

# CMB x LSS Simulations

“Synergies of Large Scale Structure Surveys with CMB-S4”  
CMB-S4 2021 Summer Collaboration Meeting

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with inputs from Marcelo Alvarez, Yuuki Omori, Francois Lanusse

# Why Cross-Correlations?

See all other talks in this session

# Why Simulations?



**Survey pipeline:**

map → statistics → cosmology

**Systematics — instrumentation**

(beam, bandpass, electronics...)

**Astrophysics**

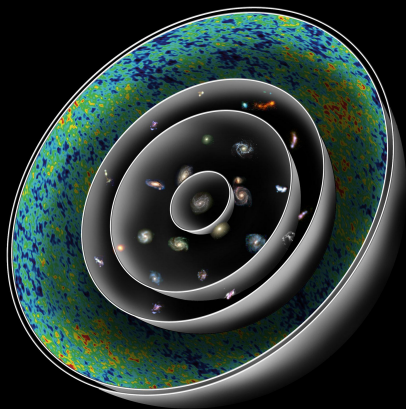
(baryons, intrinsic alignments...)

**Covariance**

**Modeling**

More  
Stringent

# Why Correlated Simulations?



Survey pipeline:

e.g. 6x2 analysis, CMB FGs x LSS

**Systematics** — **instrumentation**

(~~beam, bandpass, electronics...~~)

**Astrophysics**

(baryons, intrinsic alignments...)

**Covariance** between CMB x LSS

**Modeling** beyond the linear theory

More  
Stringent

## Typical CMB simulations

- 2D maps
- Single cosmology
- Gravity-only / approx. methods
- Full-sky / curved
- 1-10 Gpc boxes
- Painted observables

Examples: Sehgal+2010, Stein+2020 (Websky)

## Typical LSS simulations

- 2D (lensing) and 3D (galaxies)
- More cosmologies
- Gravity-only, some hydro
- Often flat-sky
- $\lesssim$ Gpc box
- HOD / Semi-analytic models

Examples: LSST DC2, Euclid Flagship, DES Buzzard, Magneticum, Aemulus, BAHAMAS, SLICS, Quijote, MassiveNuS, and many many more..

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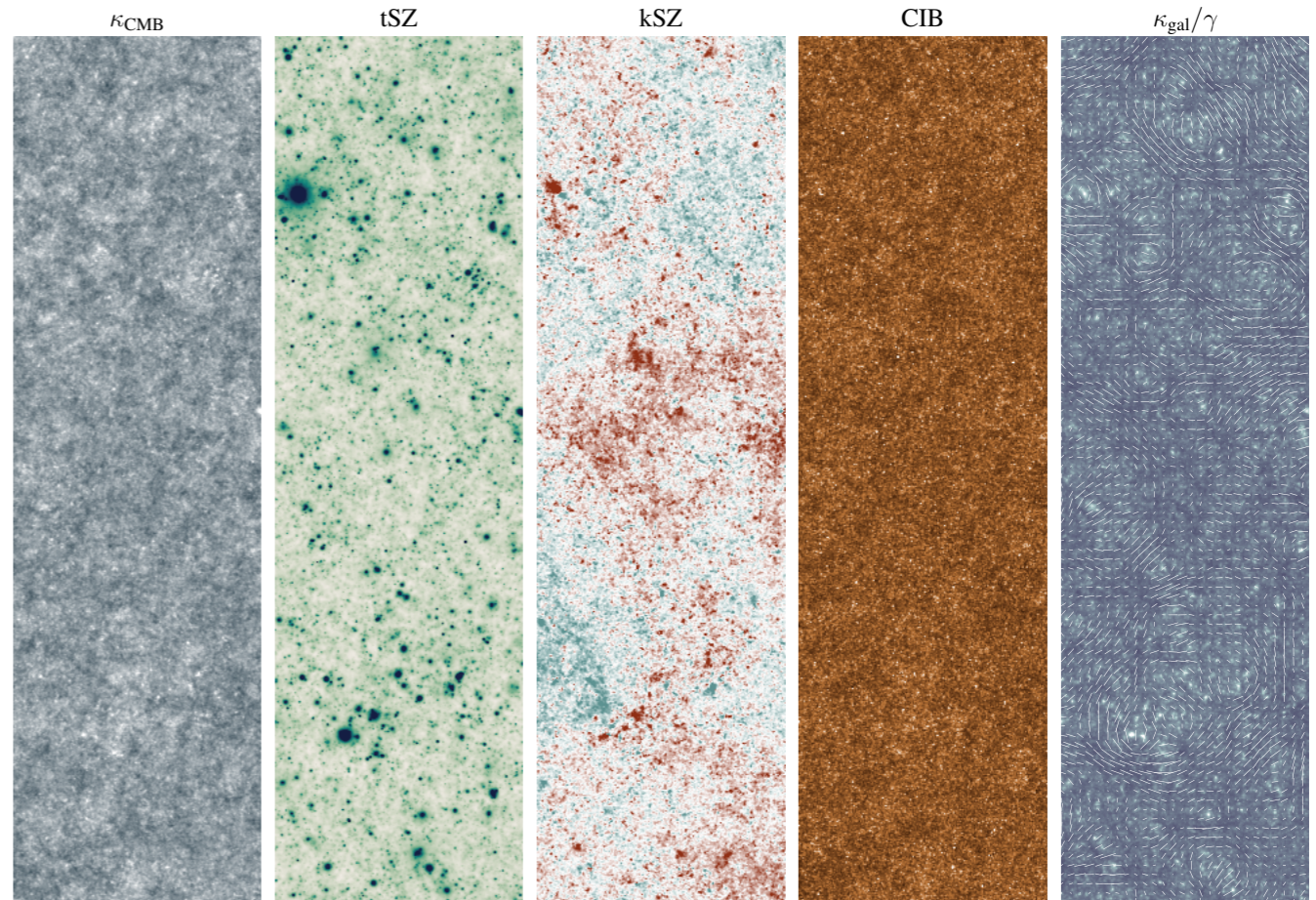
Correlated sims need to accommodate **BOTH** worlds

