

# CMB and BBN constraints on light thermally coupled WIMPs

*Rui An*

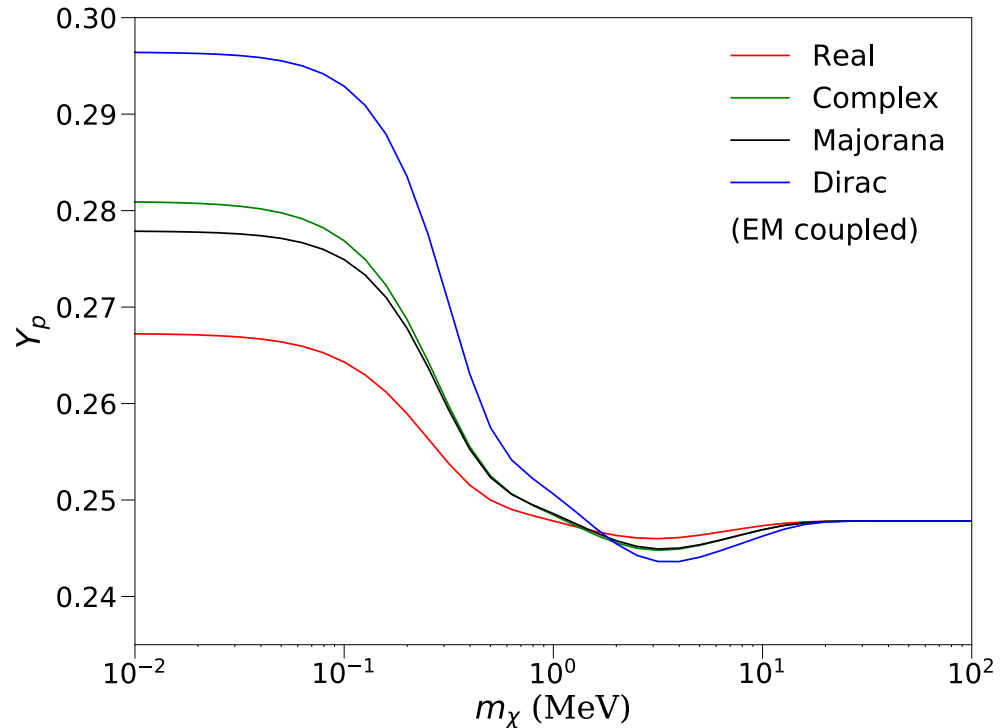
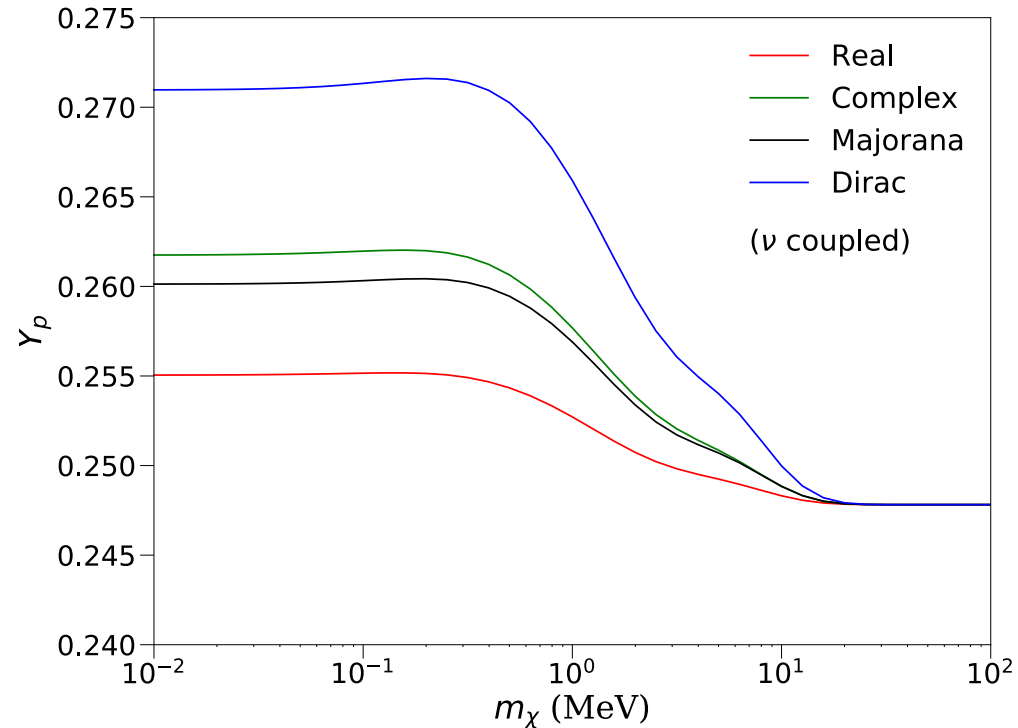
*University of Southern California*

