

### Emmanuel Schaan Chamberlain fellow

# Halo thermodynamics from kSZ with rec. vel.: Progress update

with S Amodeo, A Amon, F Ardila, H Aung, N Battaglia, J deRose, S Ferraro, C Hill, S Huang, J Lange, A Leauthaud, M Madhavacheril, D Nagai, A Roman, A Schneider, K Smith & ACT

### Why care about the gas?



Galaxies surrounded by vast gaseous halos

Gas is more extended than dark matter, due to feedback in galaxy formation → Key astrophysical unknown!

Main limiting theoretical systematic for galaxy lensing

NASA, ESA, DePasquale, Wheatley, Levay

### High S/N with CMB S4 & DESI



and lensing profiles for the same halos!

### Recent measurement: BOSS & ACT+Planck



RA=0

### **BOSS CMASS**

Spectroscopic sample ~400k galaxy groups,  $10^{13}M_{\odot}$ z = 0.4 - 0.7 BAO, Clustering, galaxy-galaxy lensing, CMB lensing



### ACT + Planck

ACT DR5 + Planck,150GHz and 98GHz

Naess+20

ACT DR4 ILC maps Madhavacheril+20, Choi+20, Aiola+20



Image: Debra Kellner

# Measurement: CMASS



# Measurement: LOWZ

### ACT + Planck (microwave)





### Hubble (optical)





# tSZ / kSZ = gas temperature!

$$\frac{\delta T_{\rm kSZ}}{T_{\rm CMB}} = \tau \frac{v_{\rm bulk}}{c} \propto n_e$$
$$\frac{\delta T_{\rm tSZ}}{T_{\rm CMB}} = f(\nu) \ \tau \left(\frac{v_{\rm thermal}}{c}\right)^2 \quad \propto n_e \ T_e$$



Amodeo Battaglia Schaan Ferraro & ACT 20

# Informing hydrodynamical simulations



Amodeo Battaglia Schaan Ferraro & ACT 20

New territory: low halo masses, outside virial radius Data suggests hotter gas in the outskirts Informs subgrid feedback prescriptions in hydro sims

# Constraining galaxy formation: challenges



**Color selection matters** red vs black

**Stellar/halo mass selection matters** red vs purple

**GNFW not expressive enough** solid vs dashed

**2-halo term matters** dashed vs dotted

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### Total mass profile and gas profile

Same halos, HOD, weighting (linear in mass, VS tSZ or Xray), angular scales → no modeling needed

# Is BOSS lensing low?

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Joint analysis of tSZ, gg lensing, RSD for BOSS galaxies Investigate low lensing tension depending on host halo mass



3 Stellar mass bins =  $[10^{11.4} - 10^{11.57}],$  $[10^{11.57} - 10^{11.75}],$  $[10^{11.75} - 10^{13.0}]$ 







## kSZ quadratic estimator



kSZ yields lower noise than galaxy density!

Several groups implementing estimator: Johnson+, Munchmeyer+, Smith+

Alex Roman & Kendrick Smith working on BOSSxACT (5 $\sigma$  currently)

### DESI has started!



Commissioning complete, main survey ongoing until 2025 5k fiber spectrograph on 4m Mayall telescope 35M redshifts over 14k deg<sup>2</sup>

In 2.5 months, DESI gathered as many redshifts as BOSS+eBOSS in 10 years!

# Route towards S4: DESI x ACT

Project idea 1: tSZ from ACT + Legacy Survey LRG, BGS, ELG, QSO high precision tSZ profiles

### Project idea 2: tSZ & kSZ from ACT + DESI LRG, BGS, ELG, QSO

highest precision kSZ profiles best constraints on galaxy formation and baryonic correction to lensing Codes <u>Thumbstack</u>, <u>MopC-GT</u>, <u>Iskay</u> are ready to use!



### **Project idea 3: large-scale velocities & f**<sub>nl</sub> **from kSZ** Smith+18, Münchmeyer+19