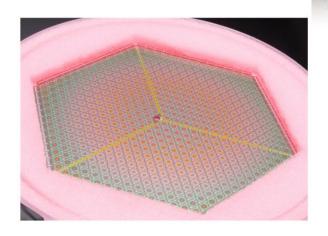
Measurement Reqts. Science Requirements	Ultra-deep low res survey MR 1.1 • Deg res small aperture telescope telescopes for systematics • 6 bands; ≥3% of sky	Ultra-deep high res survey MR 1.2 • arcminute res & overlap w/ ultra-deep low res for delensing • 7 bands; ≥3% of sky	Deep-wide high res survey MR 2.0, 3.1, 3.2, 4.0 Arcminute resolution 6 bands; 70% of sky daily cadence (for 25% of sky)
 SR 1.0 Inflation Set 95% confidence upper limit at r ≤ 0.001; If r > 0.003: measure at equivalent 5σ 	✓	✓	
SR 2.0 Light Relics • ΔNeff ≤ 0.06, 95% confidence			✓
SR 3.1, 3.2 Emergence of Clusters z > 1.5 clusters (wide) z > 1.5 clusters (deep)		(lower mass)	(more sky)
SR 4.0 mm-wave Transients • GRB afterglows > 30 mJy		√ (higher cadence)	(more sky)

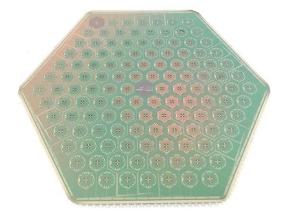


Chile: nearly full sky observable from site; provides deep maps of ~70% of the sky26

Chapter 3.2: Detectors

- Draft is work in progress. The detector PBD is the same as the Reference Design. Most downselects have been made. Several issues remain.
- Related workshop sessions:
 - Wednesday Detector Plenary/Parallel sessions (Brenna Flaugher)





Rhombus and/or Hex pixel layout?

Chapter 3.3: Readout

3.3	Readout	$(Zeesh\ Ahmed)$
	3.3.1	Overview and Scope
	3.3.2	Requirements and Interfaces
	3.3.3	Architecture: Time Division Multiplexing
	3.3.4	100mK Cold Electronics
	3.3.5	4K Cold Electronics
	3.3.6	Production Workflow for Superconducting Circuits on Silicon
	3.3.7	Microfabrication
	3.3.8	Packaging and Cryogenic Screening of Components
	3.3.9	300K Warm Electronics

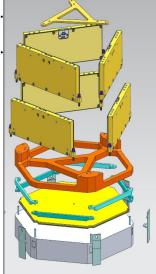
- Draft ready for comments; PBD is advanced version of the DSR Reference Design
- Related workshop session:
 - Tuesday Readout Plenary/Parallel (Zeesh Ahmed)



Chapter 3.4: Module Assembly and Testing

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	3.4.3	Interface Wafers	95
	3.4.4	Module Frame and Assembly	
	3.4.5	Detector Module Testing and Requirements	

- Draft ready for comments.
 Modules PBD is advanced version of the DSR Reference Design
- Related workshop session:
 - Thursday Module Assembly & Testing Plenary/Parallel (Brad Benson)





Chapter 3.5: Large Aperture Telescopes

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	3.5.2	South Pole Large Aperture Telescope (SPLAT) (Emerson and Carlstrom) 99
	3.5.3	Chile Large Aperture Telescopes (CHLATs) (Emerson and Niemack) 108
	3.5.4	Large Aperture Telescope Receivers (LATRs) (Benson)
	3.5.5	Large Aperture Telescope Commissioning Receivers (LATCRs) (McMahon) 121
	3.5.6	Calibration Equipment (Nagy and Bender)

- Nearly ready for comments. LAT PBD has changed significantly from DSR Reference Design; new SP-LAT, LATR, LATCR and I&C equipment added
- Related workshop session:
 - Wednesday Large Aperture Telescopes Plenary/Parallel (Mike Niemack)



Chapter 3.6: Small Aperture Telescopes

3.6	Small A	Aperture Telescopes (John Kovac)
	3.6.1	Small Aperture Telescopes Overview and Design Drivers (Kovac)
	3.6.2	Optics (Paine, Vieregg, Grimes)
	3.6.3	Cryostat (Kusaka, Saba)
	3.6.4	Mount (Pryke)
	3.6.5	Ground Shields and Exterior Baffles (Schmitt)
	3.6.6	Calibration (<i>Karkare</i>)