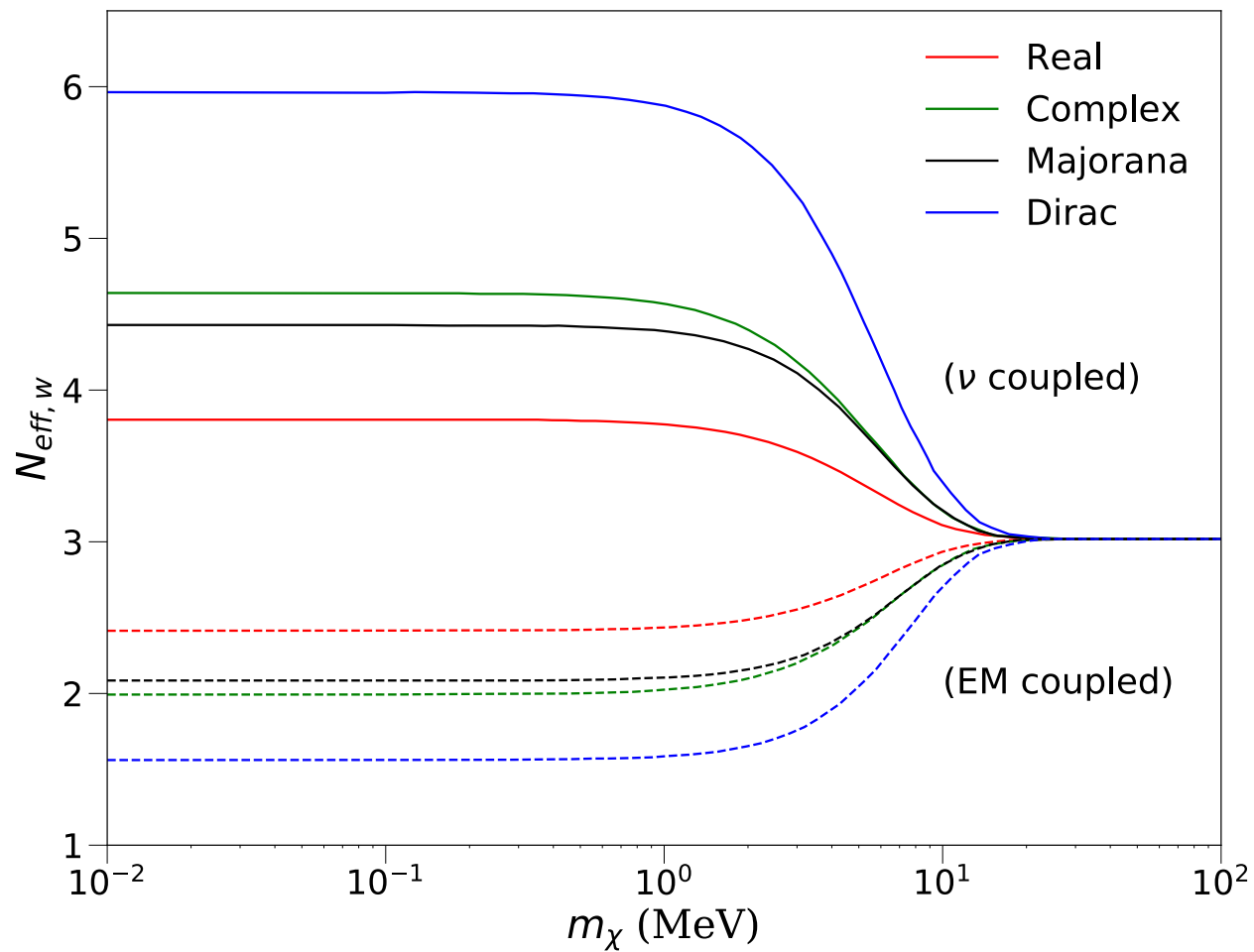
# BBN in WIMP scenarios

WIMP mass: 0.01MeV~100MeV

Types of particle: Real Scalar, Complex Scalar, Majorana Fermion, Dirac Fermino

Types of coupling: WIMPs annihilating to neutrinos ( $\nu$  coupled) or photons and  $e^\pm$  pairs (EM coupled)