# Public CMB x LSS Simulations

- From LSS experts:
  - [Takahashi+2017](#) [ $\kappa^{\text{gal}}$ ,  $\kappa^{\text{CMB}}$ ]: 108 maps x full-sky, WMAP9
  - [BAHAMAS](#) (McCarthy+2018) [ $\kappa^{\text{gal}}$ ,  $\kappa^{\text{CMB}}$ , tSZ]: 25 maps x 25 deg<sup>2</sup>, Planck15, WMAP9
  - [MassiveNuS](#) (Liu+2018) [ $\kappa^{\text{gal}}$ ,  $\kappa^{\text{CMB}}$ ]: 10,000 maps x 12.25 deg<sup>2</sup>, 100 cosmologies ( $M_\nu$ ,  $A_s$ ,  $\Omega_m$ )
- From CMB experts:
  - [MillimeterDL](#) (Han+2021) [ $\kappa^{\text{gal}}$ ,  $\kappa^{\text{CMB}}$ , tSZ, kSZ, CIB, Radio, lensed CMB]: 500 maps x full-sky, WMAP5
  - [MDPL2synsky](#) (Omori in prep) [ $\delta$ ,  $\kappa^{\text{gal}}$ ,  $\kappa^{\text{CMB}}$ , tSZ, kSZ, CIB, Radio]: 1 map x full-sky, Planck15

# MULTI DARK PLANCK 2 SYNTHETIC SKY SIMULATION

- Based on the MultiDark Planck 2  $N$ -body simulation ([Klypin 2016](#),  $L = 1 h^{-1}\text{Gpc}$  box with  $3840^3$  particles), and associated halo catalogs.
- Galaxy evolution model (used for CIB/radio galaxies) from UniverseMachine ([Behroozi 2018](#)) and TRINITY ([Zhang 2021](#)).
- 1 full-sky realization (all the components provided as  $N_{\text{side}} = 8192$  maps or catalogs).