- Ready for comments. SAT PBD basically an advanced version of the DSR Reference Design; I&C equipment added.
- Related workshop session:
 - Tuesday Small Aperture Telescopes Plenary/Parallel (John Kovac)



Chapter 3.7: Data Acquisition and Control

3.7	Data A	equisition & Control (Laura Newburgh)
	3.7.1	Observatory Control System
	3.7.2	Data Acquisition
	3.7.3	Monitoring and Alarms

- Draft ready for comments; DAQ PBD is advanced version of the DSR Reference Design
- Related workshop session:
 - Wednesday Data Acquisition and Control Plenary/Parallel (Laura Newburgh)



Chapter 3.8: Data Management

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	3.8.6	$\overline{\Gamma}_{\text{ransients}} \ . \ . \ . \ . \ . \ . \ . \ . \ . \$. 161
	3.8.7	Site Hardware	. 164

- Draft ready for comments; DM PBD is expanded and advanced version of the DSR Reference Design; Added DM for transients
- Related workshop session:
 - Tuesday Data Management Plenary/Parallel (Julian Borrill)



Chapter 3.9: Sites

3.9	Sites (Kam Arnold & Amy Bender)					
	3.9.1	Site Requirements & Considerations				
	3.9.2	Chilean Site, Integration and Commissioning (Kam Arnold)				
	3.9.3	South Pole Site, Integration and Commissioning (Amy Bender)				

- Draft in progress. The Sites PBD is similar to the DSR Reference Design. Commissioning and calibration scope needs to be drafted. I&C equipment moved to LATs and SATs
- Related workshop sessions:
 - Thursday Site Infrastructure, Integration & Commissioning Plenary/Parallel sessions (Amy Bender & Kam Arnold)



Chapter 4: Science Analysis

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- Draft in-progress and needs some updates compared to DSR. Chapter reorganized to call
 out the design drivers separately from the complementary parts of the scientific program.
- No dedicated workshop session, but sessions on design validation for science and the complementary science case will address analysis.



Chapter 5: Project Plan

- Compared to the DSR, it will be rewritten to reflect the updated project structure. Will not
 include detailed cost and schedule information. Not ready for comment at this time.
- No directly related workshop sessions, other than the previous project update plenary.

Appendices: Design Validation

A	Desig	n Valid	ation	2	01
	A.1	Technic	al Design to Measurement Requirements (Borrill)	. 2	:01
	A.2	Measure	ement Requirements to Science Requirements	. 2	:01
		A.2.1	Tensor-to-scalar ratio r (Science Requirement 1.0) (Bischoff & Wu)	. 2	:01
		A.2.2	Light relics (Science Requirement 2.0) (Loverde & Reichardt)	. 2	04
		A.2.3	Galaxy clusters (Science Requirements 3.1, 3.2) (Battaglia & Vieira)	. 2	:09
		A.2.4	Gamma-ray-burst transients (Science Requirement 4.0) (Battaglia & Vieira) $\ $. 2	209

Appendices: Design Validation

A.1 <u>Technical to Measurement Requirements</u> Status:

- Not yet drafted. Path to be drafted by April 30; Completed by end of Summer
- Relevant workshop sessions to advance the validation of the <u>T2M</u>:
 - Tuesday morning "flowdown" Plenary (Instrument parameters, J. Ruhl)
 - Tuesday Data Management Plenary/Parallel sessions (J. Borrill)
 - Thursday Technical to Measurement Plenary/Parallel sessions (Reijo Keskitalo)

A.2 Measurement to Science Requirements Status:

- Partly in place; mainly updating from the DSR; Complete by April 30
- Relevant workshop sessions on the validation of the M2S:
 - Wednesday Plenary/parallels:
 Tensor-to-Scalar Ratio (Colin Bischoff & Kimmy Wu)
 Galaxy Clusters (Nicholas Battaglia)
 - Thursday Plenary/parallels:

Light relics (Christian Reichardt & Marilena Loverde) Transients (Joaquin Vieira)



Questions?

Next step for PBDR is to do the work of the workshop.