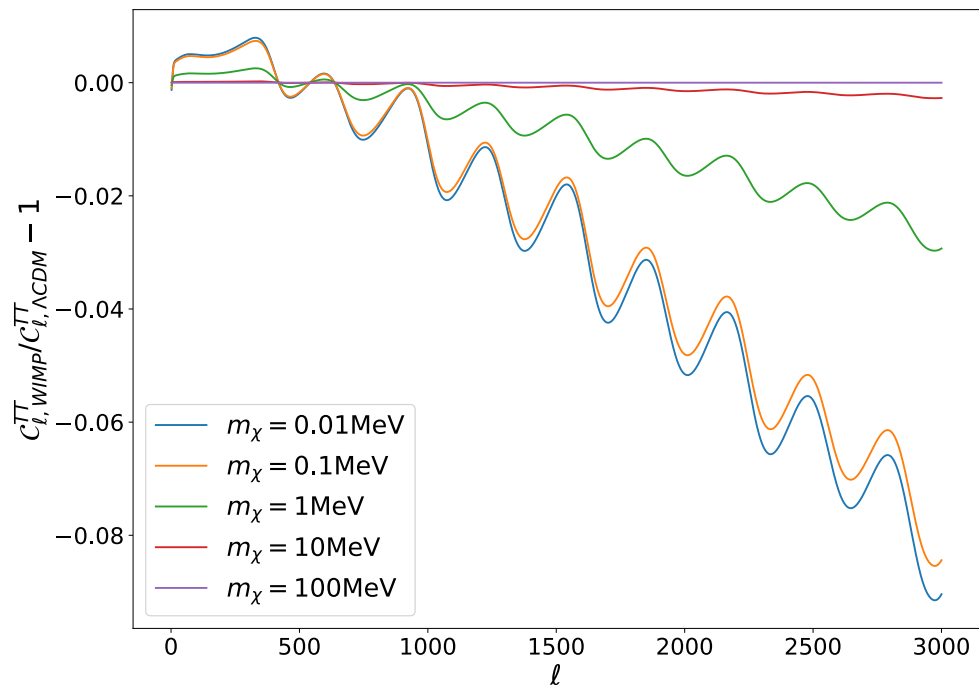


$$N_{eff} = N_{eff,w} \left( 1 + \frac{\Delta N_\nu}{3} \right)$$

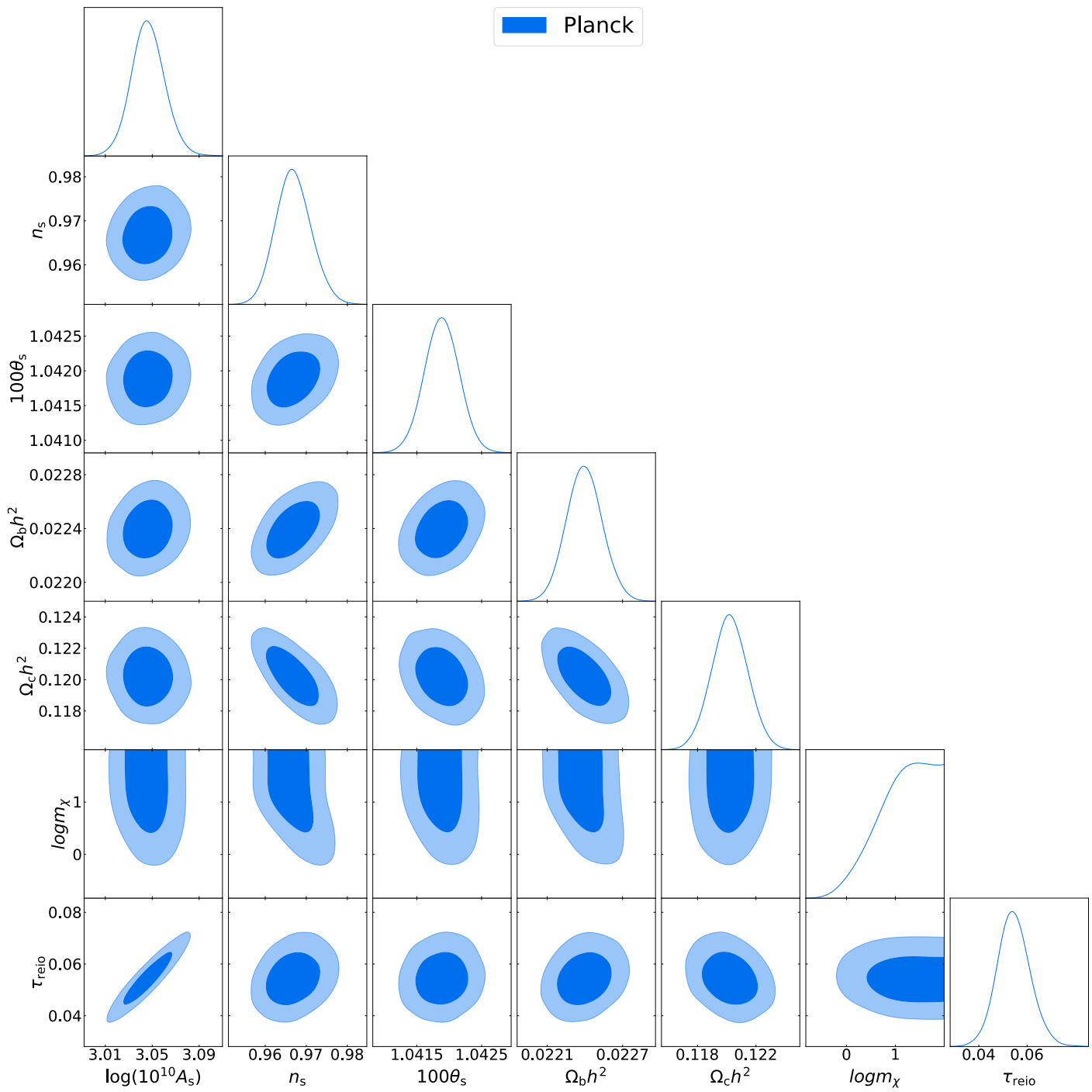
$$N_{eff,w}: 3 \left[ \frac{11}{4} \left( \frac{T_\nu}{T_\gamma} \right)_0^3 \right]^{4/3}, \text{ altered by WIMPs}$$