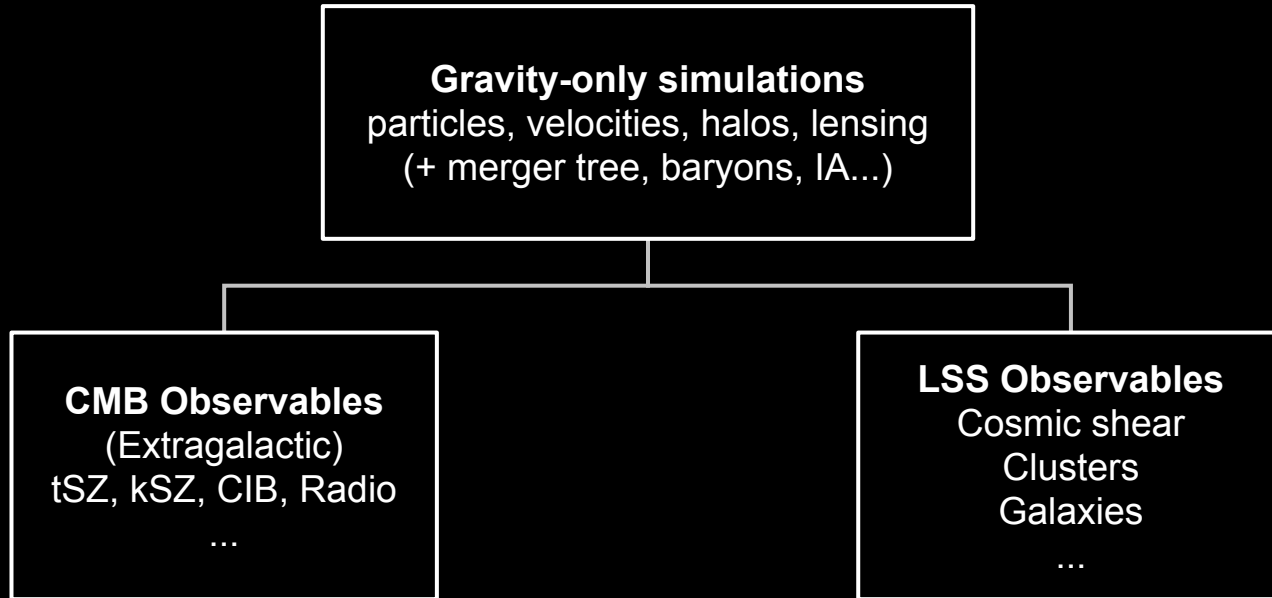


- Ray-tracing computed on  $N_{\text{side}} = 16384$ . All the CMB secondaries are lensed using the lensing field.
- Accuracy of the simulation tested for various types of analyses (high- $\ell$   $TT$ ,  $6\times 2\text{pt}$ , tSZ cross-correlation, component separation, delensing etc).
- Plans to implement various galaxy types via HOD.

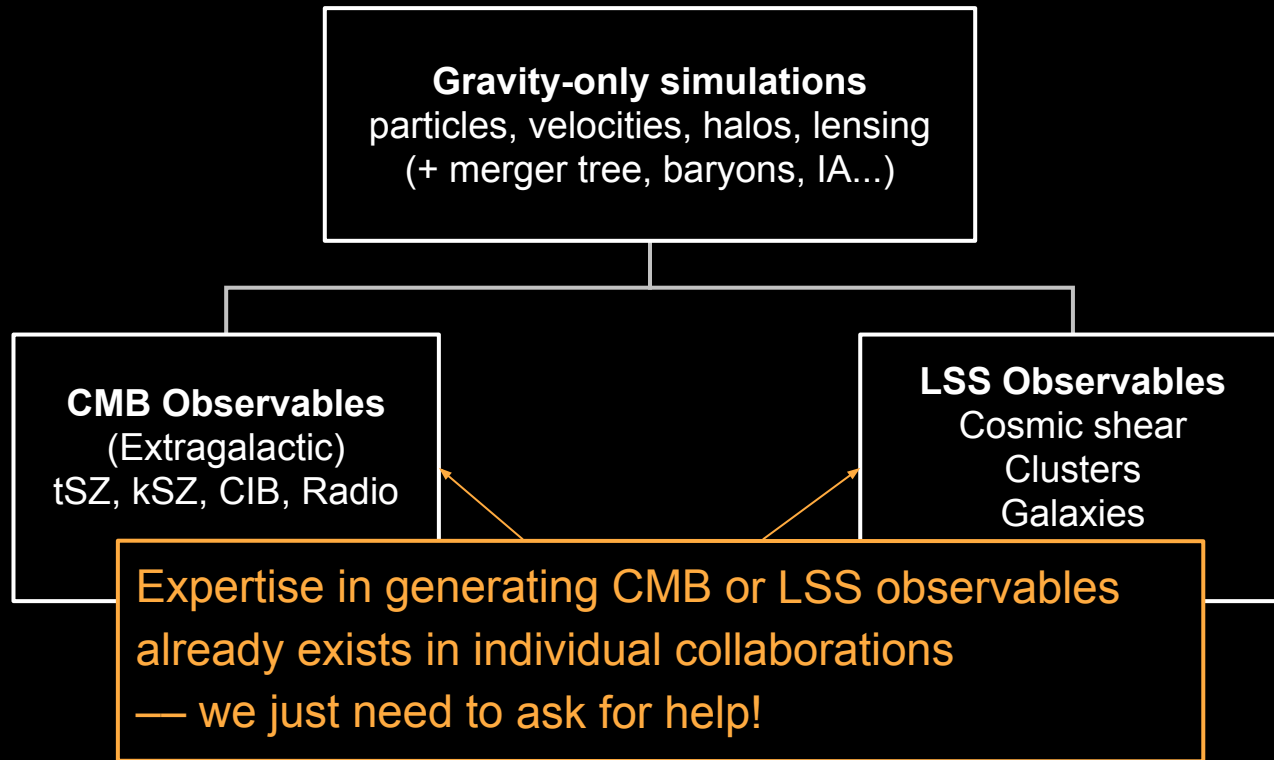
Observable	Redshift range
CMB lensing	$0 < z < 1089$
tSZ	$0 < z < 3.0$
kSZ	$0 < z < 3.0$
radio	$0 < z < 4.0$
CIB	$0 < z < 8.6$
$\delta$ (lensed)	$0 < z < 8.6$
$\kappa_{\text{g}}, \gamma_1, \gamma_2$	$0 < z < 8.6$