$\Delta N_\nu$ : the number of equivalent neutrinos beyond the standard model

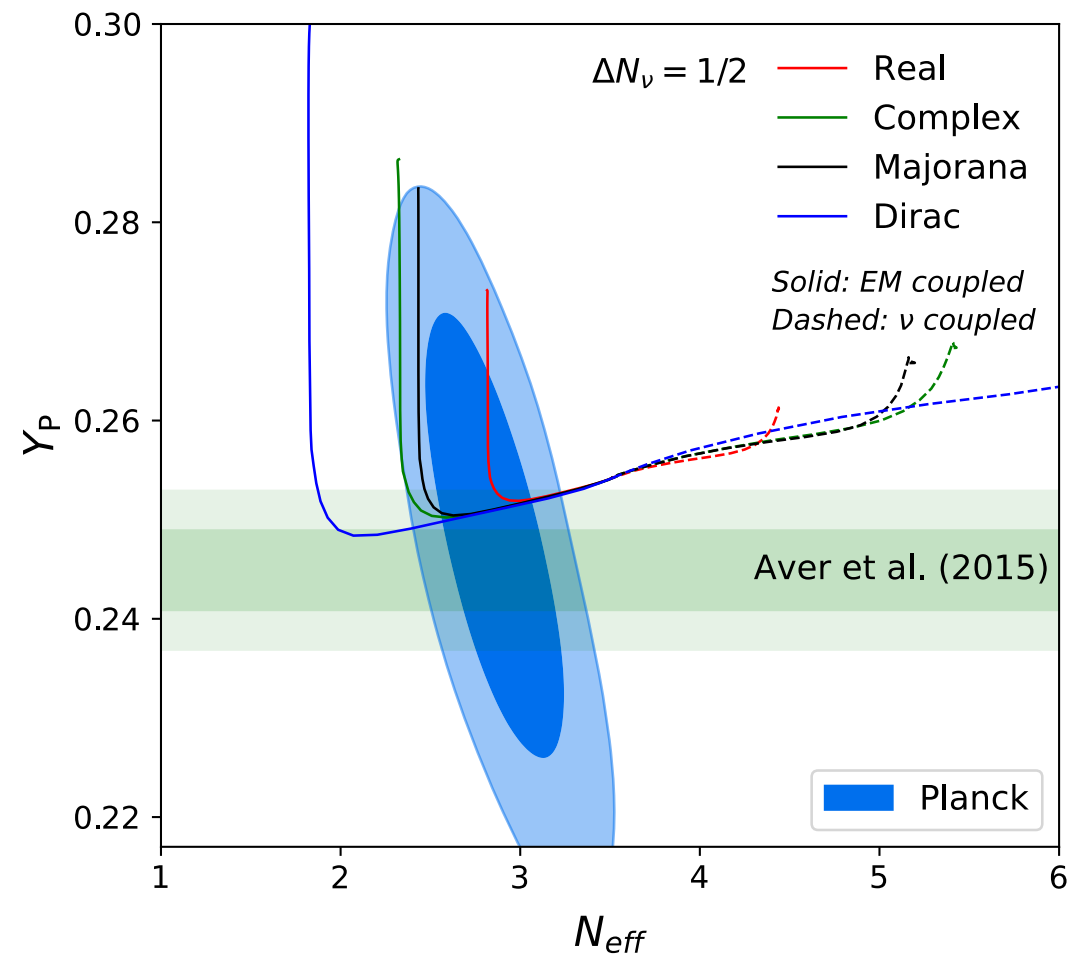