# Roadmap to Stage 4 Correlated Simulations (**my 2 cents**)



# Roadmap to Stage 4 Correlated Simulations (my 2 cents)



We can either (1) tile small/high-res boxes or (2) run a full-sky light cone with fast gravity-only codes (e.g. FastPM, COLA)

**Gravity-only simulations**  
particles, velocities, halos, lensing  
(+ merger tree, baryons, IA...)

**CMB Observables**  
(Extragalactic)  
tSZ, kSZ, CIB, Radio

**LSS Observables**  
Cosmic shear  
Clusters  
Galaxies

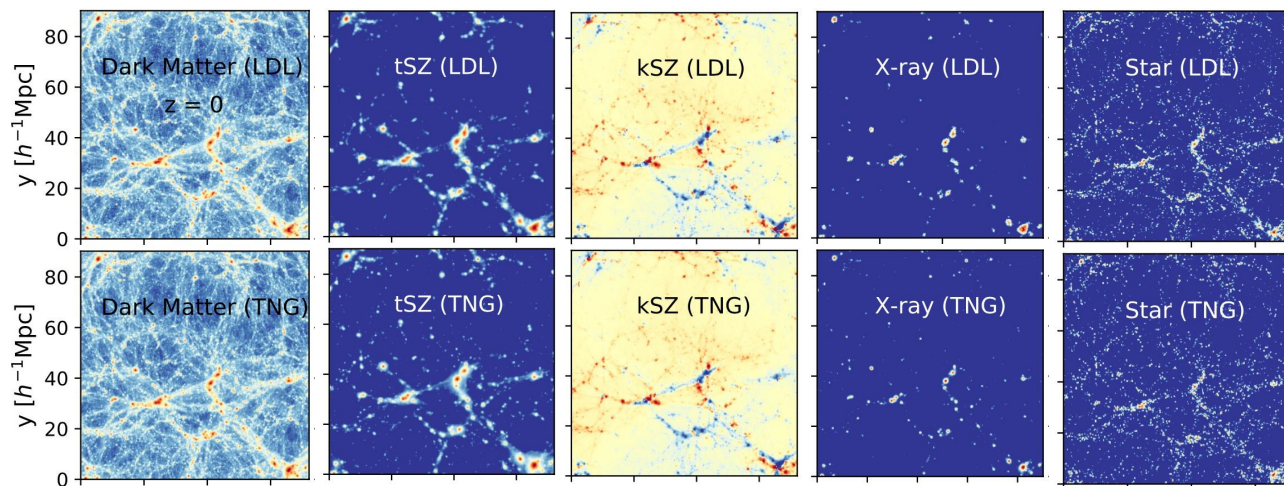
Expertise in generating CMB or LSS observables  
already exists in individual collaborations  
— we just need to ask for help!

# *FastPM* for the Gravity-only Simulations

**Code:** A particle-mesh gravity solver, with Potential Gradient Descent to improve small-scale clustering

**Performance:**  $10^7$  faster than hydro simulation;  $10^4$  faster than tree-PM N-body simulation; x10 slower than Websky (2LPT+peakpatch); used by the DESI collaboration

**Opportunities:** Lagrangian Deep Learning (astrophysics+painting); MADlens (differentiable lensing); flowPM (tensorflow-based/differentiability)



FastPM: Feng+2016  
PGD: Dai+2018  
LDL (*left img*): Dai & Seljak 2020  
MADlens: Böhm+2020  
FlowPM: Modi+2020

# Challenges

- **Computing time** (for 10-100 correlated simulations)
  - A few x 10 million CPU-hours (need a large fraction of a supercomputer at once)
- **Storage (~PB)**
  - ~50TB per model (can be smaller if only keeping maps and catalogues)
- **Joint requirements and validation**
  - With limited resources, what are the key requirements?
  - Validation can take a long time
- **Maintenance**
  - Documentation
  - Data release
  - Future upgrades
- **Personnel**
  - Training and acknowledging the simulation scientists

# Interested & want to get involved?

- PhD projects
- Postdoc positions (to be announced in the fall)
- or simply share thoughts!

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[jjaliu@berkeley.edu](mailto:jjaliu@berkeley.edu) (email address expires on 8/31/2021)

<https://liu479.github.io/> (for updated contact later)