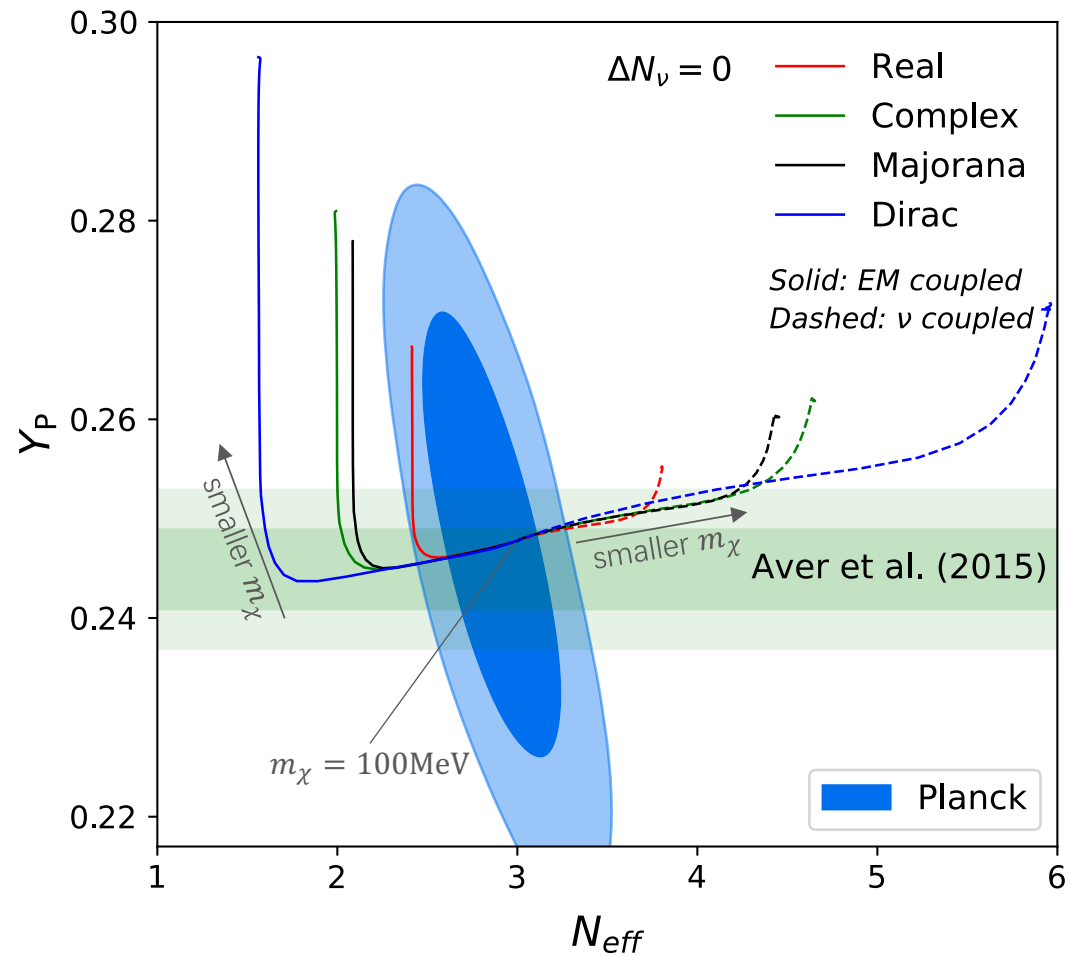
# Effects of WIMPs on CMB



(EM coupled Majorana WIMP)



# CMB and BBN constraints



Thanks!